# Atmospheric Neutrino Oscillations in IceCube DeepCore

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AT School 2024 Obertrubach-Bärnfels

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- Ice Cherenkov v detector
- 1.5 2.5 km under ice
- 5,160 DOMs on 86 strings
- 1 km<sup>3</sup> volume
- High energy array spacing
  - Δ*z*=17m
  - $\Delta(x, y) = 125m$
- LE extension: DeepCore
  - Δ*z*=7m
  - $\Delta(x, y)$ =40-70m



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#### The DeepCore data

- Use all/most events starting in the DeepCore region
  - Strong atm.  $\mu$  background suppression still, some will remain
  - Mostly contained energy estimator is reasonable
  - Information from interaction available: topology  $\rightarrow$  flavor



Color indicates time (red=early, blue=late). Sphere size is proportional to number of photons observed.

## Measurements of neutrino oscillations (DeepCore)



# Analysis strategy for oscillations



# Analysis strategy for oscillations



#### What are we doing today?

Explore the data release for IceCube's 2018 result - PRL 120, 071801 (2018)

- Load the data, look at the "observables" that IceCube uses
- Understand what signature is expected from oscillations
- Learn how you compute an "oscillated flux of neutrinos"
- Test the impact of oscillations on atmospheric neutrinos
- Extract the oscillation parameters using the simplified 2-flavor formula

$$P_{\nu_{\mu} \to \nu_{\tau}} \simeq \sin^2 2\theta_{23} \, \sin^2 \left( \frac{\Delta m_{32}^2}{4E} L \right) = A \, \sin^2 \left( 1.267 \, \frac{\Delta m_{32}^2}{eV^2} \, \frac{L/km}{E/GeV} \right)$$

### Links and tools

- Coding in Google Colaboratory
  - Link to the folder with code and data
  - Open the code and copy to your own google drive
  - Download dragon\_data.pkl and upload to your own google drive
- Follow the code, think about the questions, propose answers
- Fill in the gaps (write the missing code)
- Run the fit