

# Broadband Lightning Radio Interferometry at the Pierre Auger Observatory

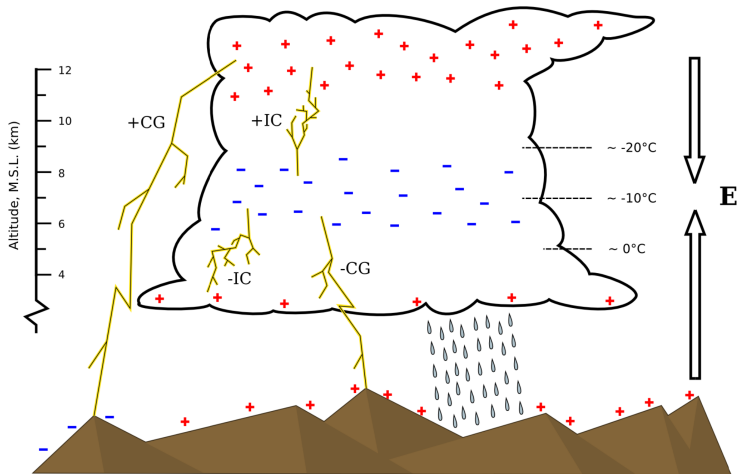
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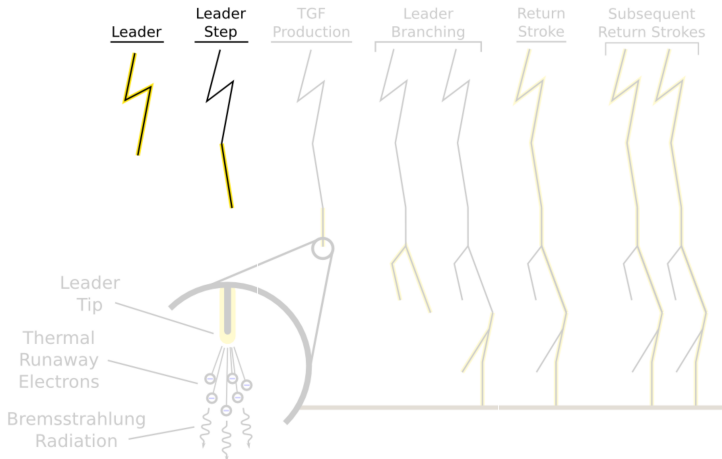
19th Erlangen School for Astroparticle Physics,  
Obertrubach-Bärnfels,  
October 08, 2024

# Thunderstorms and Lightning



Source: J. Remington's PhD thesis [1]

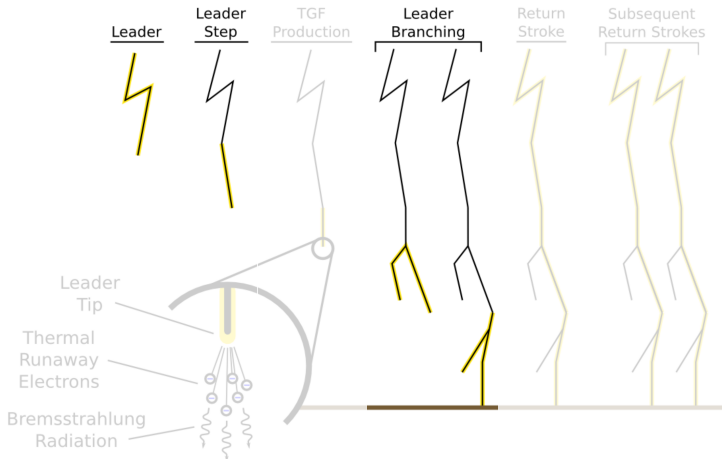
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- Many radio pulses from Leaders in VHF (MHz) regime,

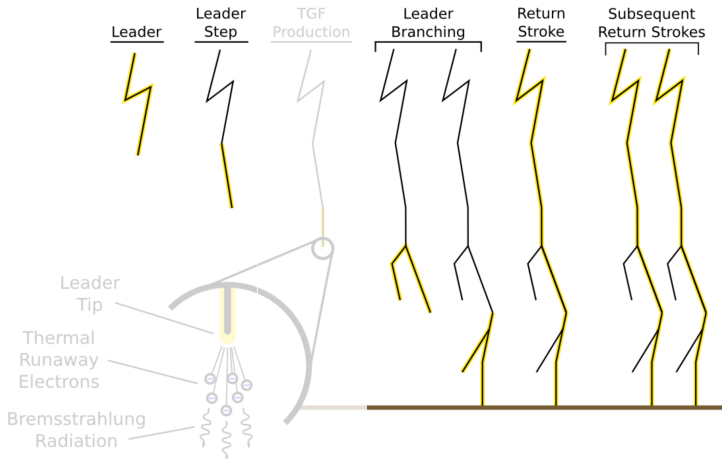
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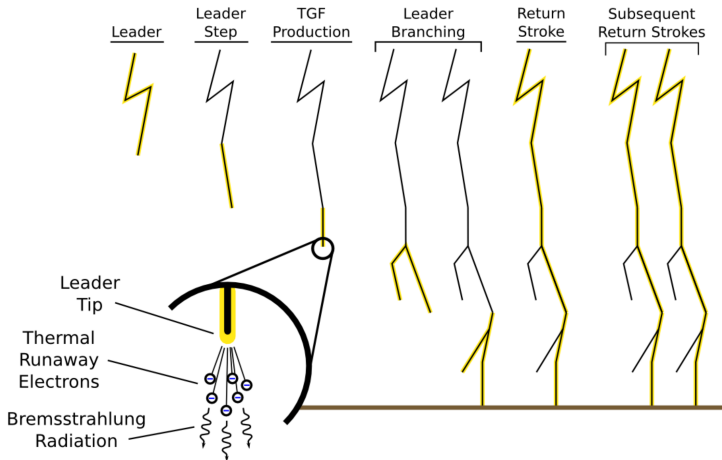
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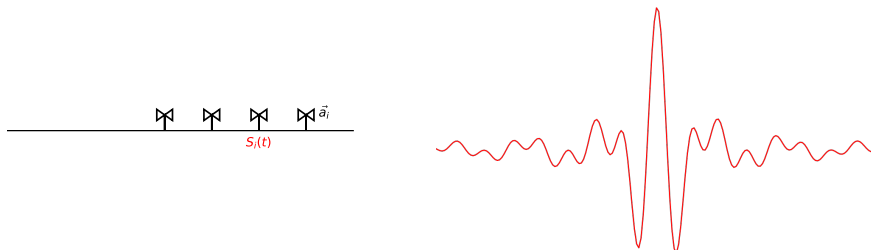
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# Radio Interferometry

Interferometry: Amplitude + Timing information of the signal

- Measure signal  $S_i(t)$  at antenna  $\vec{a}_i$



# Radio Interferometry

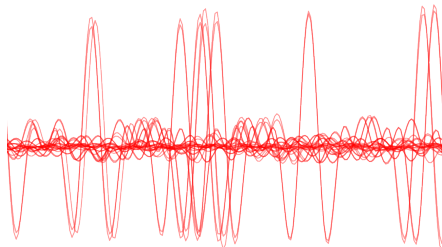
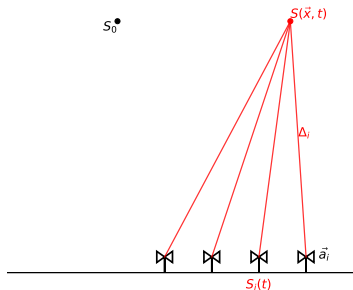
Interferometry: Amplitude + Timing information of the signal

- Measure signal  $S_i(t)$  at antenna  $\vec{a}_i$
- Calculate light travel time

$$\Delta_i(\vec{x}) = \frac{|\vec{x} - \vec{a}_i|}{c} n_{eff}$$

- Sum/Multiply waveforms accounting for time delay

$$S(\vec{x}, t) = \sum S_i(t + \Delta_i(\vec{x}))$$





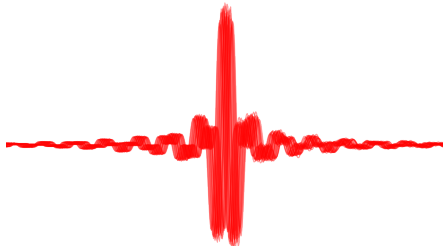
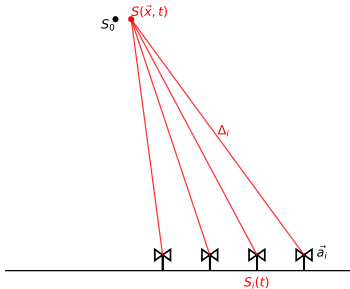
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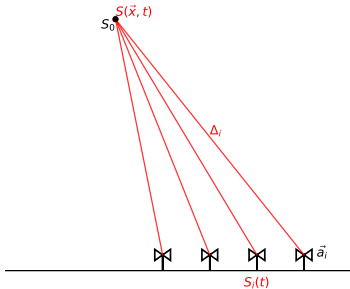
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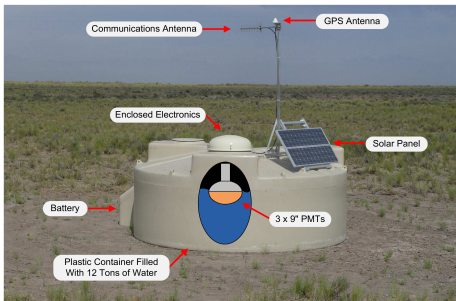
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# The Pierre Auger Observatory

Hybrid cosmic ray experiment located in the Pampa Amarilla in Argentina

## Surface Detectors (x1660)

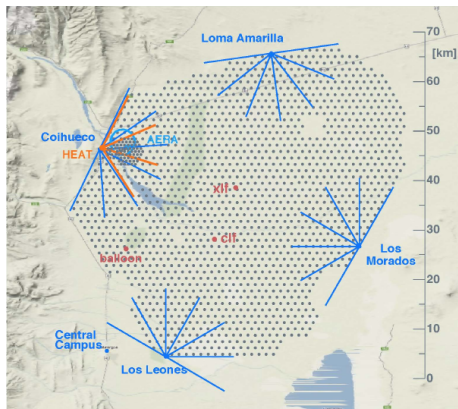


Source: L. Niemietz's PhD thesis [2]

## Fluorescence Detectors (x4)



# AERA: Auger Engineering Radio Array



Source: L. Niemietz's PhD thesis [2]

Sensitivity: 30 MHz to 80 MHz  
Sampling Rate: 180 MHz  
Tracelength:  $2048 \times 5.5 \text{ ns} \sim 11 \mu\text{s}$

## Phase I: LPDA (x24)



## Phase II/III: Butterfly (x130)



# Radio detection for Lightning versus Air Showers

Rule of Thumb for Radio Interferometry:

'Timing accuracy must be better than  $1/4$  ( $1/12$ ) of the period.'

.. but we are doing (digital) broadband interferometry.

## Air Showers

- relativistic emitter
- single pulse  
( $\lesssim 10$  ns)
- localised footprint  
(projected around shower axis)

## Lightning

- extended emitter
- multiple pulses /  
substructure  
(10 ns to  $10 \mu\text{s}$ ,  $\Delta \sim 10$  ms)
- large footprint

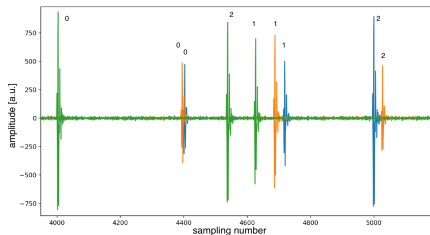
# Lightning Interferometry: reusing AERA

## Requirements:

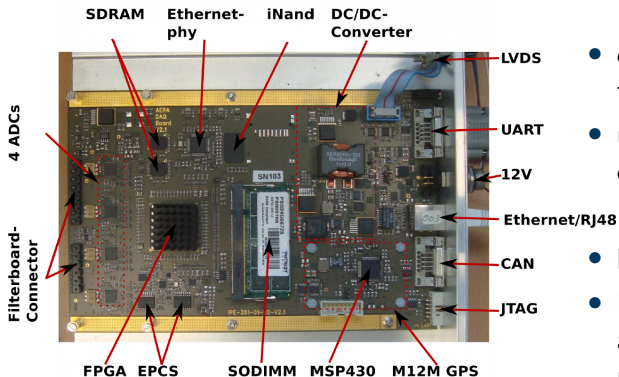
- modifications for AERA stations:
  - extending trace lengths from 11  $\mu\text{s}$  to the order of 1 s,
  - transmitting the data from the stations.
- trigger development
- nanosecond synchronisation across stations with distances up to several kilometers

## Planned configuration:

- 4 Core stations, baselines 58 m to 127 m (for pulse identification/separation)
- 3 Medium range, baselines 1 km to 2.5 km
- 4 Remote stations, baselines 3.5 km to 66 km



# AERA Digitizer



- custom FPGA gateway for continuous sampling
- uClinux for high-level coordination
- last development 2017
- FPGA Cyclone III, and NIOS II CPU, unsupported...

Source: B. Zimmermann's PhD thesis [3]

# Summary and Outlook

- Geophysical sources of high-energy particles correlated with lightning.
- Idea: hybrid lightning measurements at the Pierre Auger Observatory, using SD, FD, E-field mills and repurposed AERA stations.
- AERA requires modification for lightning interferometry.

## Outlook:

First repurposed AERA station to feature extended trace lengths by  
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*Thank you!*

# Bibliography I



Jackson Remington. *Analysis of Downward Terrestrial Gamma-ray Flashes Using a Large-Area Cosmic Ray Scintillation Detector*. PhD thesis, University of Utah, December 2021. arXiv: 2308.05662 [astro-ph.HE].



Lukas Niemietz. *Lightning Detection at the Pierre-Auger-Observatory*. PhD thesis, Bergische Universität Wuppertal, June 2017. URL: <https://nbn-resolving.org/urn:nbn:de:hbz:468-20171012-105730-3>.



Benedikt Hermann Zimmermann. *Radio-detection of ultra-high energy cosmic rays: multi-channel data acquisition and combined analysis of timing and energy information*. PhD thesis, Karlsruher Institut für Technologie (KIT), 2017. 105 pages. DOI: 10.5445/IR/1000076009. 51.03.03; LK 01.