

### VLBI Probes of Jet Physics in Neutrino-Candidate Blazars

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Radio Images of PKS 1424-418 before and after a possibly coincident ( $2\sigma$ ) Neutrino Event (Kadler et al. 2016)















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→ Need for more statistics and individual case-studies to understand the physics of neutrino production





• TeV Effelsberg Long-term Agn MONitoring





Effelsberg 100-m Telescope

Kadler et al. (2021, ICRC PoS), Eppel et al. (2023, in prep.)



**VLBI** Probes of Jet Physics in Neutrino Candidate Blazars



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- TeV Effelsberg Long-term Agn MONitoring
- Est. Aug 2020
- Monitoring radio spectra of TeV-Blazars and candidate neutrino-assoc. AGN
- Sample of ~60 sources, mostly HBLs
- High cadence monitoring observations
  - Every 2-4 weeks
  - From 5 GHz up to 45 GHz (45mm, 20mm, 14mm, 7mm)
- Neutrino follow-up observations





Effelsberg 100-m Telescope

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Main Goals (long term):

• correlation (especially with upcoming CTA)

-detections.





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Publicly available data distributed via website: <u>https://telamon.astro.uni-wuerzburg.de</u>







PKS 0215+015

# **Good Things Come In Threes**

#### TXS 0506+056



#### PKS 1502+106



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PKS 0215+015

# Good Things Come In Threes

#### TXS 0506+056





 $\rightarrow$  All radio-flaring at time of neutrino event and followed up with VLBI ToO observations



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Multiwavelength-Flare associated with Neutrino Event IC170922A



43 GHz VLBA images and surface brightness profiles of TXS 0506+056 after the associated neutrino event providing evidence of limb-brightening (Ros et al. 2020)



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Multiwavelength-Flare associated with Neutrino Event IC170922A

 VLBA DDT/ToO Campaign (Oct 2017, Nov 2017 & May 2018)



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Multiwavelength-Flare associated with Neutrino Event IC170922A

- VLBA DDT/ToO Campaign (Oct 2017, Nov 2017 & May 2018)
- 32 MOJAVE observations at 15 GHz available in the archive



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Multiwavelength-Flare associated with Neutrino Event IC170922A

- VLBA DDT/ToO Campaign (Oct 2017, Nov 2017 & May 2018)
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- GMVA observations at 43 GHz & 86 GHz in Oct 2020 and Apr 2021



43 GHz VLBA images and surface brightness profiles of TXS 0506+056 after the associated neutrino event providing evidence of limb-brightening (Ros et al. 2020)



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- Spectral-Index Analysis
- Kinematic Analysis

Possible new jet component found





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# TXS 0506+056 (ongoing work)



#### **Main Goals**

- Investigate super-luminal coreexpansion found by Ros et al. 2020 with more VLBI epochs and at higher frequencies
- Spectral-Index Analysis
- Kinematic Analysis
- Core-Shift Analysis

Possible new jet component found




# TXS 0506+056 (ongoing work)



#### **Main Goals**

- Investigate super-luminal coreexpansion found by Ros et al. 2020 with more VLBI epochs and at higher frequencies
- Spectral-Index Analysis
- Kinematic Analysis
- Core-Shift Analysis
- Investigate magnetic field and spinesheath structure with polarization information

Possible new jet component found





Kun et al. 2021





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Kun et al. 2021

Radio-Flare coincident with neutrino event IC190730A but NO gamma-ray flare







Kun et al. 2021

Radio-Flare coincident with neutrino event IC190730A but NO gamma-ray flare

 VLBA DDT/ToO Campaign (Aug 2019)







Kun et al. 2021

Radio-Flare coincident with neutrino event IC190730A but NO gamma-ray flare

- VLBA DDT/ToO Campaign (Aug 2019)
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Kun et al. 2021

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43 GHz Image of PKS 1502+106 ~1 month after the Neutrino Event

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Main Goals



43 GHz Image of PKS 1502+106 ~1 month after the Neutrino Event

Main Goals

• Study kinematics and evolution of EVPA after the neutrino event



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43 GHz Image of PKS 1502+106 ~1 month after the Neutrino Event

#### Main Goals

- Study kinematics and evolution of EVPA after the neutrino event
- Investigate possible EVPA swings which have been observed before



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43 GHz Image of PKS 1502+106 ~1 month after the Neutrino Event

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- Spectral-Index Analysis



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43 GHz Image of PKS 1502+106 ~1 month after the Neutrino Event

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43 GHz Image of PKS 1502+106 ~1 month after the Neutrino Event

### Main Goals

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- Investigate possible EVPA swings which have been observed before
- Spectral-Index Analysis
- Core-Shift Analysis
- Investigate spine-sheath structure in combination with lepto-hadronic SED models





Eppel et al. 2023

### PKS 0215+015: A New Neutrino Emitter?



Gamma-ray and radio flaring of **PKS 0215+015** in coincidence with the IceCube Neutrino Alert IC220225A



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Eppel et al. 2023



Gamma-ray and radio flaring of **PKS 0215+015** in coincidence with the IceCube Neutrino Alert IC220225A



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Gamma-ray and radio flaring of PKS 0215+015 in coincidence with the IceCube Neutrino Alert IC220225A







Eppel et al. 2023



Gamma-ray and radio flaring of  $\mbox{PKS}$  0215+015 in coincidence with the IceCube Neutrino Alert IC220225A

VLBA ToO-Proposal

- Total of 6 epochs after the neutrino event (1 per month)
- Three-point spectral-index (20mm, 14mm, 7mm) and Faraday-rotation measurements



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Eppel et al. 2023



Gamma-ray and radio flaring of **PKS 0215+015** in coincidence with the IceCube Neutrino Alert IC220225A

VLBA ToO-Proposal

- Total of 6 epochs after the neutrino event (1 per month)
- Three-point spectral-index (20mm, 14mm, 7mm) and Faraday-rotation measurements
- Probe jet kinematics



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Eppel et al. 2023



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- Three-point spectral-index (20mm, 14mm, 7mm) and Faraday-rotation measurements
- Probe jet kinematics
- Search for characteristic similarities to TXS 0506+056 and other candidate neutrino-emitting







Eppel et al. 2023



Gamma-ray and radio flaring of PKS 0215+015 in coincidence with the IceCube Neutrino Alert IC220225A

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- Three-point spectral-index (20mm, 14mm, 7mm) and Faraday-rotation measurements
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 $\rightarrow$  First VLBA observation took place on **March 24**, **2022** (total of 6 epochs)











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### **VLBA observations:**

- 6 epochs with 20 mm, 14 mm and 7 mm
- 3 additional MOJAVE epochs with only 20mm available







#### VLBA observations:

- 6 epochs with 20 mm, 14 mm and 7 mm
- 3 additional MOJAVE epochs with only 20mm available

### **Current Status:**

- All epochs calibrated in AIPS & rPICARD/CASA
- Preliminary imaging and selfcal performed in DIFMAP







### VLBA observations:

- 6 epochs with 20 mm, 14 mm and 7 mm
- 3 additional MOJAVE epochs with only 20mm available

### **Current Status:**

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- Preliminary imaging and selfcal performed in DIFMAP

### To Do:

- Consistency check of final images
- Modelfitting
- Polarization calibration/imaging









### PKS 0215+015: First Look at 15 GHz



#### 2022-01-03 (MOJAVE) - pre-neutrino



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### PKS 0215+015: First Look at 15 GHz



2022-01-03 (MOJAVE) - pre-neutrino

2022-03-24 - post-neutrino



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#### PRELIMINARY



#### Ku-Band (15 GHz)



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**VLBI** Probes of Jet Physics in Neutrino Candidate Blazars

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VLBI Probes of Jet Physics in Neutrino Candidate Blazars

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Linear Polarization Fraction of **PKS 0215+015** after the associated Neutrino Event



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Linear Polarization Fraction of **PKS 0215+015** after the associated Neutrino Event • TELAMON data shows elevated polarization state close to neutrino event









Linear Polarization Fraction of **PKS 0215+015** after the associated Neutrino Event

- TELAMON data shows elevated polarization state close to neutrino event
- Preliminary VLBA polarization image produced in CASA with PolSolve (Marti-Vidal et al. 2020)











Linear Polarization Fraction of **PKS 0215+015** after the associated Neutrino Event

- TELAMON data shows elevated polarization state close to neutrino event
- Preliminary VLBA polarization image produced in CASA with PolSolve (Marti-Vidal et al. 2020)
- Polarized emission core dominated







# Summary & Outlook





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# Summary & Outlook

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- We have triggered VLBA ToO observations to follow up the pc-scale structure of the jet in total intensity and polarization






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- We have triggered VLBA ToO observations to follow up the pc-scale structure of the jet in total intensity and polarization
- Main Goal: Search for similarities/differences in order to understand the physics of neutrino production in blazars

> We now have a unique data set at hand to investigate neutrino emission mechanisms in blazars







# **Questions?**



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# **Ku-Images**









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# K-Images



#### PRELIMINARY





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Density [Jy]

Flux I



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2 -

Relative DEC. [mas] 1-

-2

3

2 .

Relative DEC. [mas]

0 -

-2

3

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