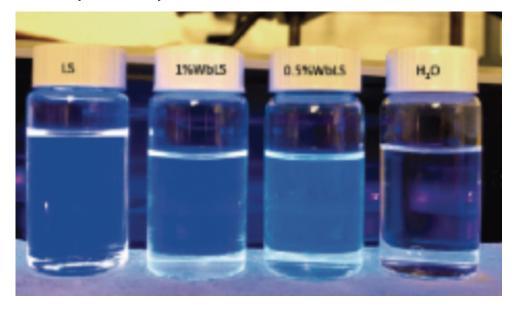


DISCO Experiment

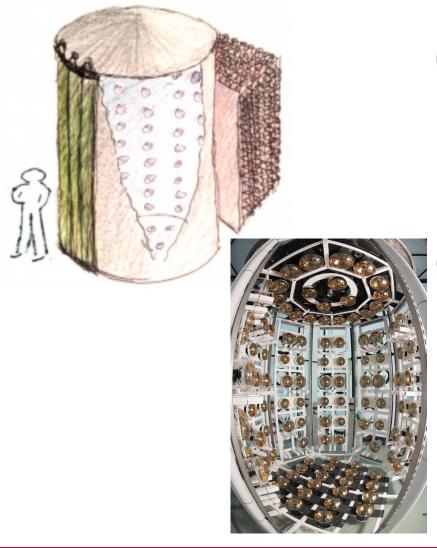
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Astroparticle School 2023

INTRODUCTION

- Directionality In Scintillation Observer (DISCO) detector
 - Demonstrate the sepeartion of Cherenkov and Scintillation light
 - Charecterize Water Based Liquild Scintillators (WbLS)
- WbLS → water + scintillator
- Tunable light yield and timing profile
- Properties
 - Low energy threshold
 - High light yield
 - Directional information
 - Increased attenuation length



WATER BASED LIQUID SCINTILLATORS (Wbls)

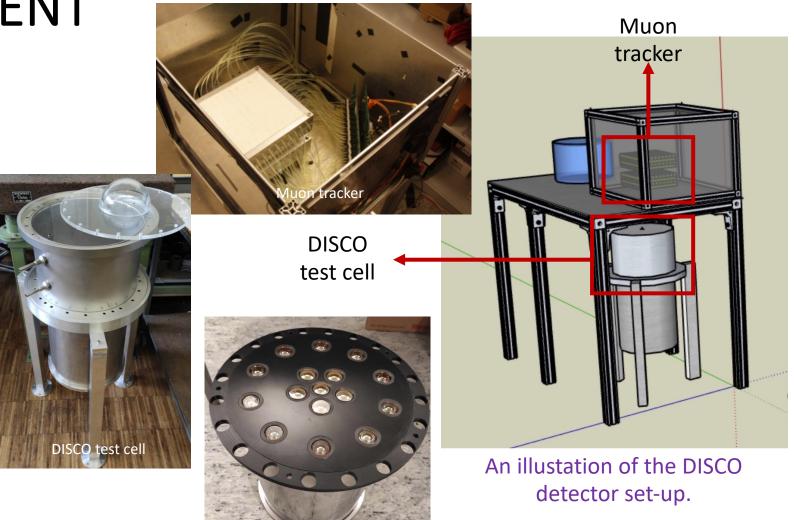




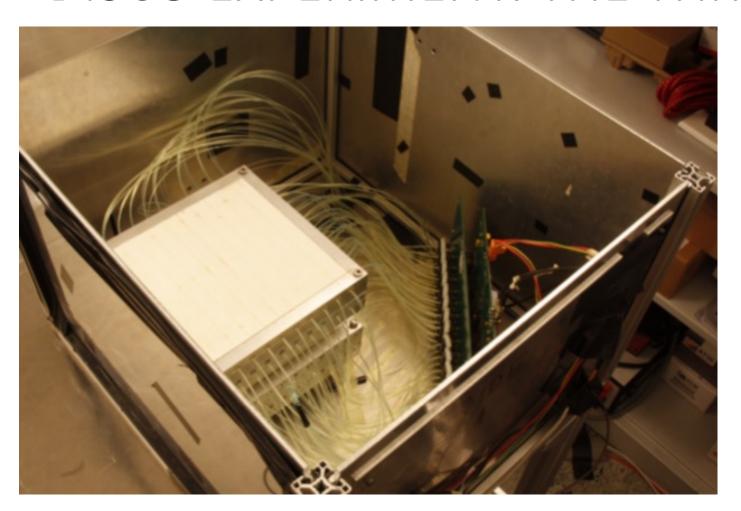
- Why use WbLS??
 - Enables detection of both Cerenkov and Scintillation light
 - Construct large detectors
 - Low energy threshold
 - Use in future experiments like THEIA, ANNIE, etc.

DISCO EXPERIMENT

- 3 main components:
 - Muon tracker
 - Test cell
 - Light detectionsystem
- Test particles muons



DISCO EXPERIMENT: THE THREE COMPONENTS



MUON TRACKER

- Eight planes, each plane consists of eight 19 cm long scintillating rods
- Read-out using SiPMs
- Provide external trigger signal
- Reconstruct muon tracks
- Placed above the test cell

DISCO EXPERIMENT: THE THREE COMPONENTS

TEST CELL

- Cylindrical steel tank
 enclosed with two
 plexiglass planes
- Detection volume ~ 14 L
- Height : 20 cm
- Diameter: 30 cm
- Can be filled with water,
 LS, WbLS
- Currently filled with water







DISCO EXPERIMENT: THE THREE COMPONENTS

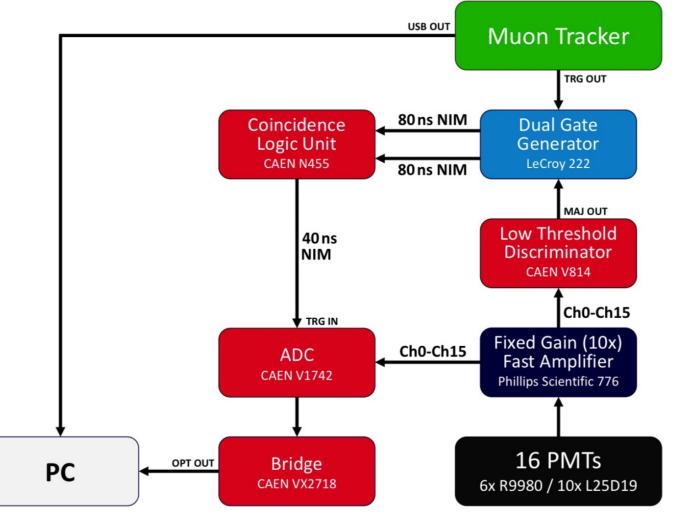


LIGHT DETECTION SYSTEM

- Cherenkov photons
 detected using sixteen
 1" PMTs placed below the
 test cell.
- PMTs arranged in two concentric circles of radii
 40 mm and 110 mm.
- PMTs are calibrated using laser of wavelength 405 nm

DATA ACQUISITION SYSTEM

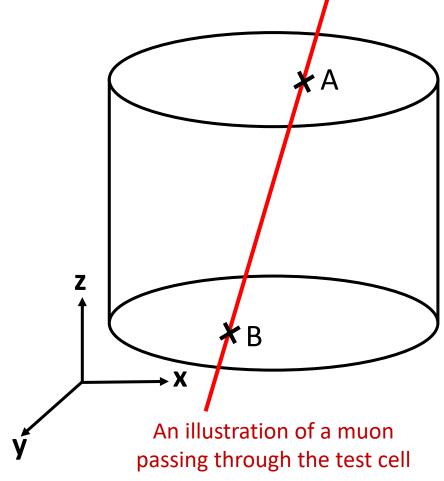
- PMT signal is amplified by a factor of 10
- External trigger:
 - Muon tracker + PMT signal
- PMT signals digitalized on receiving external trigger.
- Offline analysis → Extract hit time and charge



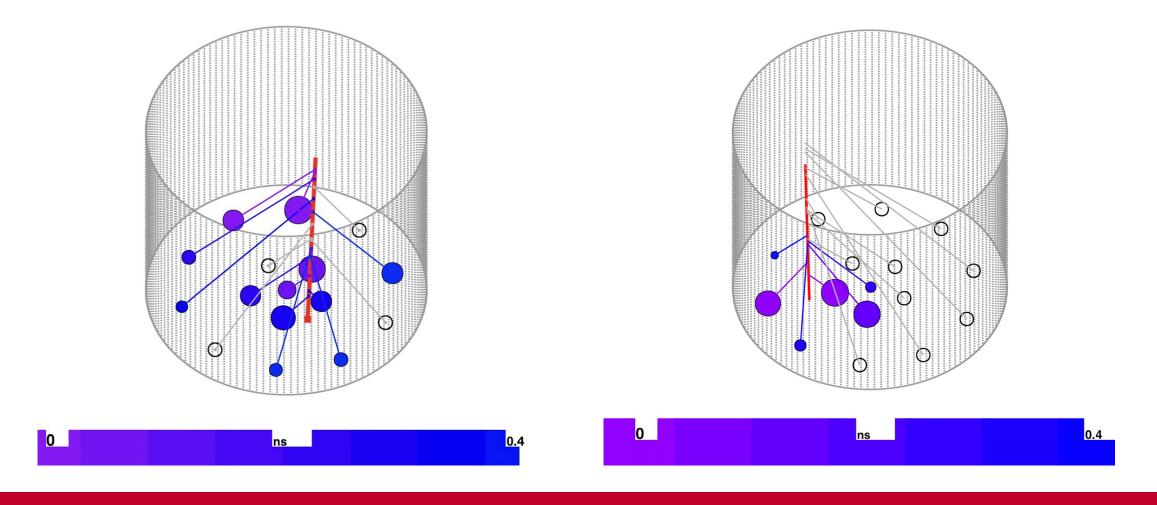
A schematic of the DAQ

TRACK RECONSTRUCTION

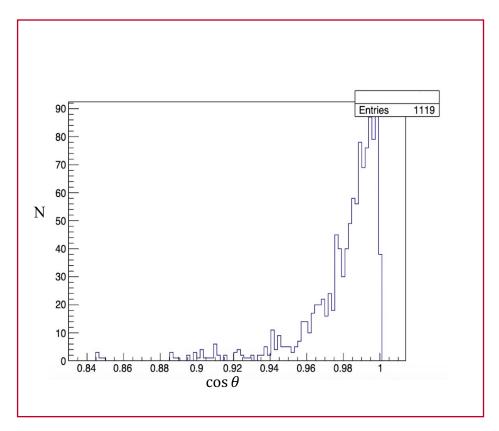
- Muon tracks described by:
 - \circ Entry point (A) (x_1, y_1, z_1)
 - o Exit point (B) (x_2, y_2, z_2)
- Track parameters are obtained by minimizing the log likelihood function.
- Qualitative assessment of reconstruction procedure: look at the angular distribution of reconstructed tracks.



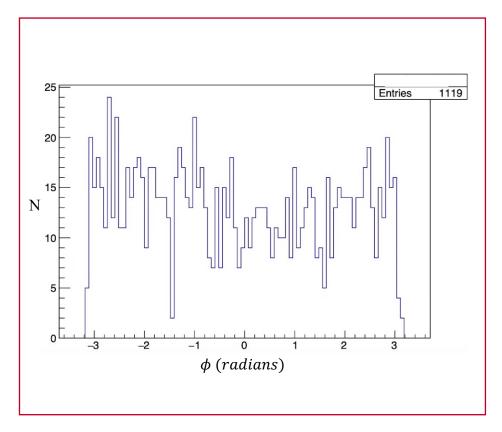
EVENT DISPLAYS



ANGULAR DISTRIBUTIONS OF RECONSTRUCTED TRACKS



Distribution of muons as a function of zenith angle



Distribution of muons as a function of azimuthal angle

SUMMARY AND NEXT STEPS

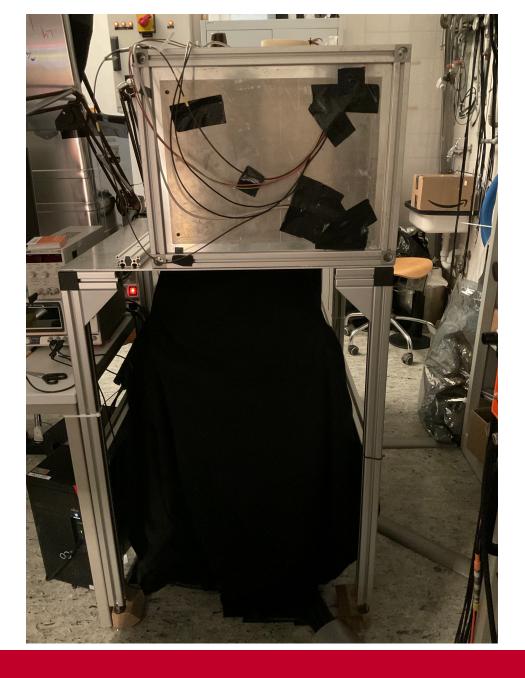
- DISCO: lab scale experiment to charecterize WbLS
- Tested DISCO with water

• Muon tracks were reconstructed by combining the PMT hit-time (χ_t^2)

and charge information (log likelihood)

- Next Steps:
 - Replace water with WbLS in test cell and take data
 - Replace PMTs at the bottom with LAPPD → enables precise track reconstruction

BACK-UP



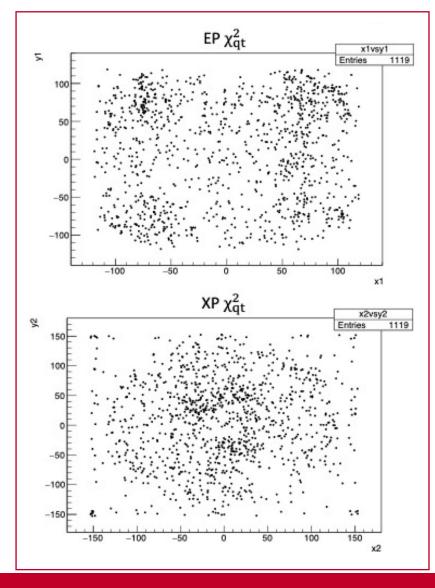
PMT SPECIFICATIONS

- Inner ring: 6 Hamamatsu R9980 PMTs
 - Mainly for the detection of scintillation light
 - Bias Voltage: 1300 V
- Outer ring: 10 ADIT L25D19 B546 PMTs
 - Mainly for the detection of Cherenkov light
 - Bias Voltage: 1250 V

MEASURES OF GOODNESS OF RECONSTRUCTION PROCEDURE

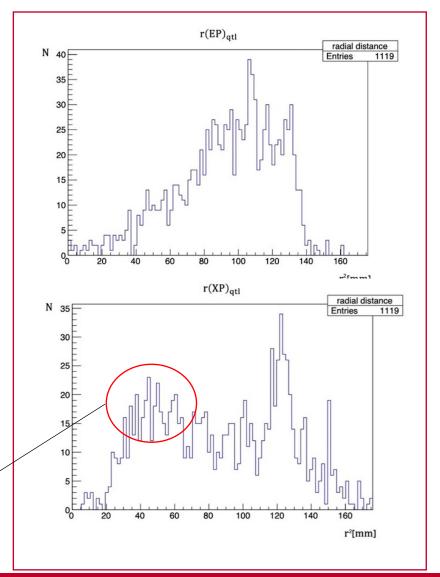
- Quantitative assessment: Compare reconstructed tracks with tracks obtained from muon tracker
- Perform qualitative assessment of track reconstruction algorithm using the distributions:
 - Entry Points
 - Exit Points
 - Radial distances of entry and exit points
 - Angular distributions Zenith and Azimuthal angle

DISTRIBUTION OF ENTRY AND EXIT POINTS



- Number of muons detected on a surface of radius $r \propto \pi r^2$
- Mostly uniform distribution
- Slight clustering of events towards the inner PMTs

Clustering near inner PMTS



ADDITIONAL IMPROVEMENTS

 Inserted black shields inside the test cell to minimize internal reflections

