

# Reconstruction of faint non-standard model particles in IceCube

Nick Jannis Schmeißer  
11.10.2023

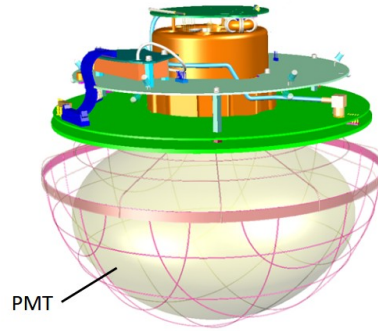


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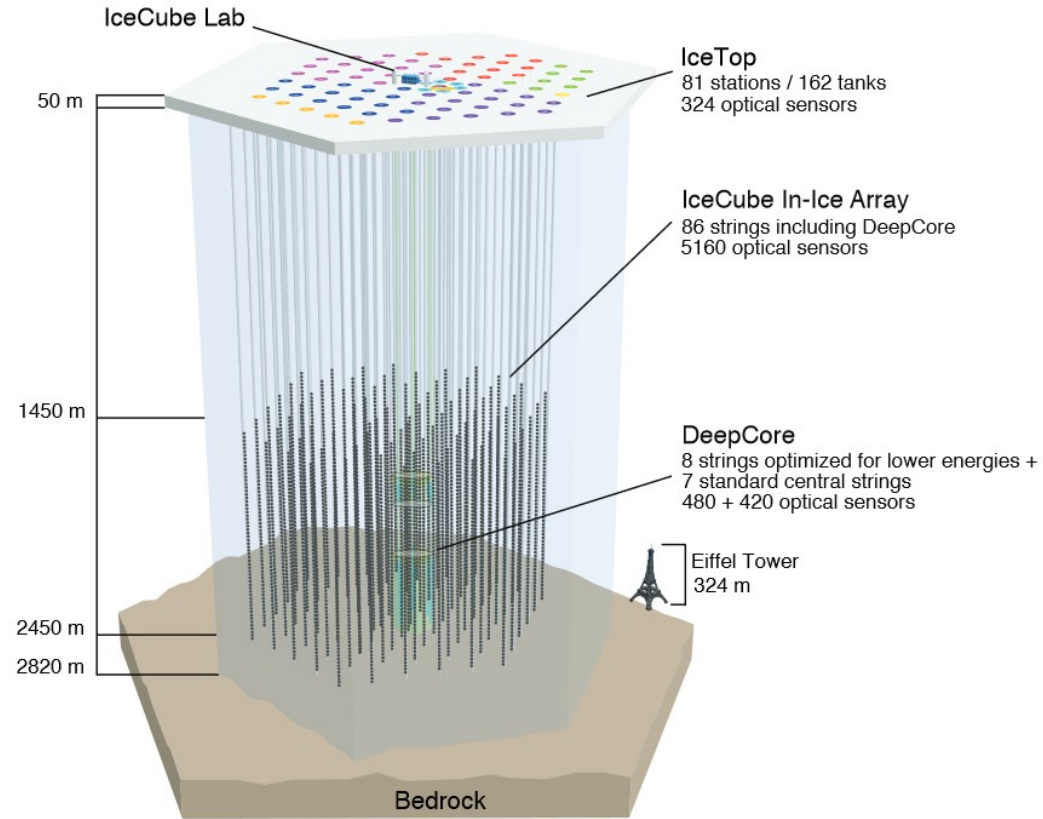


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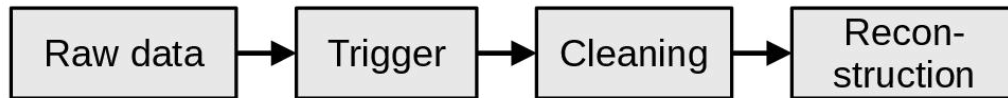
# The IceCube Neutrino Observatory



- Different types of events possible:
  - Track-like events
  - Spherical events, cascades
- Prior to high level analyses events need to be reconstructed, e.g. direction, position, energy



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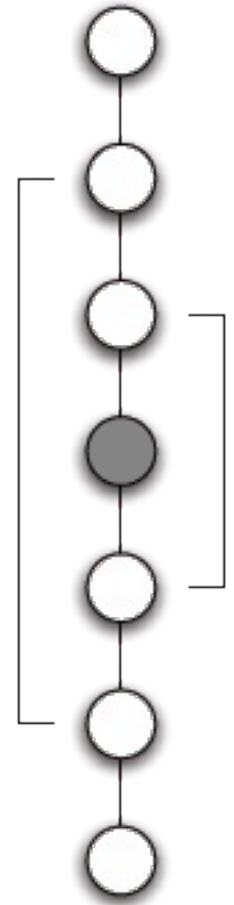


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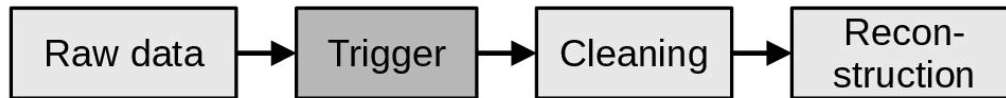
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# The faint particle trigger (FPT)

- Usual triggers based on local coincidences
- Problematic for search of **faint signals**,  
e.g. low energy events or fractionally charged particles  
→ produce **multiple isolated hits**
- Trigger uses isolated hits
- **Will be deployed for IceCube DeepCore in November**
- Looks for **direction and velocity consistent signal-combinations**
- **Existing reconstruction algorithms** also based on local coincidences → **needs to be adjusted**



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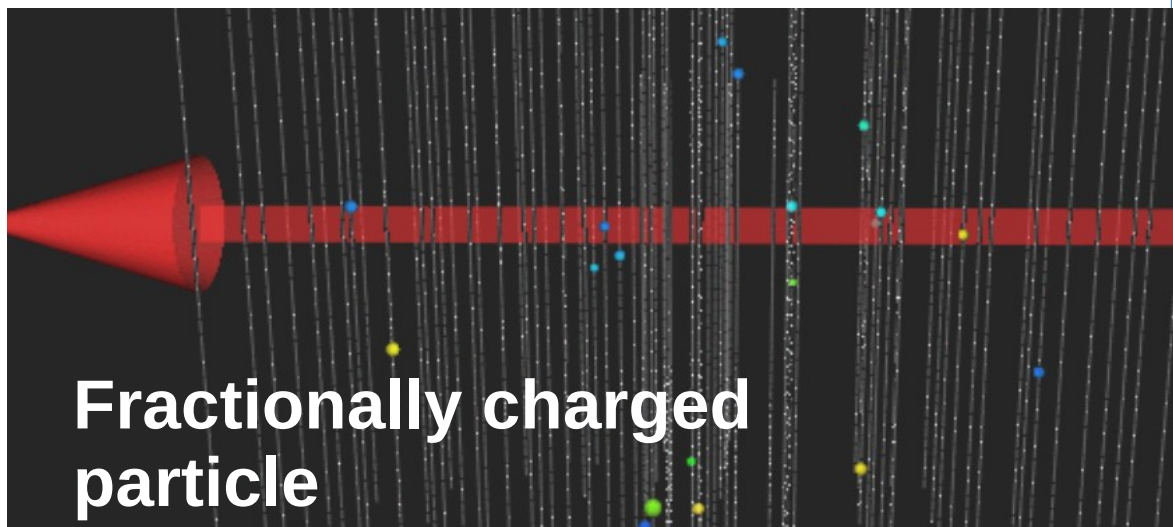
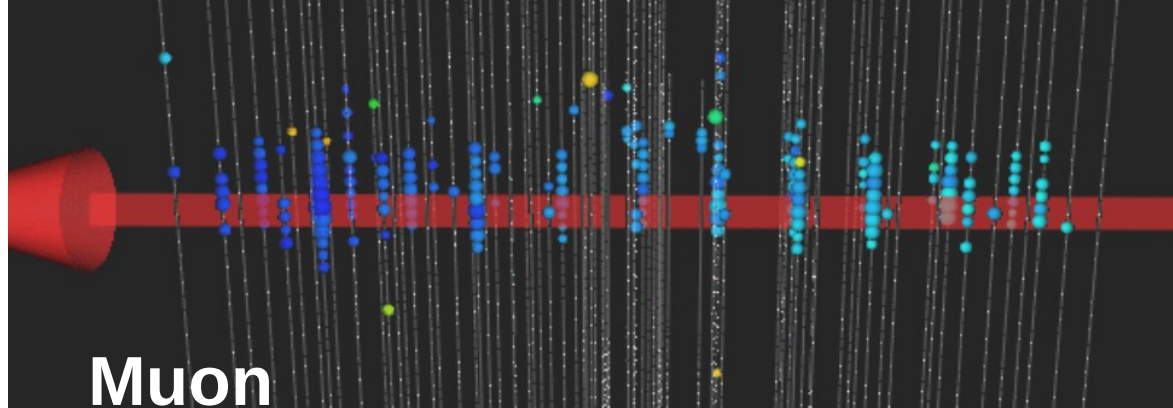
# Signal events

- **Simulate** fractionally charged particles

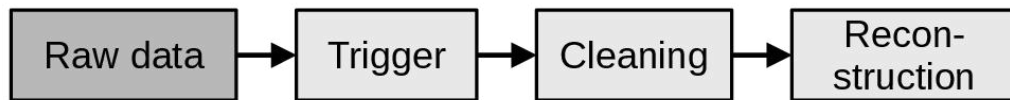
- Charge  $1/3 e$

→ lower signal-to-noise ratio

- Homogeneous distribution around detector
- Use faint particle trigger on events



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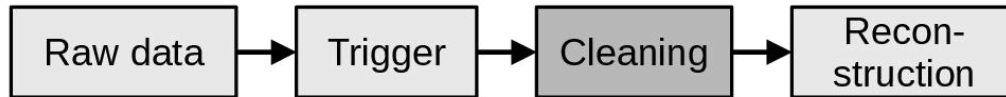
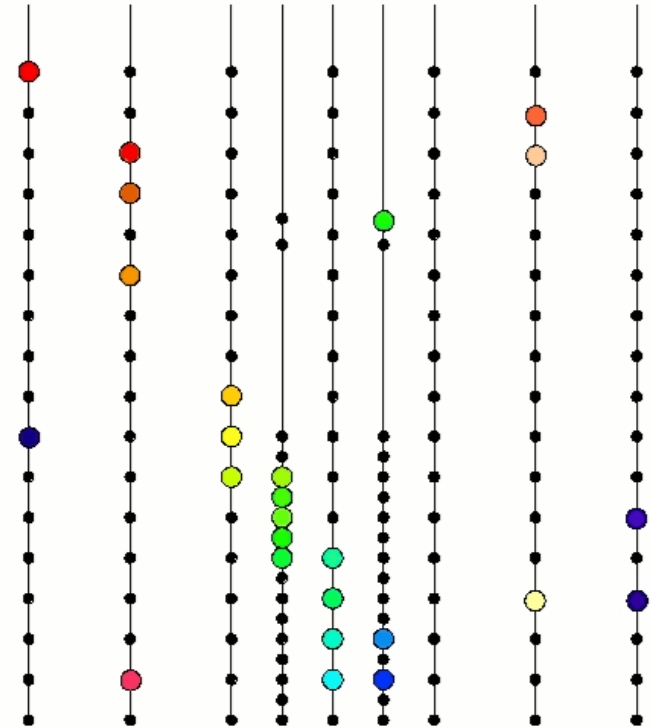


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# Seeded RT Cleaning

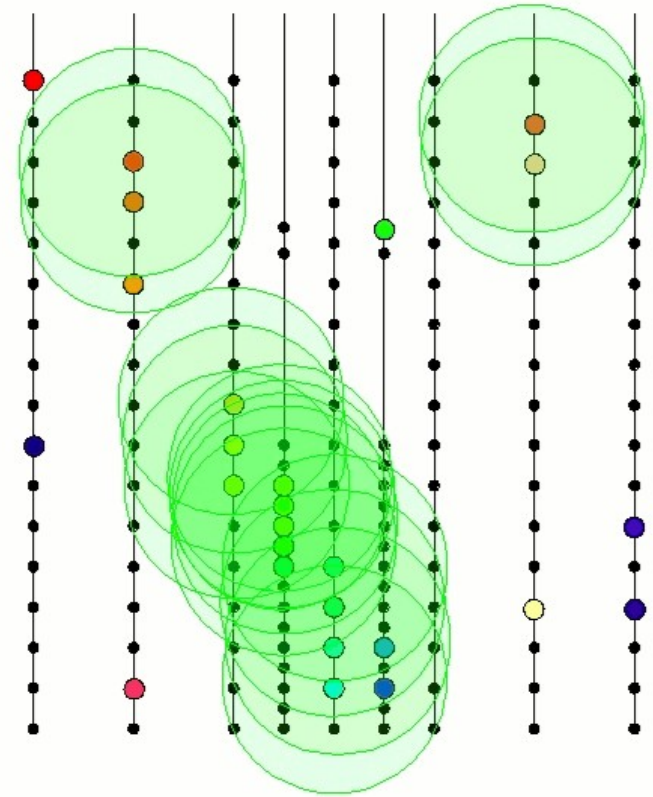
- Standard cleaning algorithm  
→ reduce noise before reconstructions
- Use local coincidences as seed
- Look for hit pairs within radius  $R$  and time  $T$
- Neglect other hits
- **Problem: Cleaning based on local coincidences, which are not likely in FPT-triggered events**



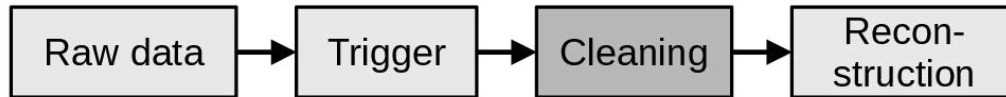
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Looking for hits in RT-range of HLC hits ...



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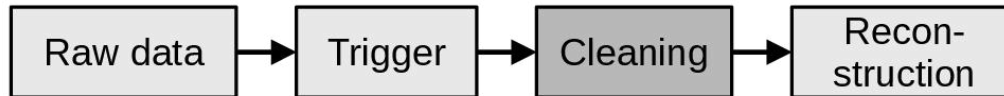
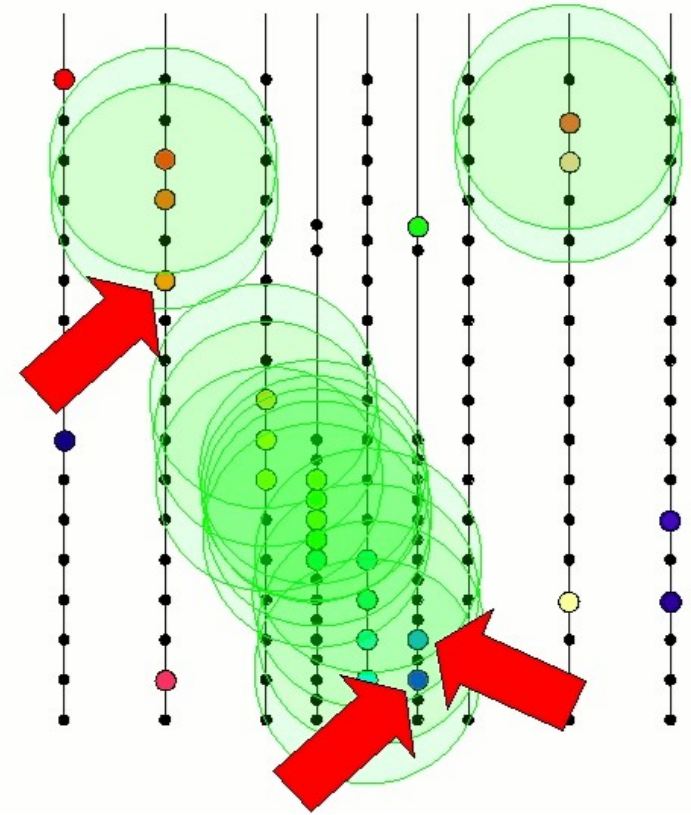
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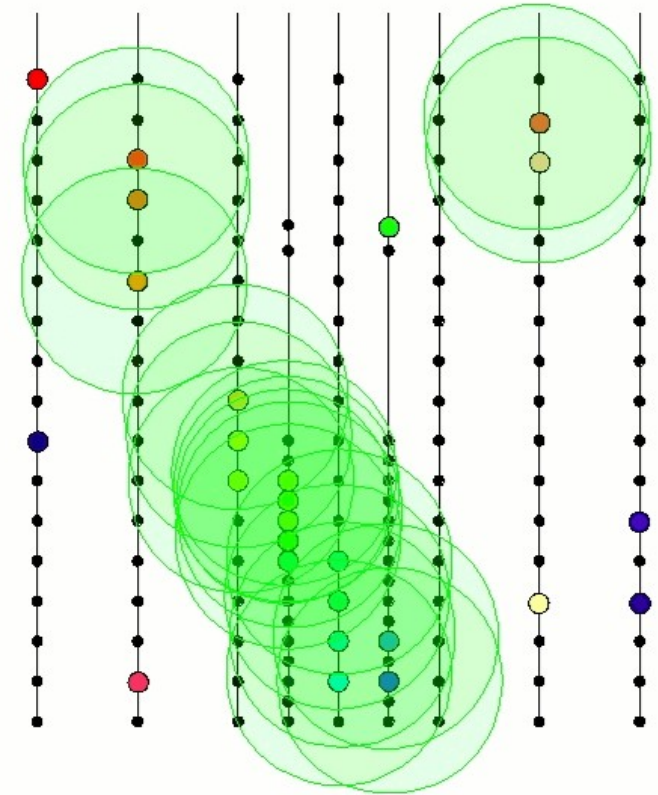
... keep these hits and look for further hits in their RT-range ...



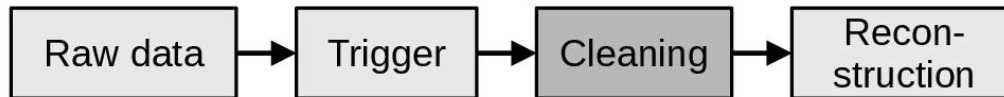
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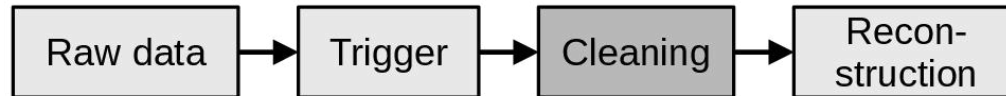
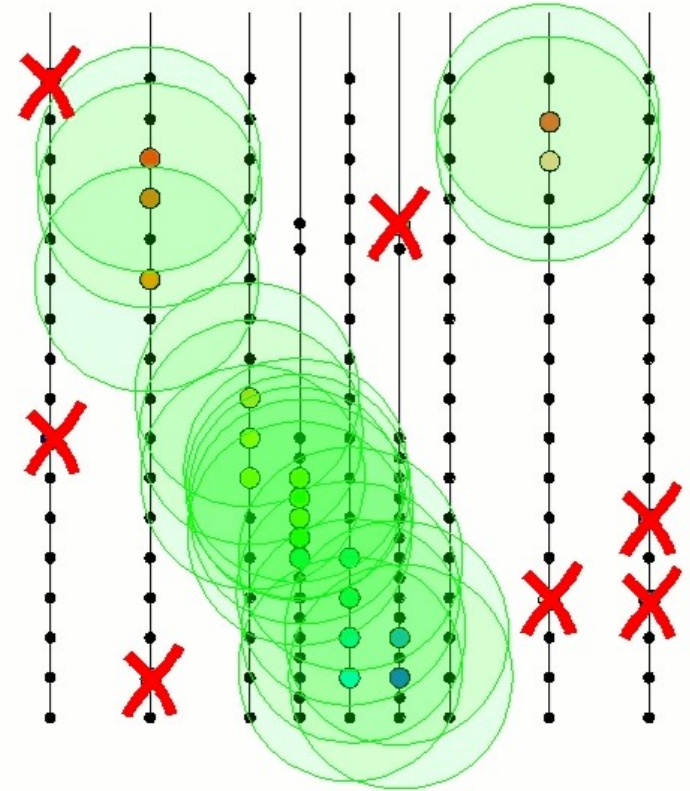
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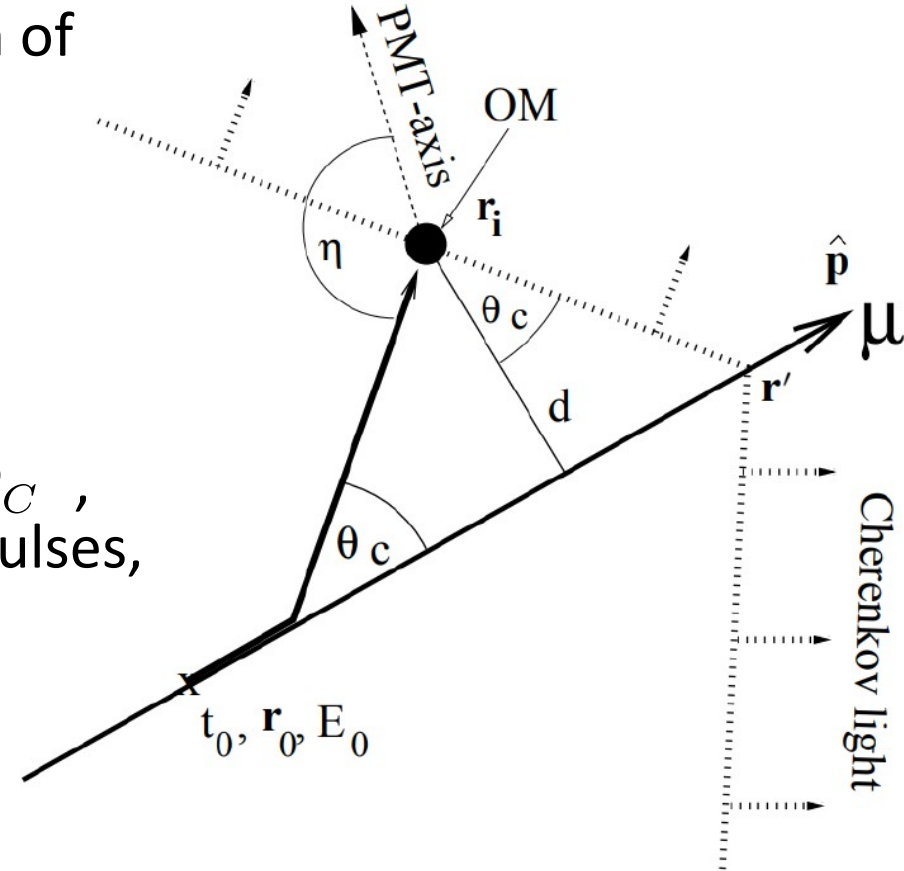
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... iterate until there are no more changes.



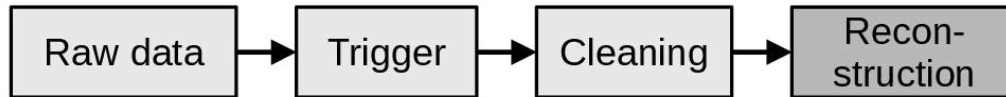
# Track reconstruction

- Used as a **basic direction reconstruction** of events
- Use **Line-Fit as seed**, which estimates simple particle track  $\vec{r}' = \vec{r} + \vec{v}t$  by assuming perpendicular wavefront
- Improve seed by including sensor-orientation  $\eta$ , Cherenkov-cone  $\theta_C$ , time-consistency, Multi-Photoelectron pulses, etc.
- Use **angular deviation  $\Delta\Phi$**  between reconstructed direction and simulated direction to **check precision of reconstruction**



Source: Van Driessche, Ward: Search for particles with anomalous charge in the IceCube detector, 2019, p. 123

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# Influence of cleaning (1)

## With Cleaning:

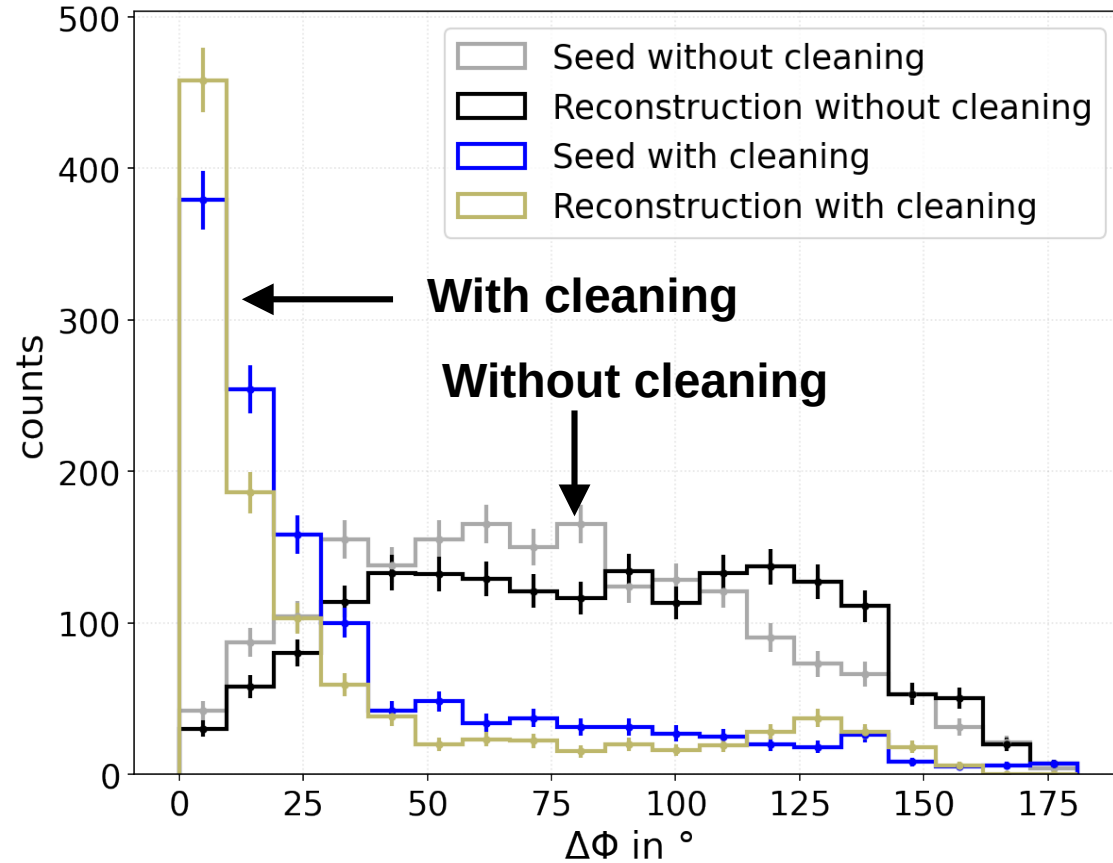
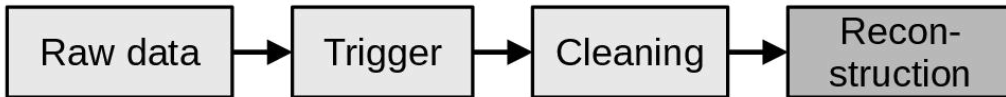
- **~1/3 of events are lost**
- Especially events only triggered by FPT → no local coincidences

## Without Cleaning:

- Reconstruction much worse
- Data dominated by noise

Notice: Seed also not precise

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# Influence of cleaning (2)

Idea:

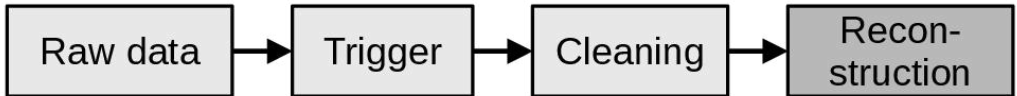
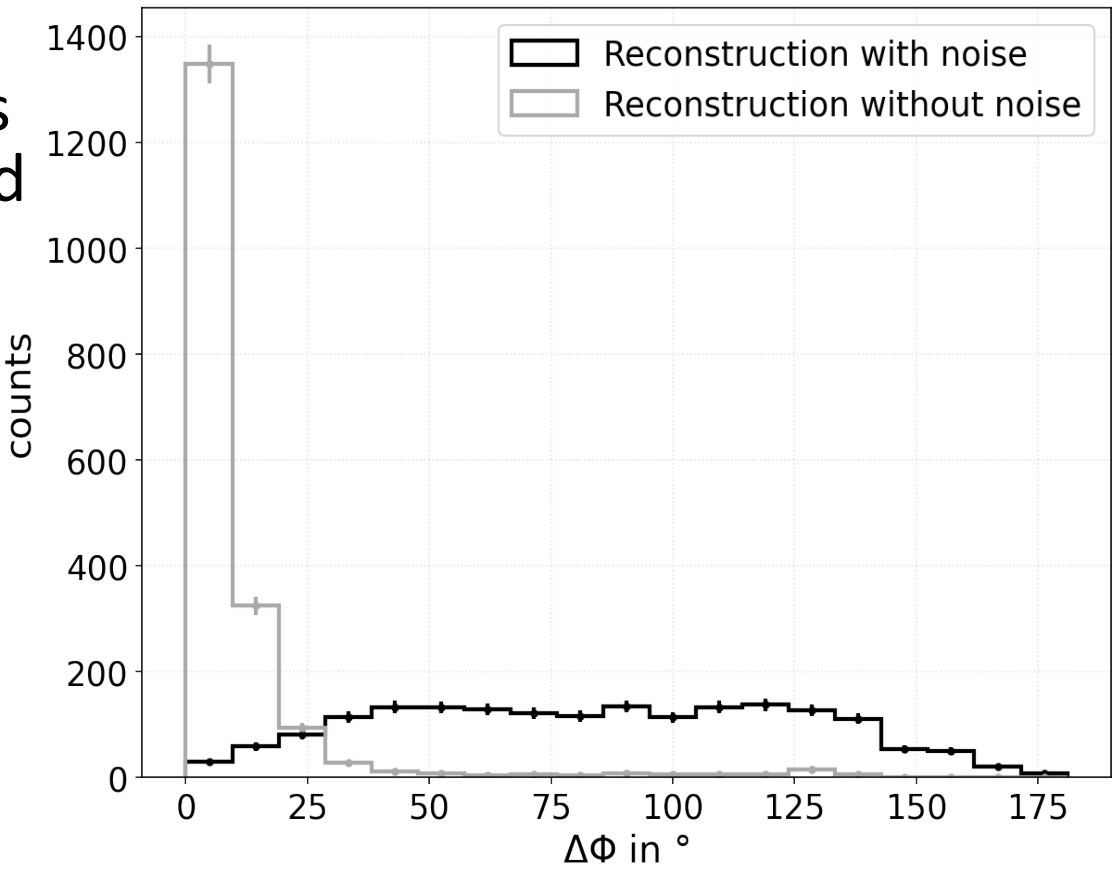
Adjust Cleaning to new events  
→ Include isolated hits as seed

Problem:

Not possible with  
existing algorithm  
→ work in progress

Goal:

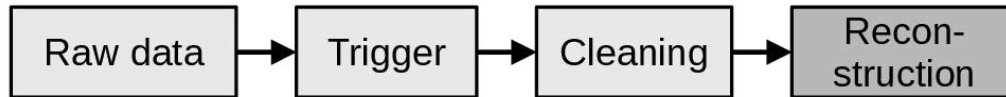
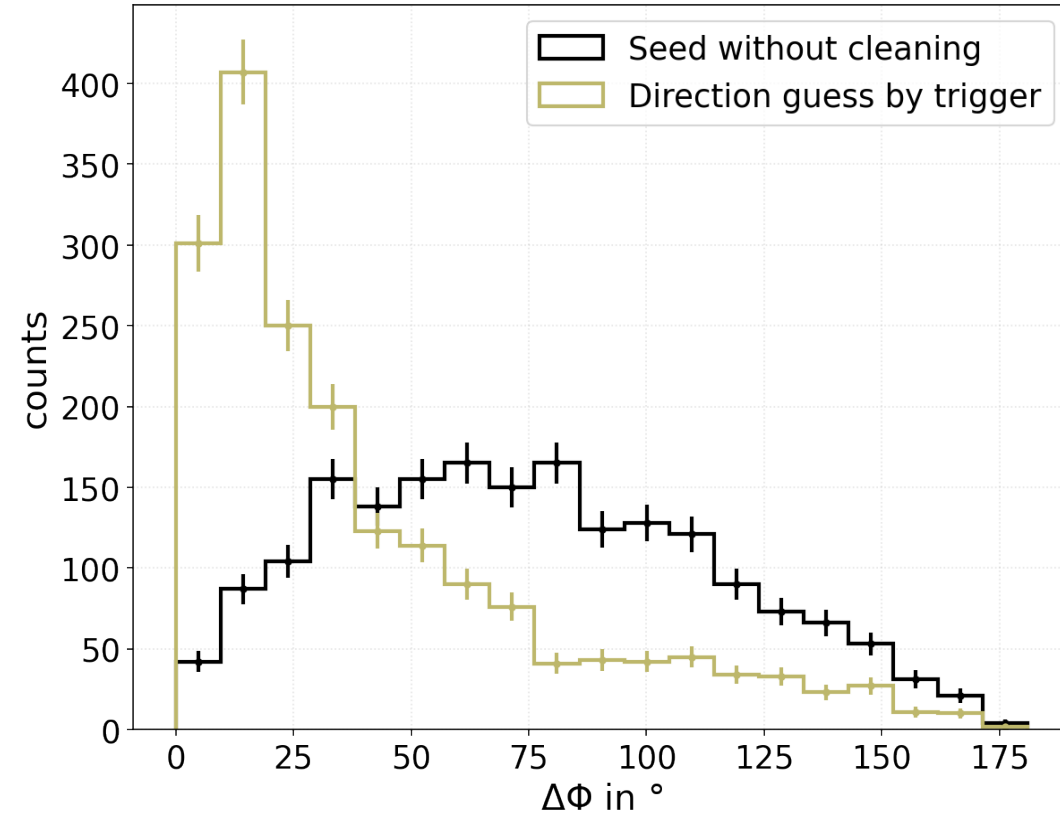
Reduce noise hits





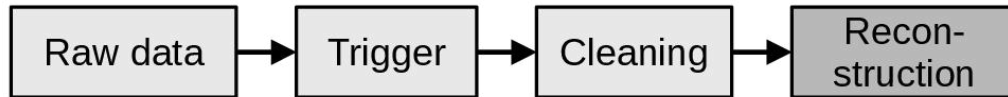
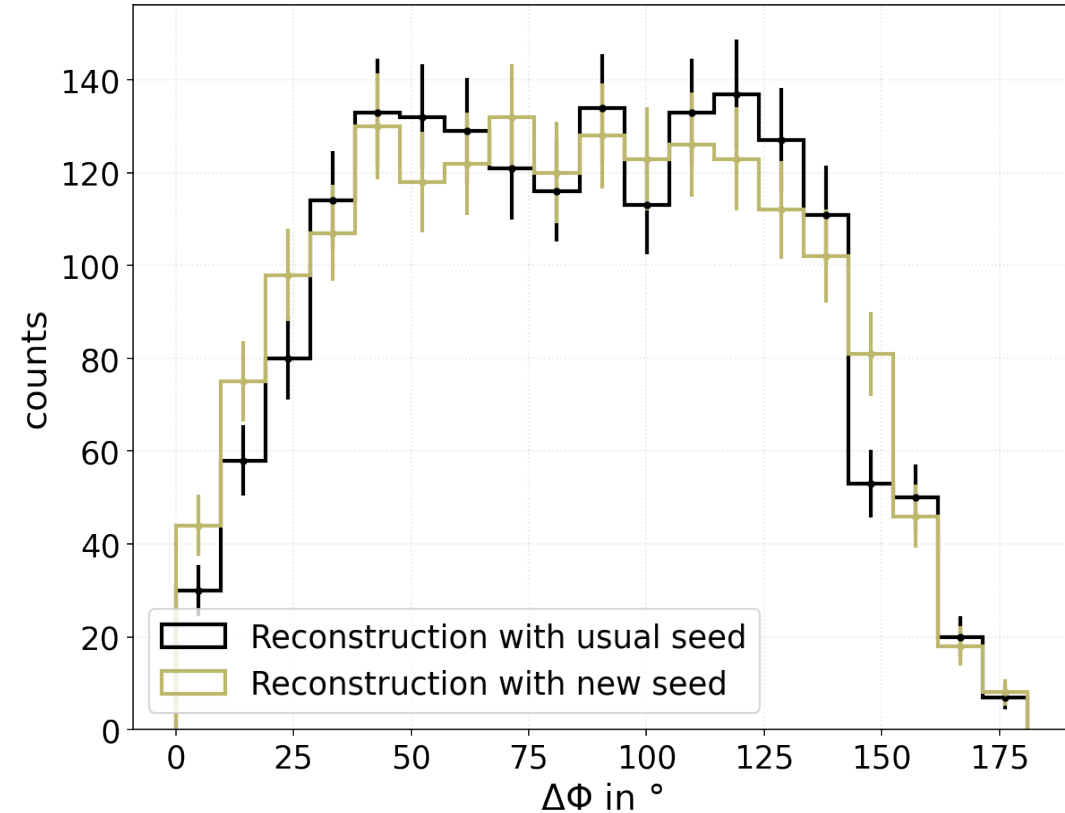
# Direction guess by trigger (1)

- Faint particle trigger able to give first direction guess
- Looks at clustering of hit pairs in zenith and azimuth
- Better as seed which is used without the cleaning
- Can maybe be used as seed for reconstruction



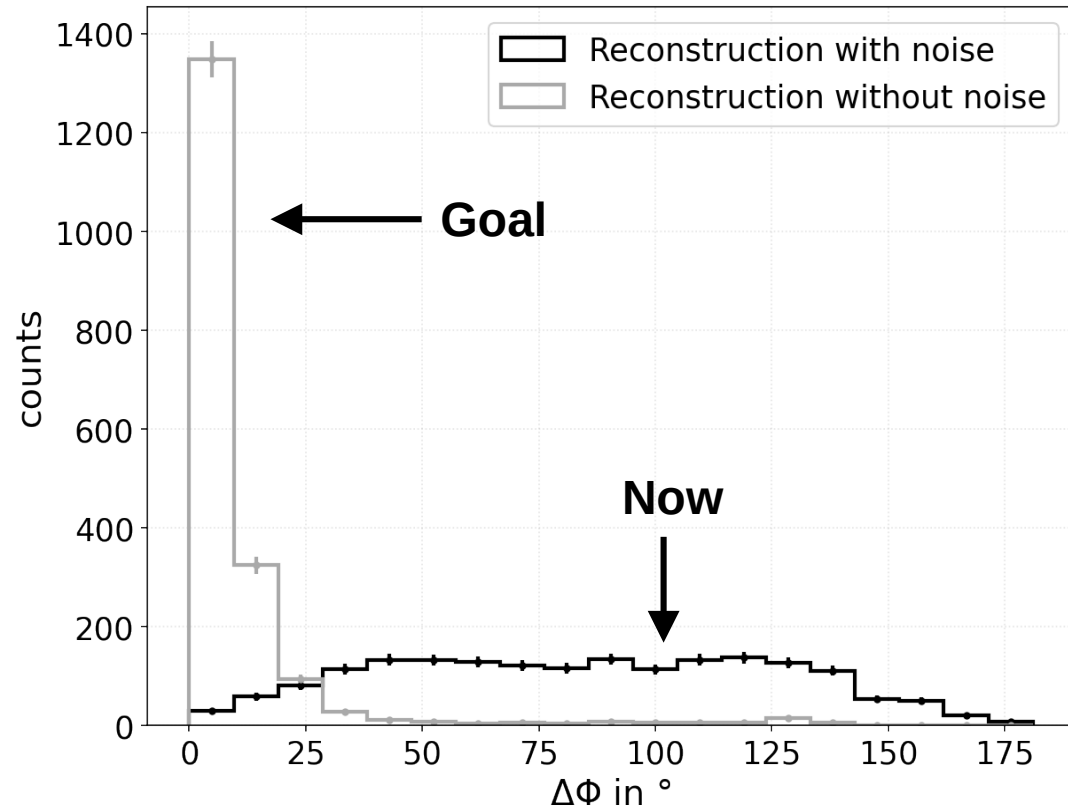
# Direction guess by trigger (2)

- Use direction guess as seed for reconstruction
- Gives small improvement
- Reconstruction still bad  
→ still influenced by noise



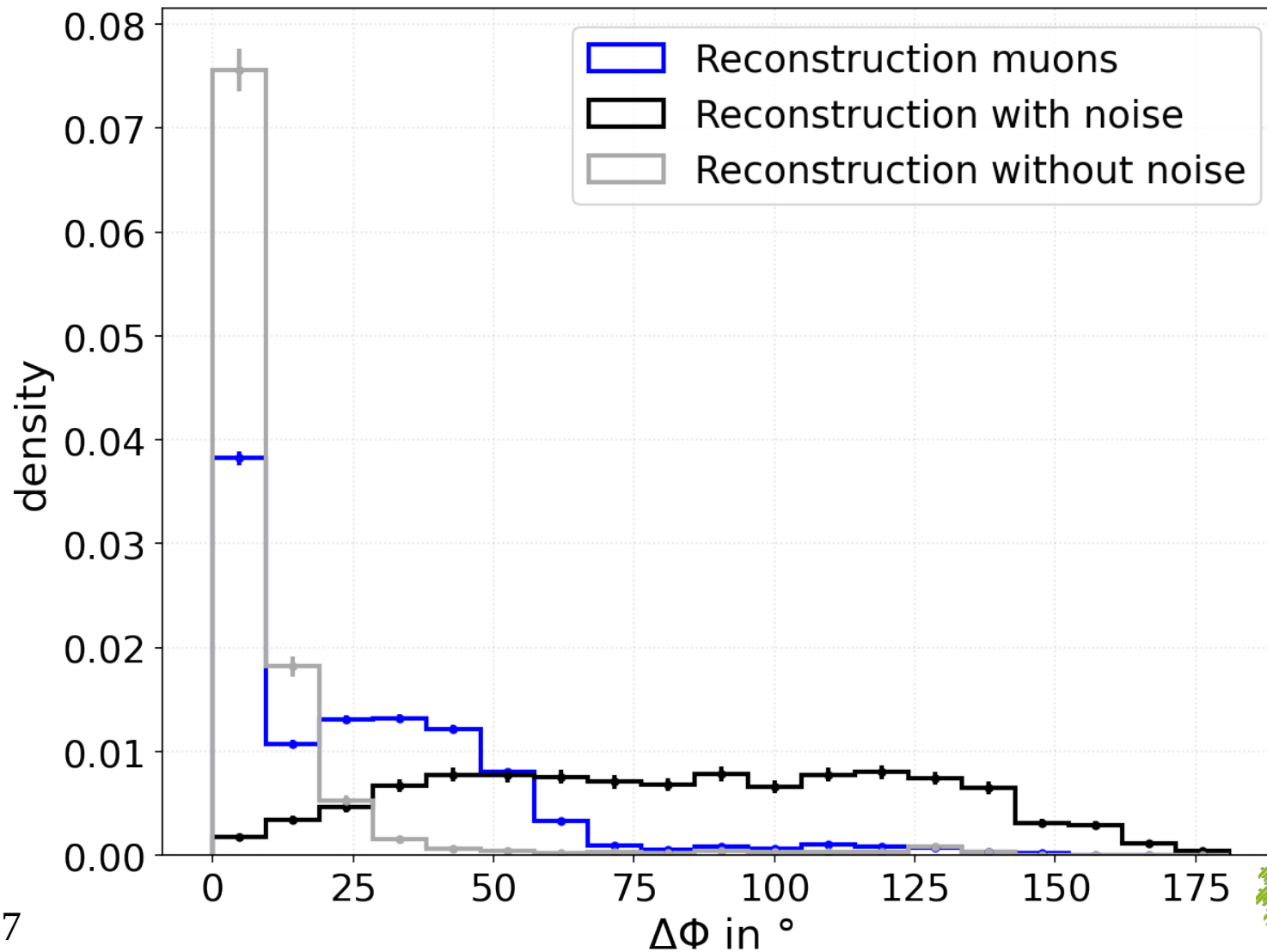
# Conclusion and outlook

- Adjustment of existing reconstruction algorithms needed  
→ especially cleaning
- Maybe more information from trigger can be used as seed
- Check energy reconstruction
- Goal: Implement adjustments into standard processing pipeline so that events by new trigger can be used for analysis
- Later: Reconstruction with graph neural network for comparison



# Back-Up





# Line-Fit

- Estimate simple particle track  $\vec{r}' = \vec{r} + \vec{v}t$
- Assumes perpendicular wavefront traveling with particle
- Get analytical solutions for  $r$  and  $v$  by minimizing

$$S(\vec{r}, \vec{v}) = \sum_{i=1}^{N_{hit}} \rho^2 = \sum_{i=1}^{N_{hit}} (\vec{r}_i - \vec{r}')^2 = \sum_{i=1}^{N_{hit}} (\vec{r}_i - \vec{r} - \vec{v}t_i)^2$$

- $N_{hit}$  - number of hit modules
- Analytical solution:

$$\vec{r} = \langle \vec{r}_i \rangle - \vec{v} \langle t_i \rangle \quad \vec{v} = \frac{\langle \vec{r}_i t_i \rangle - \langle \vec{r}_i \rangle \langle t_i \rangle}{\langle t_i^2 \rangle - \langle t_i \rangle^2}$$

# Improved Line-Fit

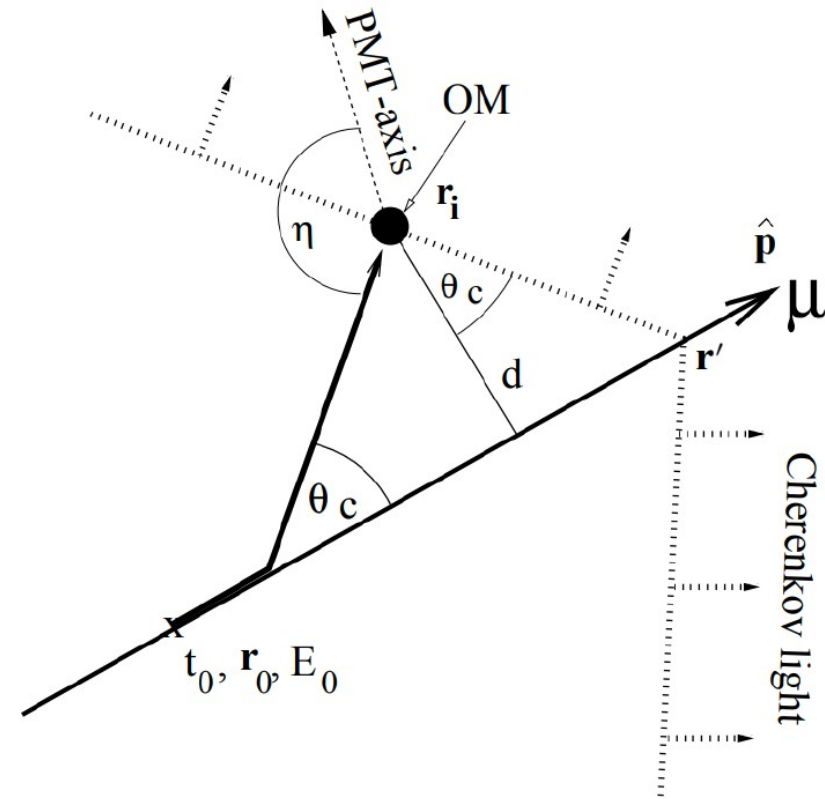
- Line-Fit doesn't account for the Cherenkov profile, scattering effects and noise hits which can occur far from the track
- Can improve accuracy by removing scattered hits with simple filter
- Algorithm looks for each hit  $h_i$  at all hits within neighborhood of radius  $\mu$  and if there is an earlier hit in this neighborhood then  $h_i$  is considered a scattered hit

- Calculate track by minimizing 
$$\sum_{i=1}^{N_{hit}} \phi(\rho) = \sum_{i=1}^{N_{hit}} \phi(|\vec{r}_i - \vec{r}'|)$$
with 
$$\phi(\rho) \equiv \begin{cases} \rho^2 & \text{if } \rho < \mu \\ \mu(2\rho - \mu) & \text{if } \rho \geq \mu \end{cases}$$

- Used as seed for following reconstructions

# SPE Fit

- Uses Line-Fit as a seed for reconstruction
- Takes Cherenkov cone into account as well as time smearing effects due to noise effects, PMT jitter, light from secondary interactions, DOM orientation, etc.
- Initial particle position and direction found by minimizing likelihood and iterating a couple of times  $\rightarrow$  global maximum
- Single PhotoElectron (SPE) fit





# MPE Fit

- Multi-PhotoElectron (MPE) fit
- Takes into account that only first pulse is recorded if multiple photons hit module in a short period of time
- Early photons scattered less in the ice
- For this likelihood function is adjusted by replacing single photon p.d.f.
- General comment: SPE-Fit and MPE-Fit in Offline Filter are normally only activated if certain filters were passed (muon, deepcore, highQ, etc. )

# Calculation of angular diviations

- From all fits as well from the original simulated particles one gets a zenith ( $\vartheta$ ) and azimuth ( $\phi$ ) angle
- Convert angles into unit vector with cartesian coordinates
- Calculate angle between different unit vectors

$$x = \sin(\vartheta) \cos(\phi)$$

$$y = \sin(\vartheta) \sin(\phi)$$

$$z = \cos(\vartheta)$$

$$\alpha = \arccos\left(\frac{\vec{a} \cdot \vec{b}}{|\vec{a}| |\vec{b}|}\right)$$

# Seeded RT Cleaning

- Cleaning algorithm to reduce noise&outliers
- Use a subgroup of all hits as seeds → HLC-hits as standards
- For each seed hit look for hits within a radius  $R$  and time distance  $T$  → add them to the seed list
- Standard definitions:  $R=150$  m,  $T=1000$  ns
- Iterate this multiple times
- Remove other hits