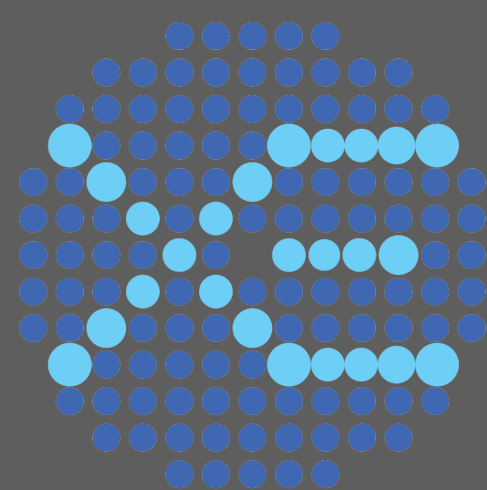


# Fast-track simulations in XENONnT

Astroparticle School 2022



Bundesministerium  
für Bildung  
und Forschung



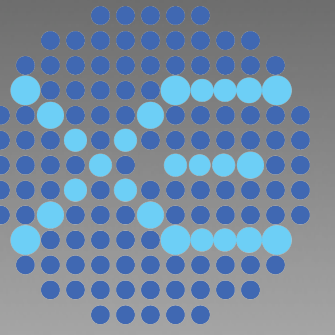
XENON



UNIVERSITÄT  
FREIBURG

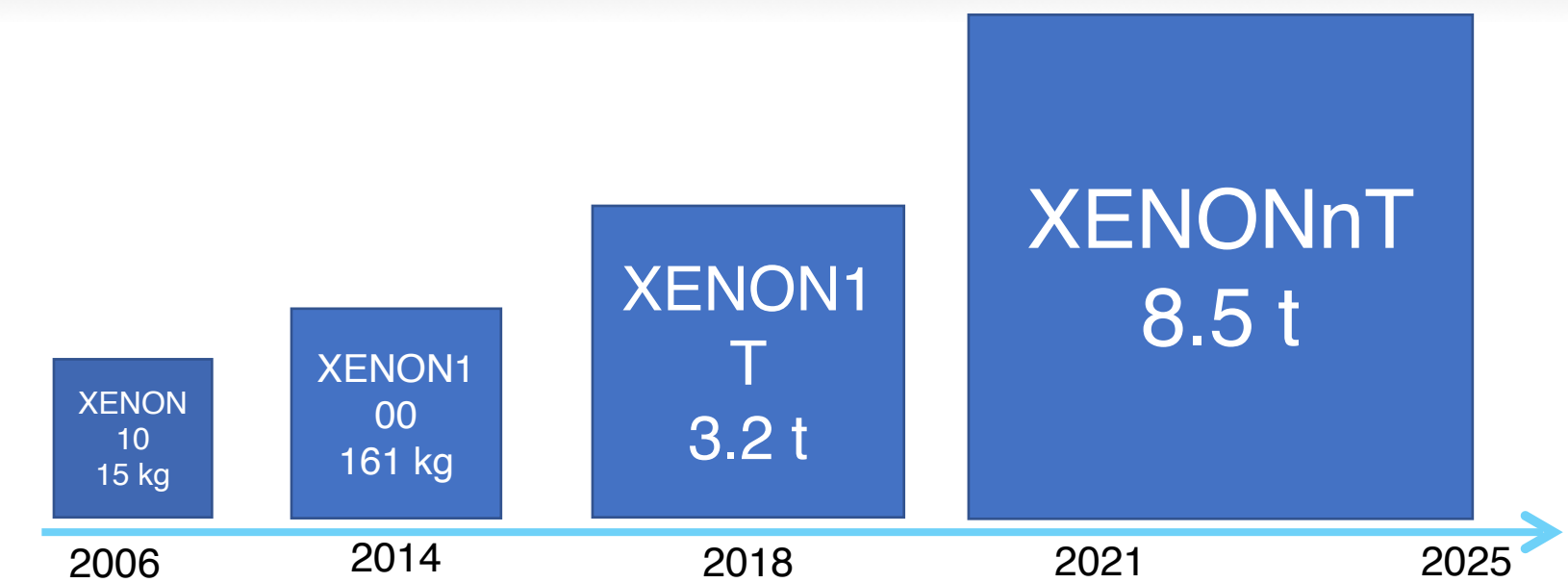
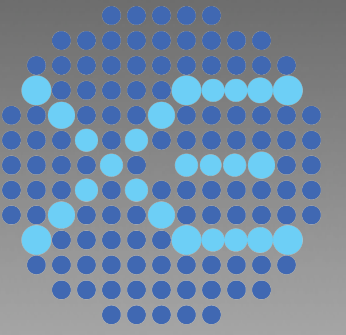
Jaron Grigat, [jaron.grigat@physik.uni-freiburg.de](mailto:jaron.grigat@physik.uni-freiburg.de)

# That's me



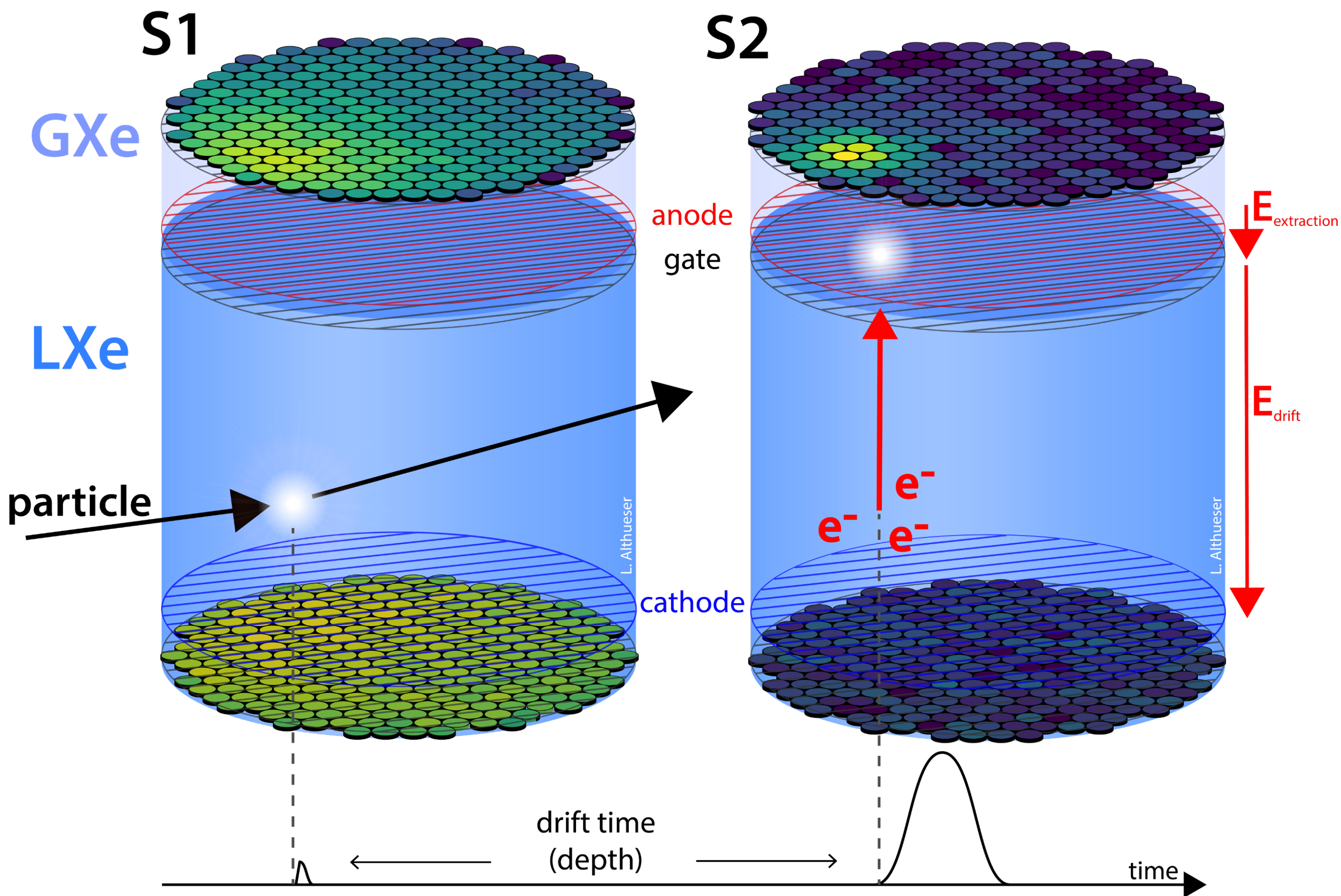
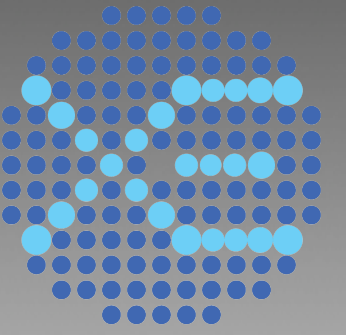
- Jaron Grigat
- Phd student since 2 years @ Uni Freiburg
- Group of Prof. Marc Schumann
- Also part of the XENON collaboration

# The XENON experiment



- Long tradition of building world-leading detectors for direct dark matter search (WIMPs)
- Many other science channels (e.g. solar axions,  $0\nu\beta\beta$ , ...)

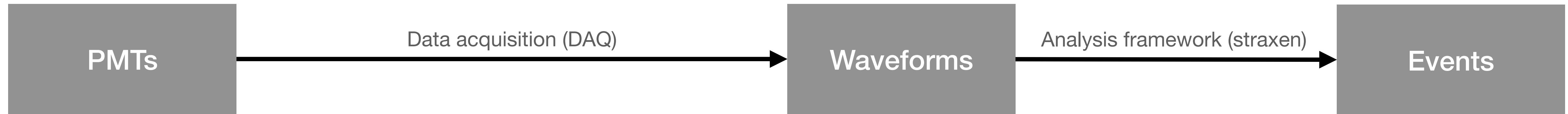
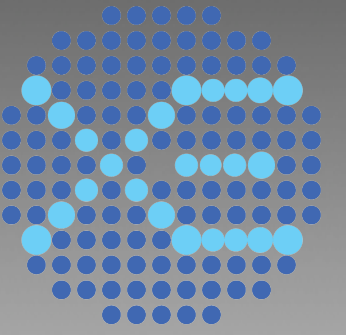
# Dual-phase time projection chamber (TPC)



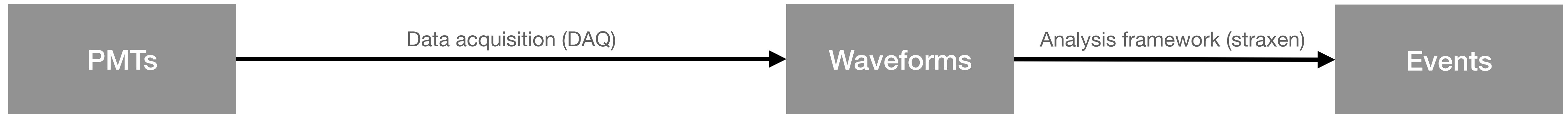
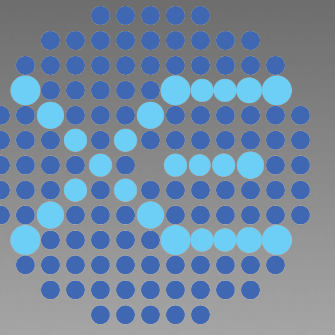
Combination of S1 and S2 signals allows for:

- 3D Position reconstruction
- Energy reconstruction
- discrimination between electronic and nuclear recoils

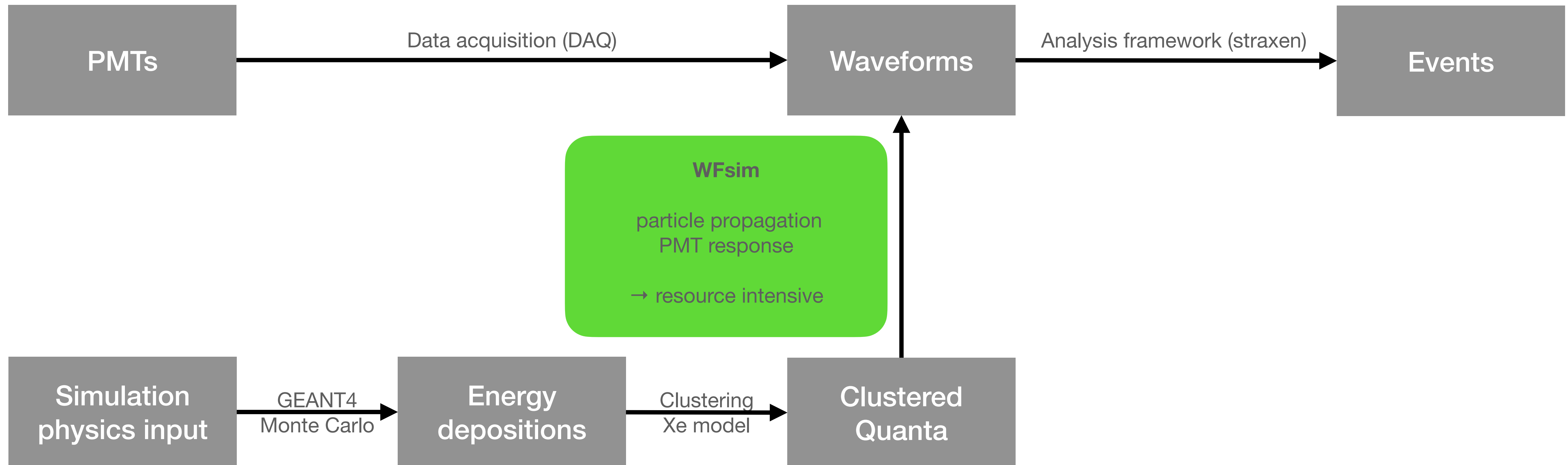
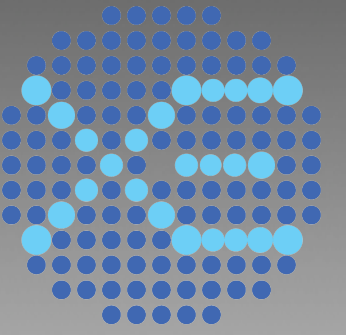
# Data processing chain



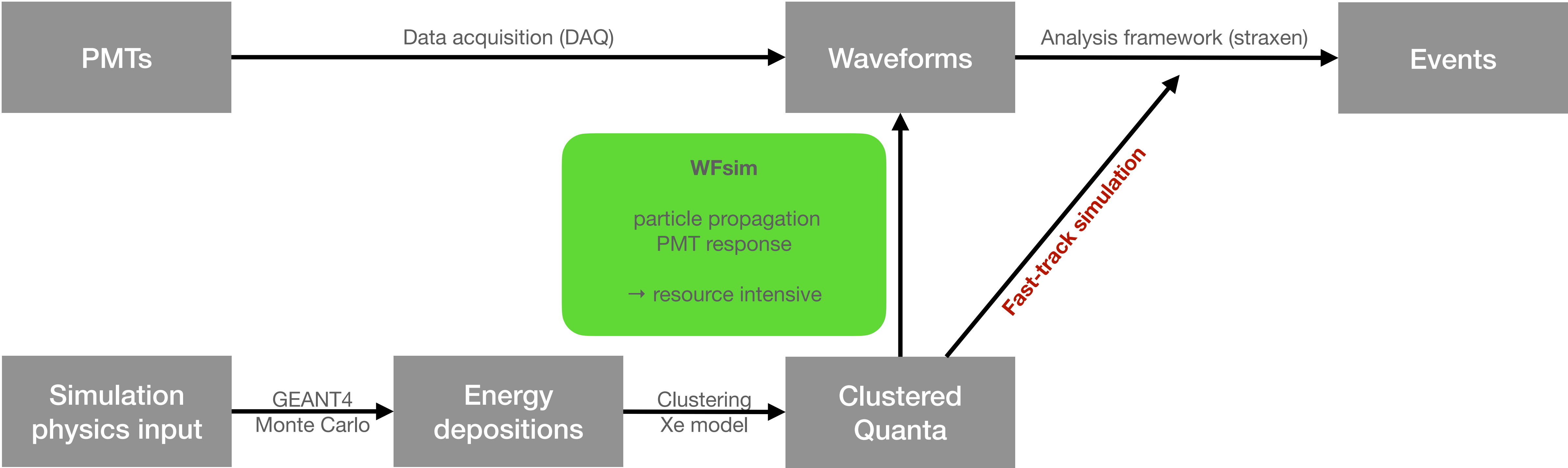
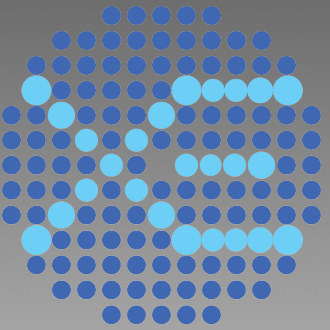
# Data processing chain



# Data processing chain

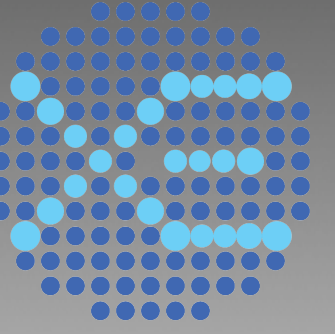


# Data processing chain



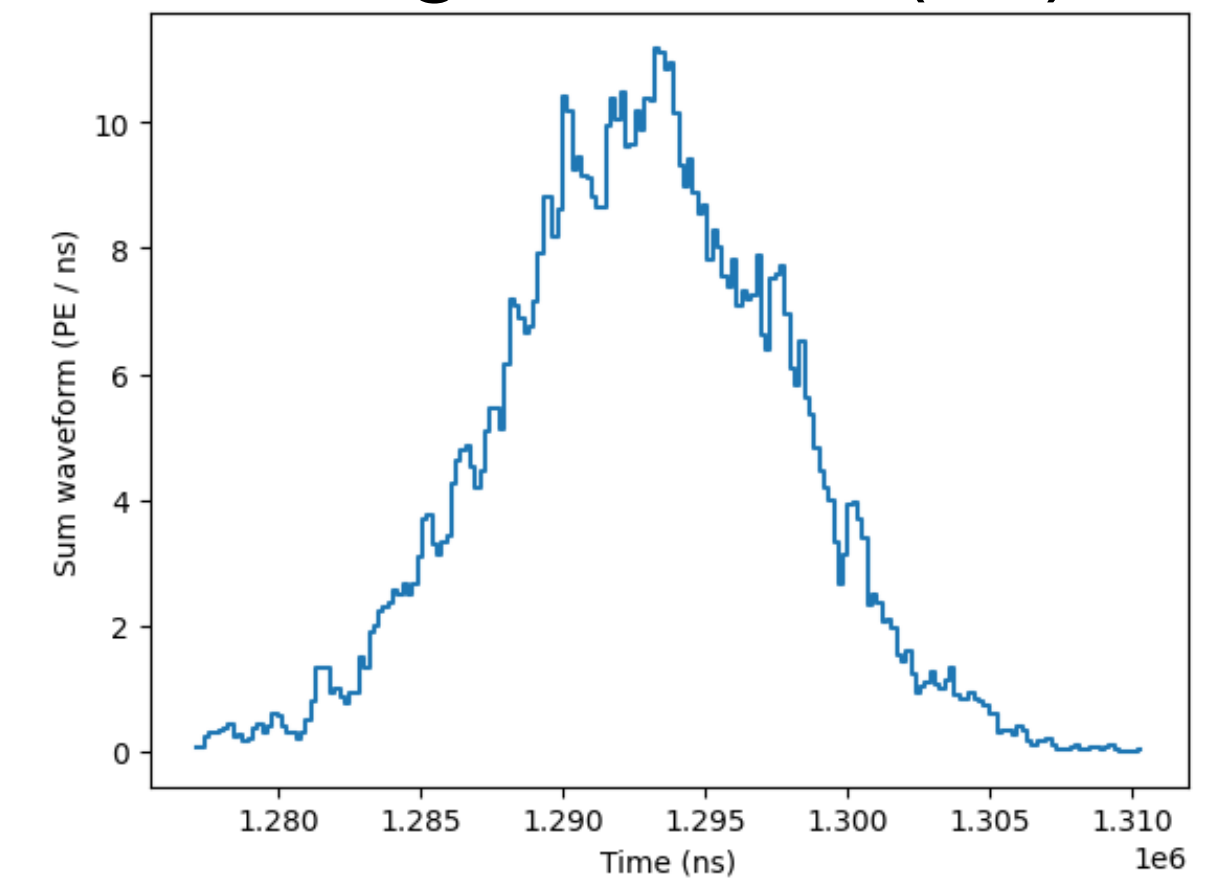


# Multi-scatter discrimination

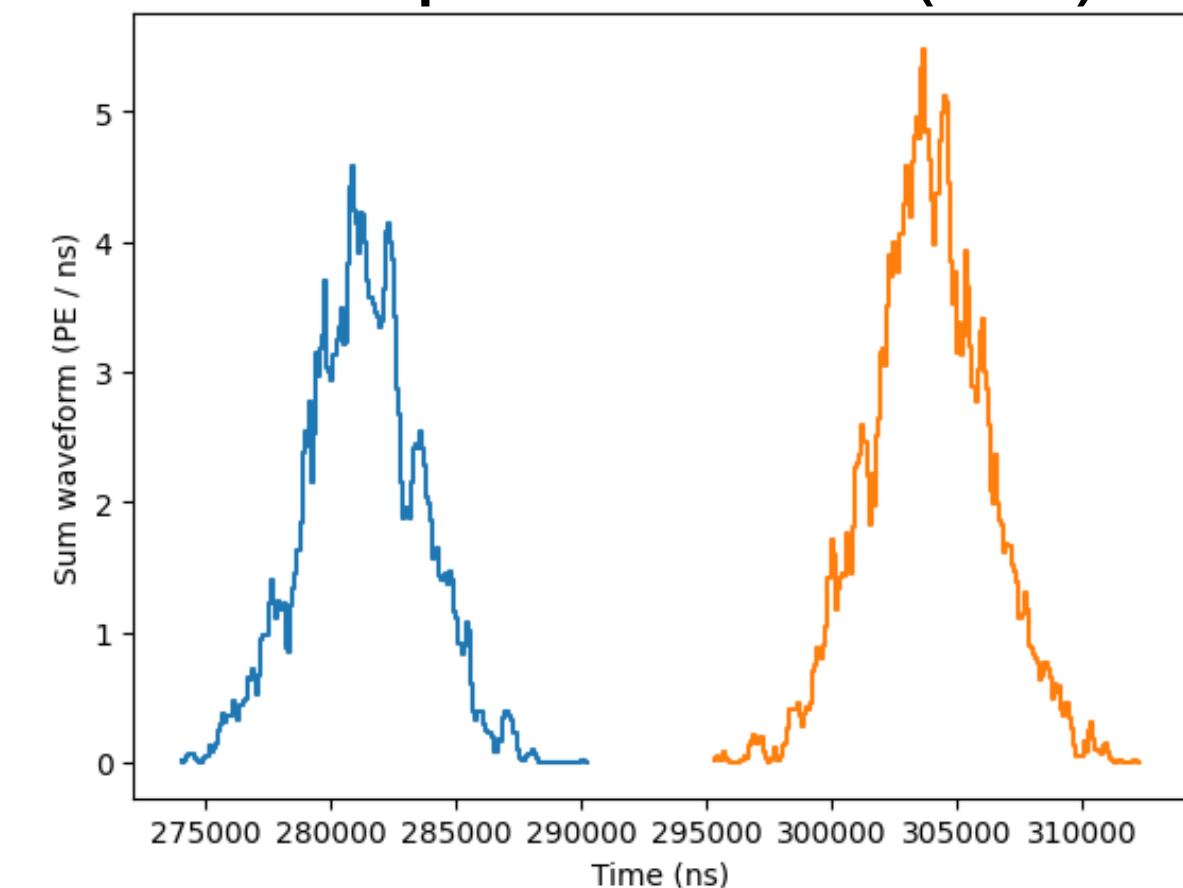


- We differentiate between events with a single scatter (SS) and events with multiple energy depositions (MS)
- Understanding the detectors efficiency to differentiate between SS & MS events is crucial for all science channels

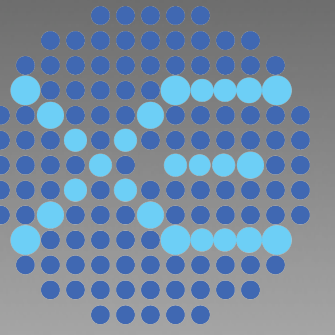
Single Scatter (SS)



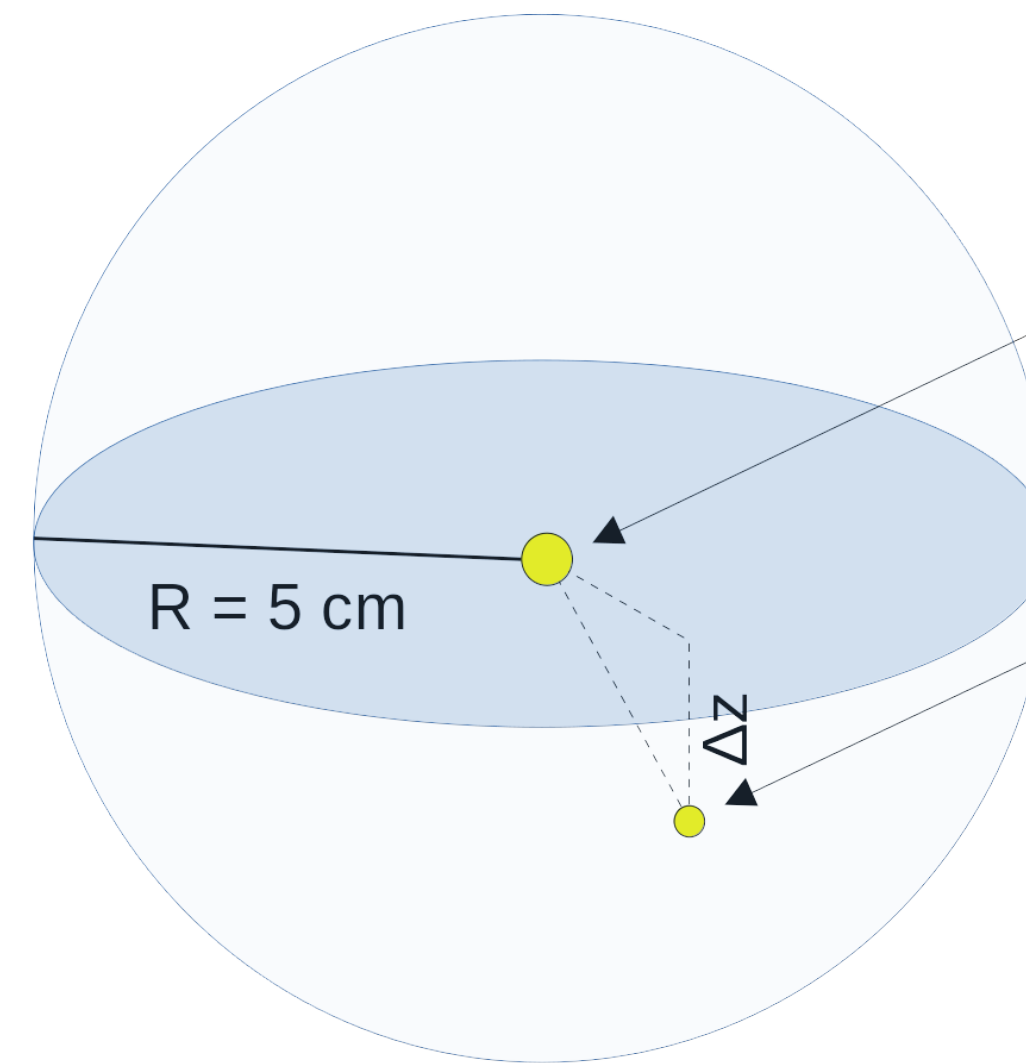
Multiple Scatter (MS)



# General idea

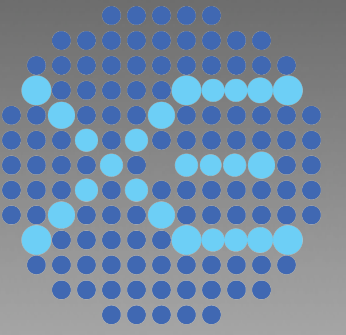


- Simulate Waveforms of events with two S2s (WFsim)
- Check if both signals are reconstructed as individual peaks
- Identify parameters that influence separation efficiency
- Define function that predicts separation based on these parameters



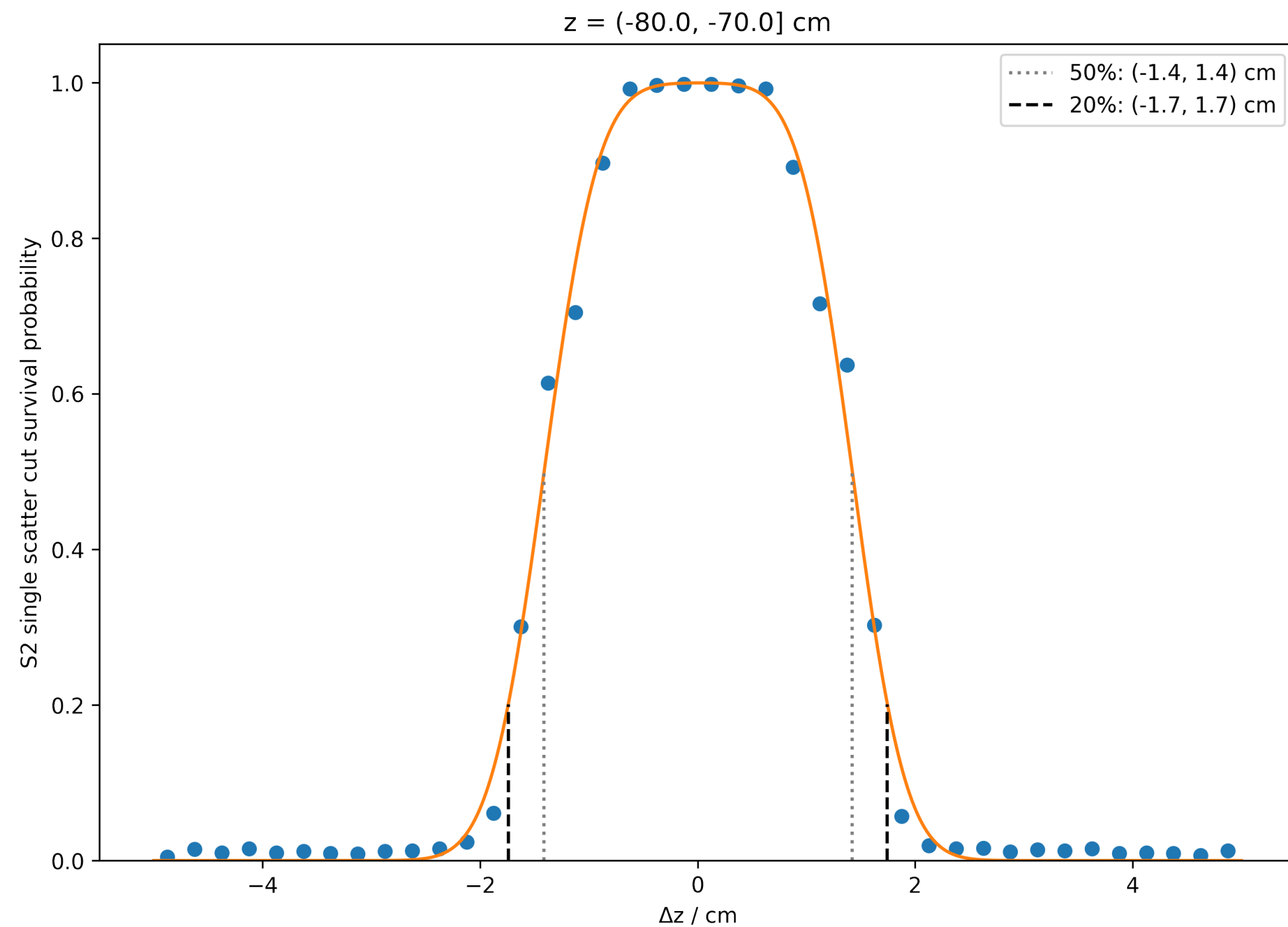
1. Generate S1 & S2a:
  - homogeneously in TPC
  - $E \in [1, 300] \text{ keV}_{ER}$
2. Generate S2b:
  - homogeneously distributed in sphere ( $R = 5 \text{ cm}$ ) around primary interaction site
  - $N_e(S2_b) \in [0.1, 0.9] N_e(S2_a)$

# Data analysis

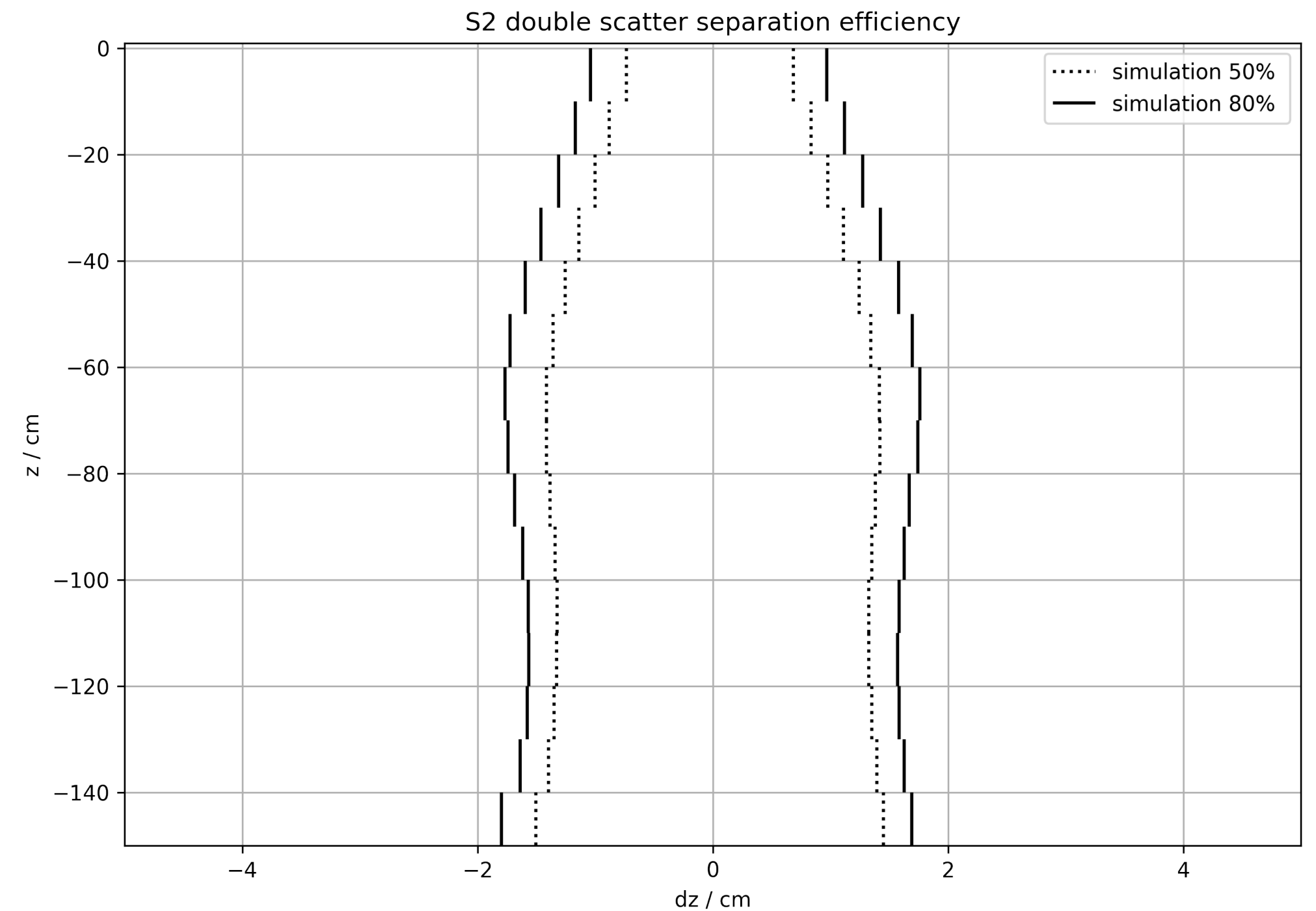


We can for example look at the MS-discrimination power as a function of the depth-difference of the two S2s ( $\Delta z$ )

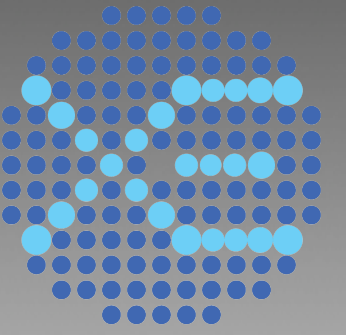
For one slice in depth ( $z$ ):



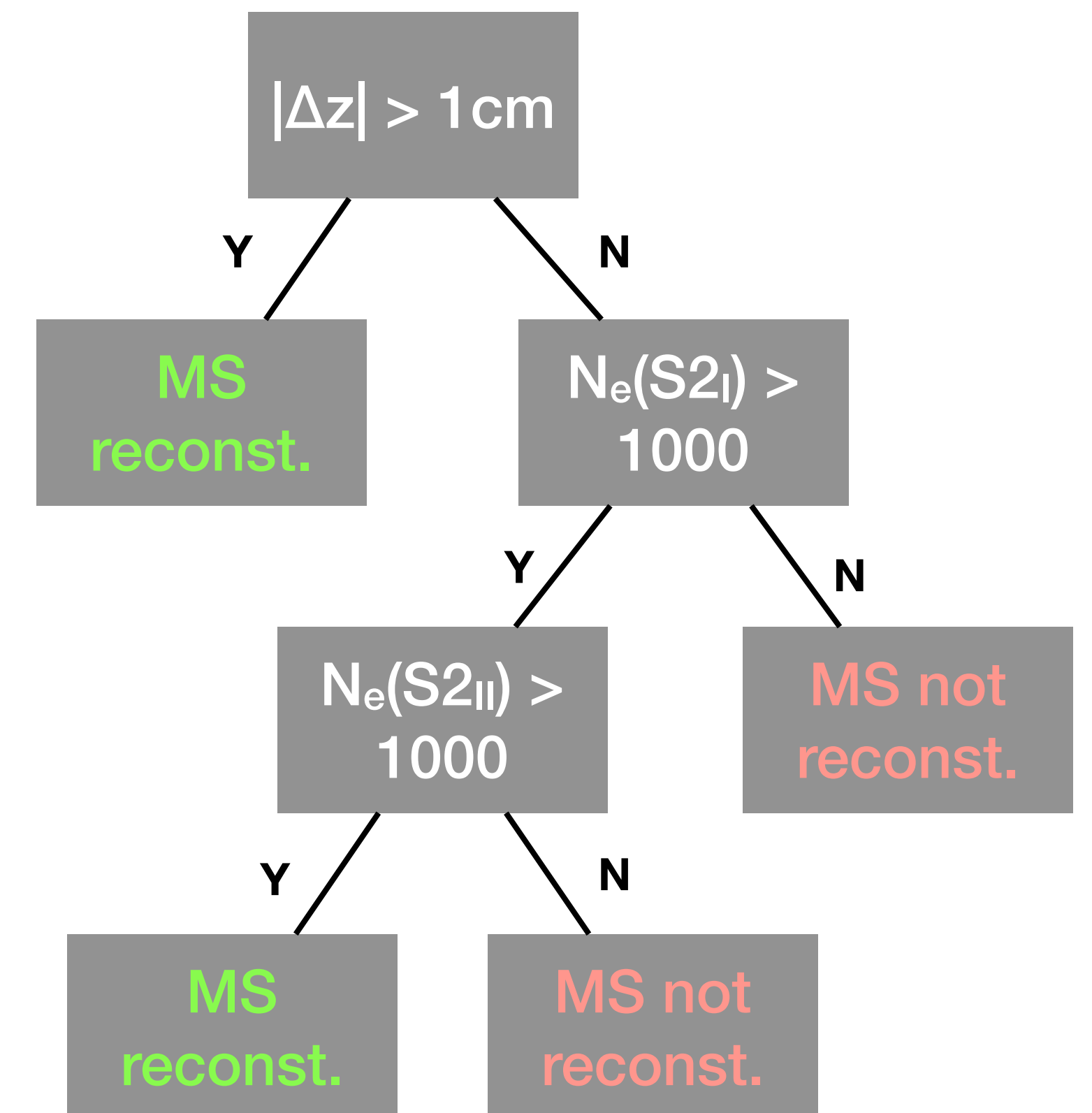
In the entire detector:



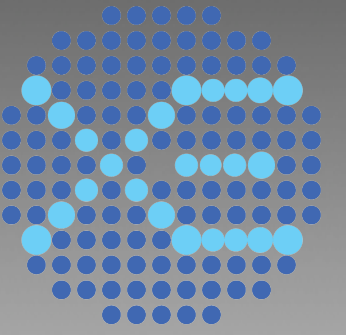
# Machine learning



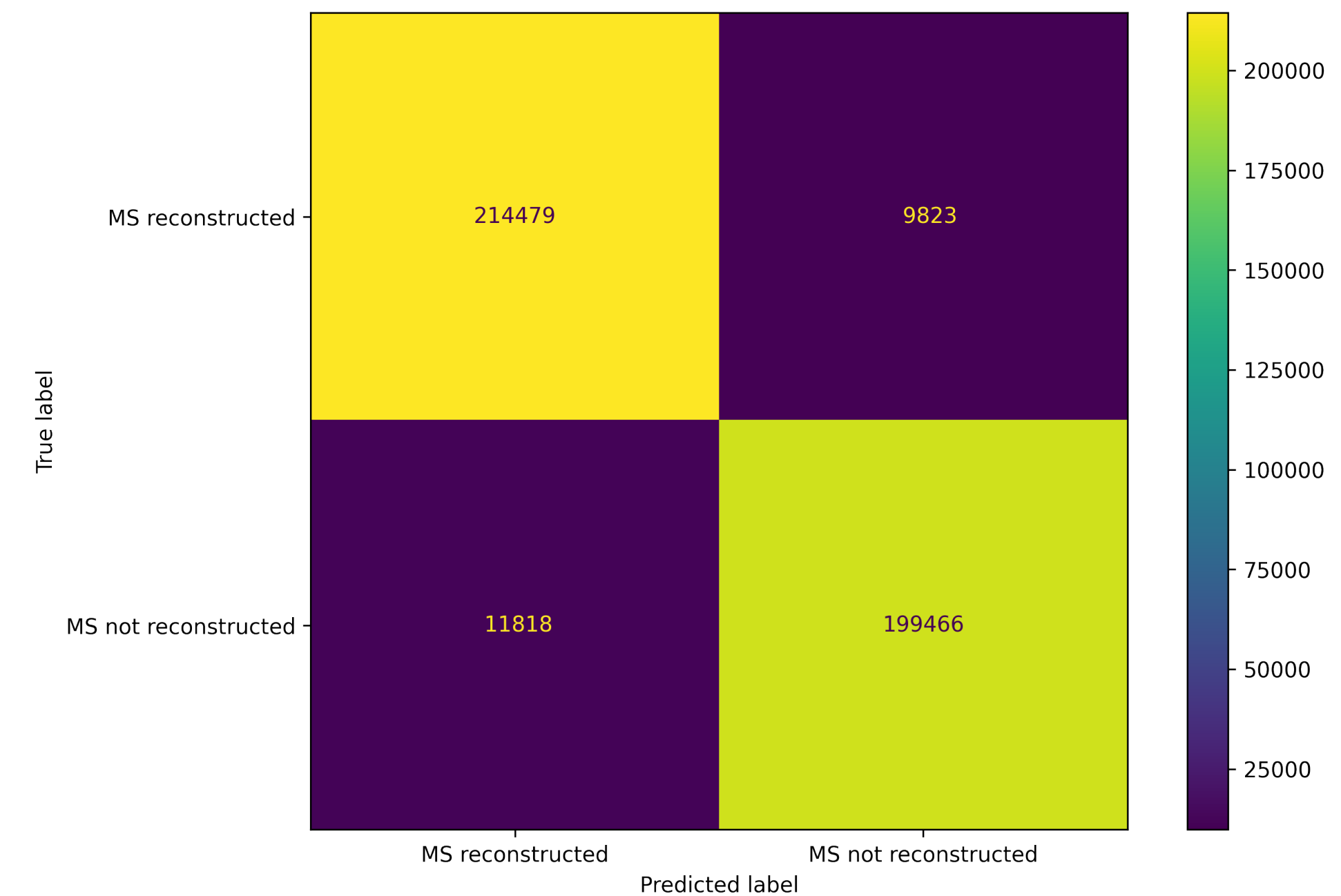
- We identified 4 main parameters that influence S2-separation:  $z$ ,  $\Delta z$ ,  $N_e(S2_I)$ ,  $N_e(S2_{II})$
- Since the functional correlation between those is not trivial, we plan on using Machine Learning techniques
- Classification Decision Tree: supervised learning approach to predict outcome based on set of inputs
- Use simulated events to train tree



# Machine learning

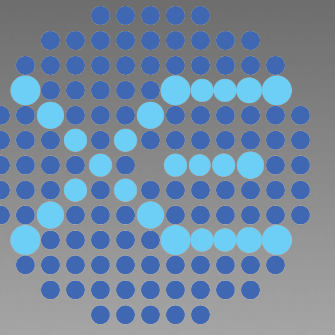


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- Classification Decision Tree: supervised learning approach to predict outcome based on set of inputs
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Already 95% accurate!

# Conclusions



- We are currently building a fast, effective simulator for XENONnT, that will be used in the analysis of many science channels of the experiment (e.g. double-beta decay, nucleon disappearance, ...)
- Machine learning techniques prove to be very fast and accurate in predicting multi-scatter resolution given a small set of inputs
- Ongoing work focuses on improving the decision tree further

## Further information

**XENON**

[xenonnt.org](http://xenonnt.org)

 [xenon\\_experiment](https://www.instagram.com/xenon_experiment)

 [@XENONexperiment](https://twitter.com/XENONexperiment)

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