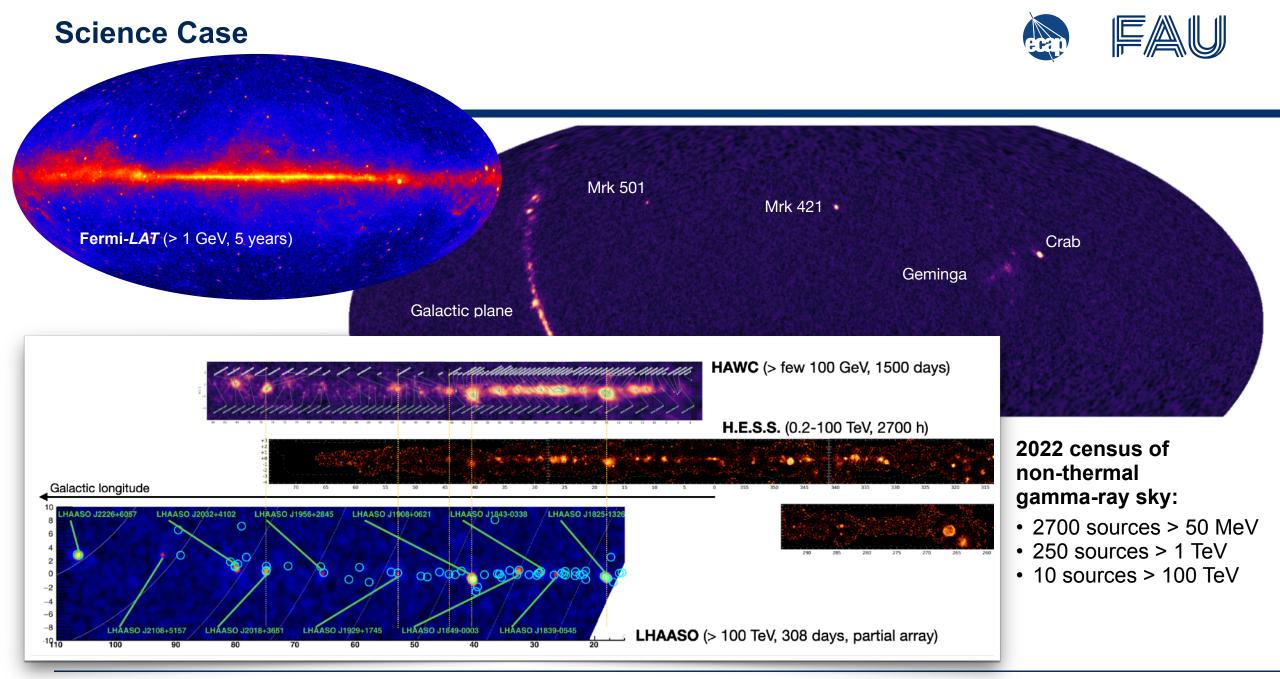


Gamma-ray Astronomy @ FAU

Christopher van Eldik • Stefan Funk • Alison Mitchell • ECAP April 4, 2022

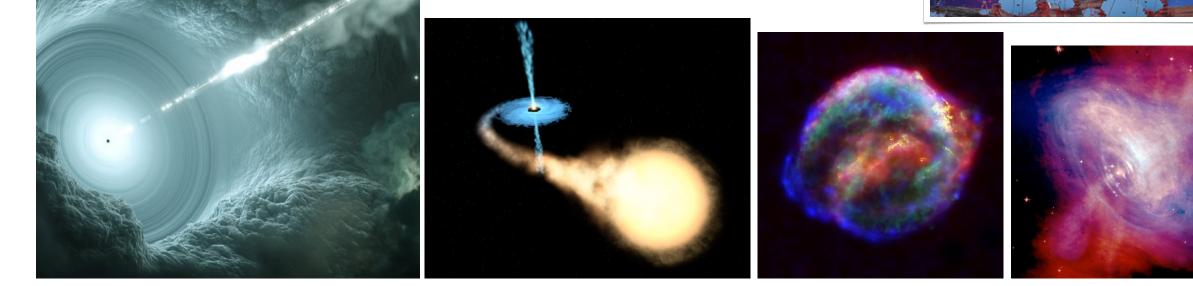


Gamma-ray Astronomy @ FAU

Science Case

(Astrophysical) key science questions:

- Where are charged particles accelerated to ultrarelativistic energies?
- How do these sources function?
- What are the acceleration processes at play?
- How does particle transport into their environments work?
- How do particles feed back on their environment?
- What is their contribution to the cosmic ray population?
- What is the nature of dark matter?



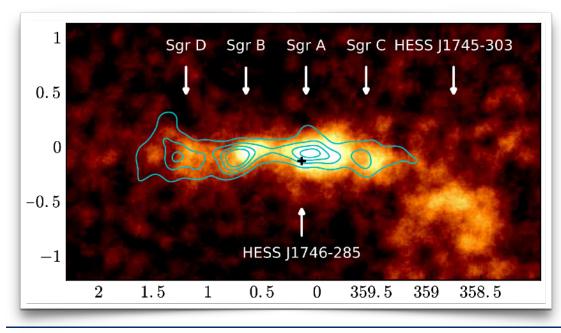


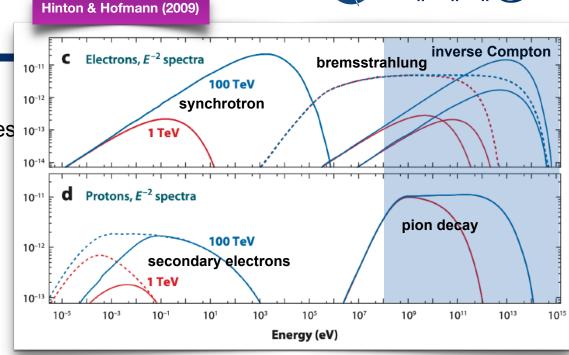
Science Case



(Astrophysical) key science questions:

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- How do these sources function?
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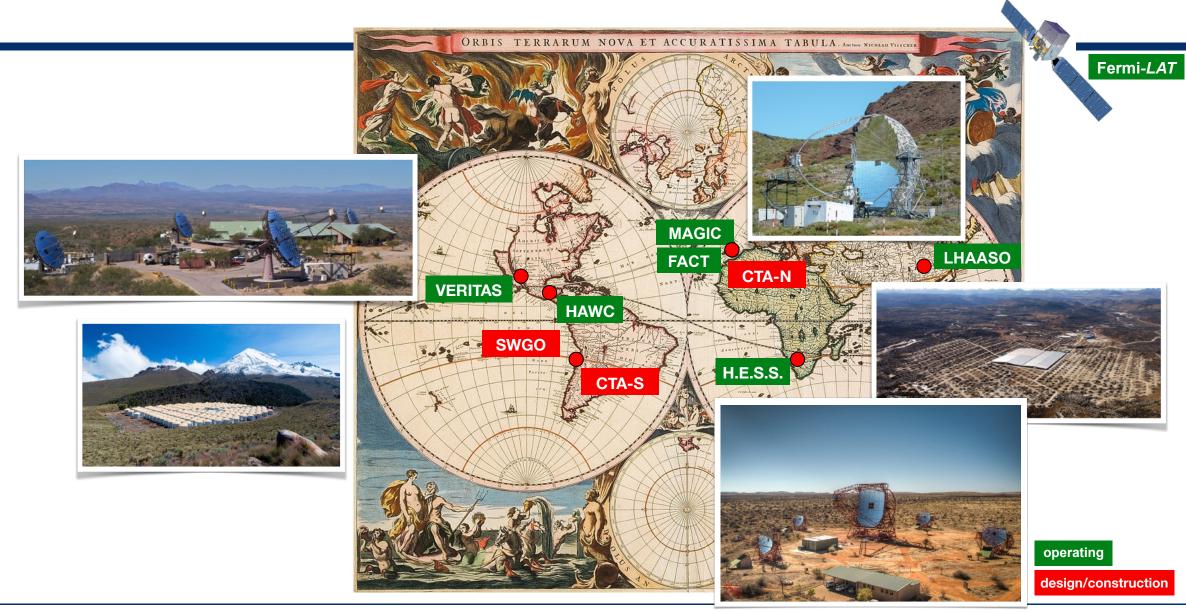
gamma rays enable

- access to non-thermal electrons (complementary to e.g. X-rays)
- unique access to non-thermal proton/ion populations

MWL/multi-messenger coverage often key:

- identification of dominant particle population
- understanding source physics through broad-band coverage
- \rightarrow contemporaneous multi-instrument campaigns
- \rightarrow ToO follow-ups

Instruments for Gamma-ray Astronomy



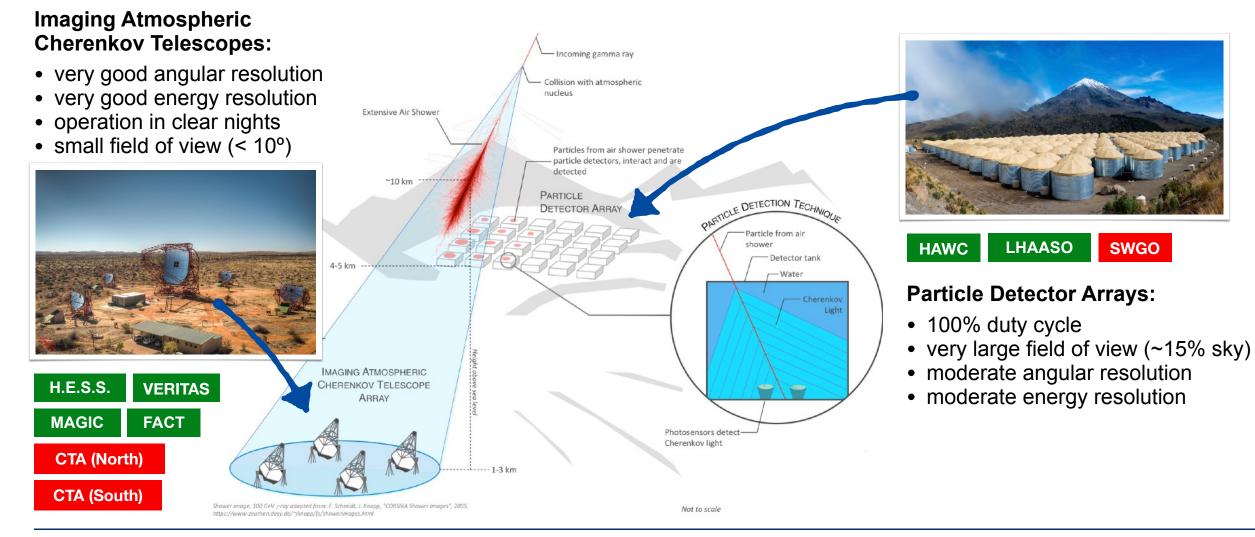
Instruments for Gamma-ray Astronomy



Instrument Complementarity

s Fau

Detection technique



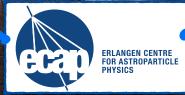
Cherenkov Telescope Array

MST



cherenkov telescope array

ARE



SST

LST

Cherenkov cameras - SSTCam



ECAP contributions to CTA



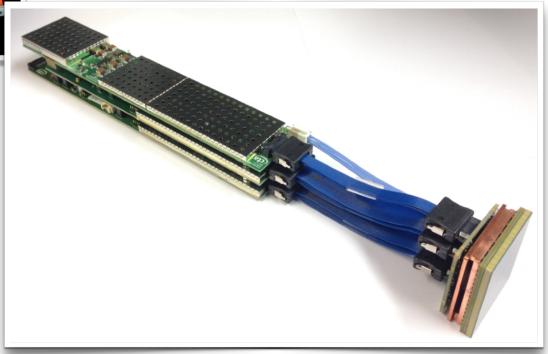


SSTCam: Compact camera using SiPMs

- 9° field of view, 2048 pixels, 1 GHz sampling rate
- 37-46 cameras for CTA-S array

ECAP:

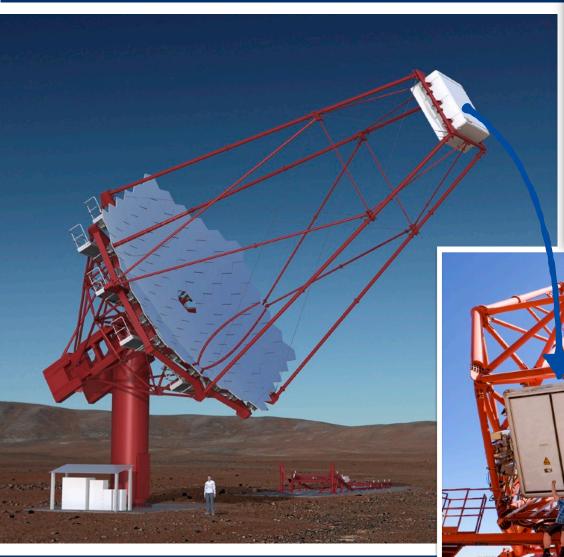
- development of read-out electronics
- signal characterisation and calibration

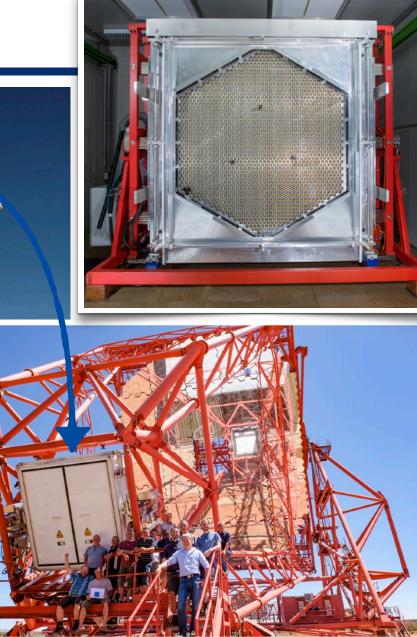


9

Cherenkov cameras - FlashCam

ECAP contributions to CTA









MST camera for CTA-S

- 1764 pixel (7.7° FoV)
- 250 MHz sampling rate
- fully digital trigger/readout

ECAP:

- tests/characterisation of photon detector planes
- camera module assembly
- calibration/science verification of FlashCam prototype in H.E.S.S.

Cherenkov cameras - FlashCam

ECAP contributions to CTA

Science

REPORTS

Cite as: H.E.S.S. Collaboration et al., Science 10.1126/science.abn0567 (2022).

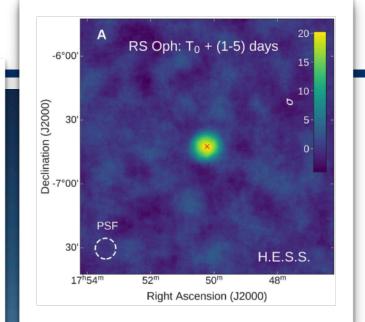
Time-resolved hadronic particle acceleration in the recurrent nova RS Ophiuchi

H.E.S.S. Collaboration*†

*Corresponding authors: contact.hess@hess-experiment.eu; Alison Mitchell (alison.mw.mitchell@fau.de); Stefan Ohm (stefan.ohm@desy.de); Brian Reville (brian.reville@mpi-hd.mpg.de)

†H.E.S.S. Collaboration authors and affiliations are listed in the supplementary materials.

Recurrent novae are repeating thermonuclear explosions in the outer layers of white dwarfs, due to the accretion of fresh material from a binary companion. The shock generated when ejected material slams into the companion star's wind can accelerate particles. We report very-high-energy (VHE, ≥100 GeV) gamma rays from the recurrent nova RS Ophiuchi, up to a month after its 2021 outburst, observed using the High Energy Stereoscopic System. The VHE emission has a similar temporal profile to lower-energy GeV emission, indicating a common origin, with a two-day delay in peak flux. These observations constrain models of time-dependent particle energization, favoring a hadronic emission scenario over the leptonic alternative. Shocks in dense winds provide favorable environments for efficient acceleration of cosmic-rays to very high energies.

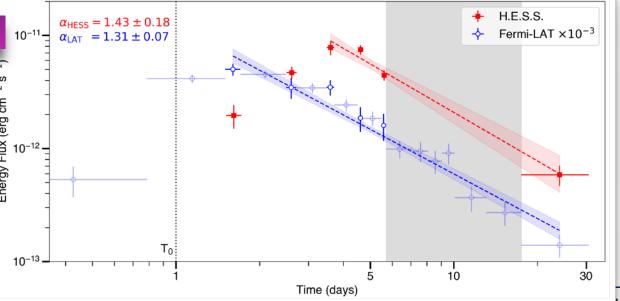




2021 outburst of recurrent nova RS Ophiuchi

- follow-up of optical outburst with H.E.S.S.
- particle acceleration to multi-TeV energies within days
- plausibly proton/ion acceleration dominates over electrons

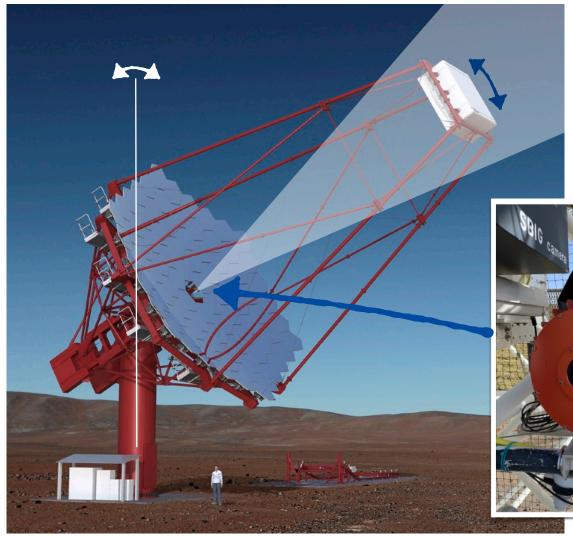




Telescope precision pointing

s fau

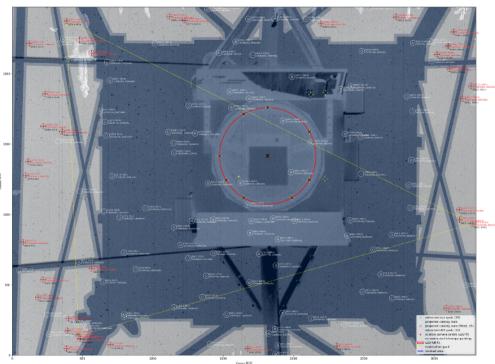
ECAP contributions to CTA



Calibration of telescope pointing to 5" precision

- camera hardware development
- camera mass production + tests
- steering software
- calibration algorithms
- observatory integration

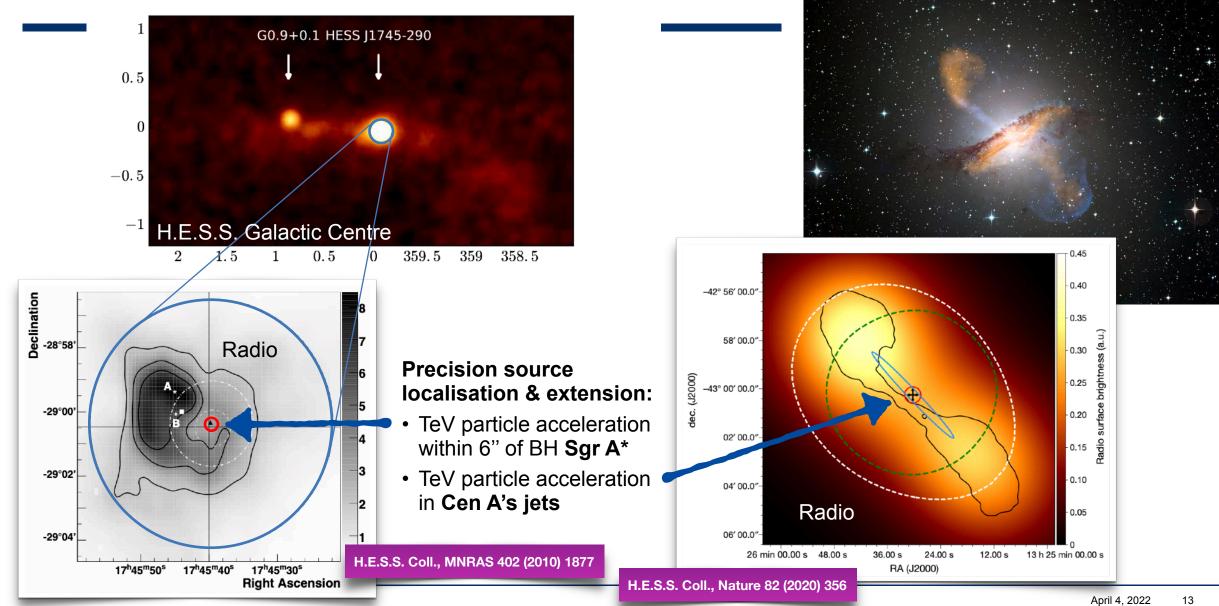




Telescope precision pointing

ECAP contributions to CTA

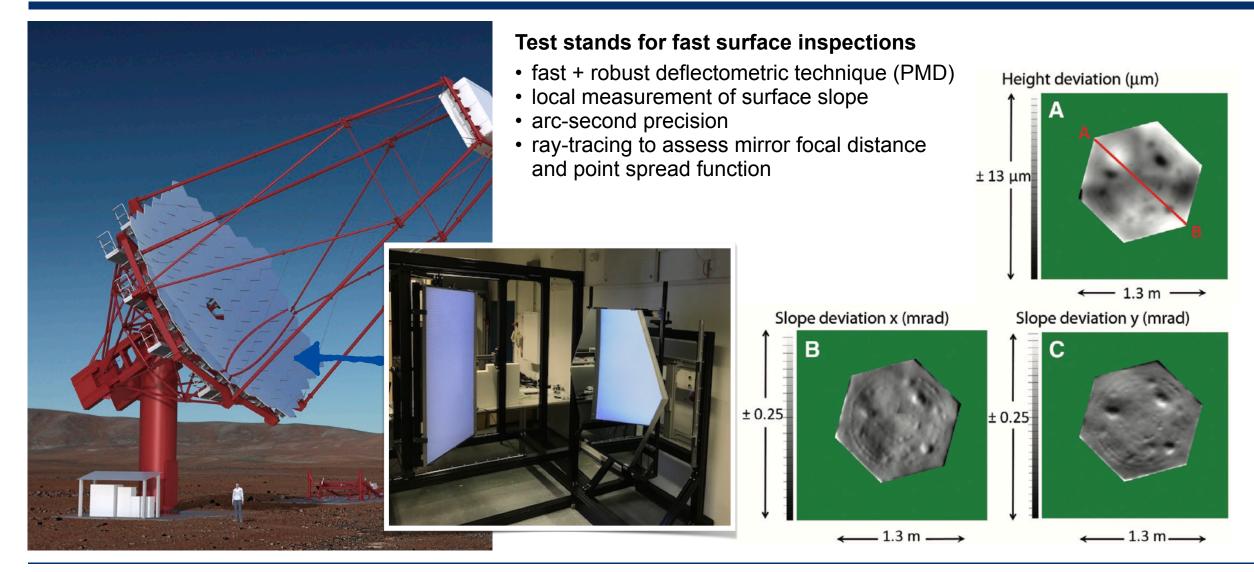




Mirror surface measurements

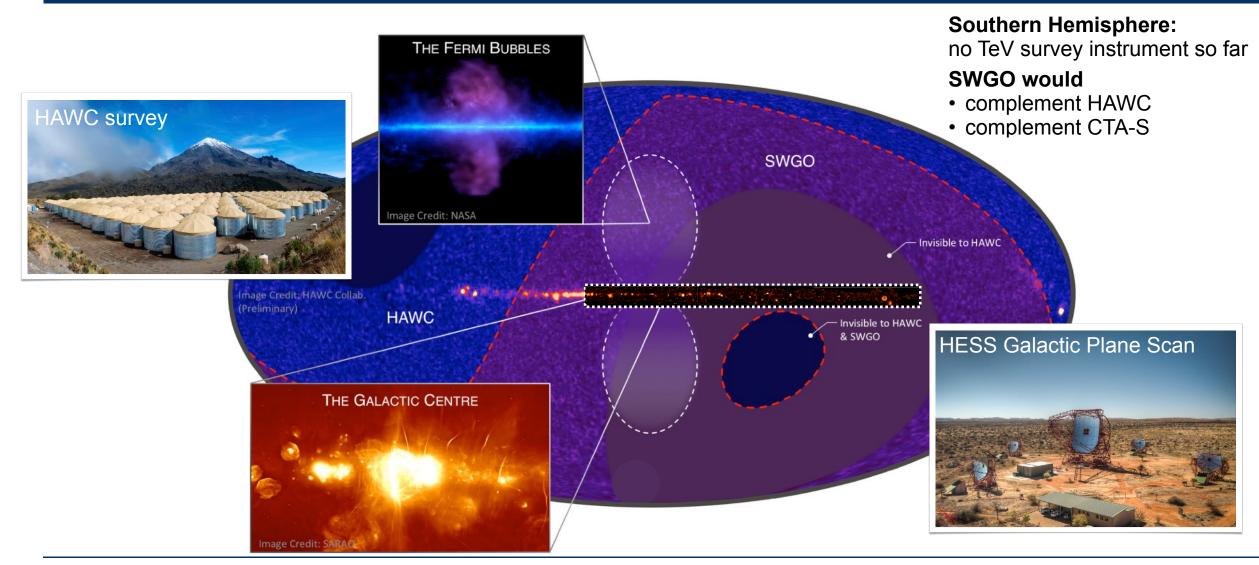
ECAP contributions to CTA





The Southern Wide-field Gamma-ray Observatory

A gamma-ray survey instrument in the southern hemisphere



The Southern Wide-field Gamma-ray Observatory



A gamma-ray survey instrument in the southern hemisphere



Reference configuration:

- 6600 water Cherenkov detectors
- 240.000 m² array footprint
- 4.700 m altitude

The Southern Wide-field Gamma-ray Observatory

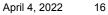
dense core array

sparse outer array Alto Tocomar (Argentina)
Cerro Vecar (Argentina)
Chacaltaya (Bolivia)
AAP Pajonal (Chile)
AAP Pampa La Bola (Chile)
Lake Sibinacocha (Peru)
Imata (Peru)
Yanque (Peru)

Site shortlisting in progress

Hinton et al. (SWGO Coll.) ICRC 2021

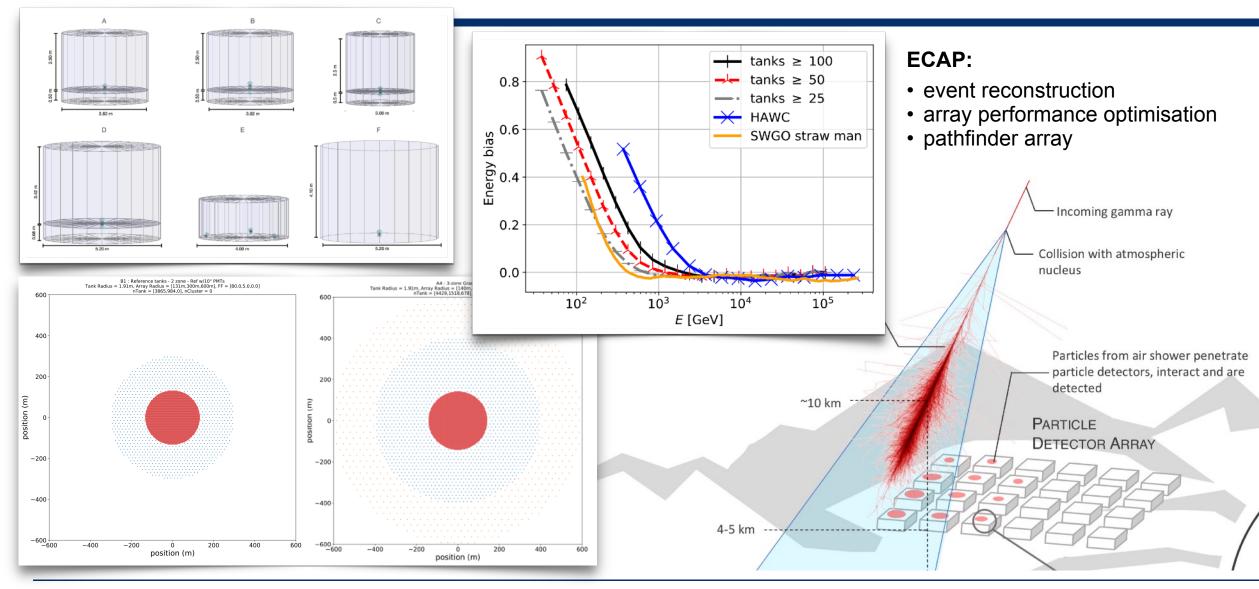
Gamma-ray Astronomy @ FAU



Reconstruction and Array Optimisation



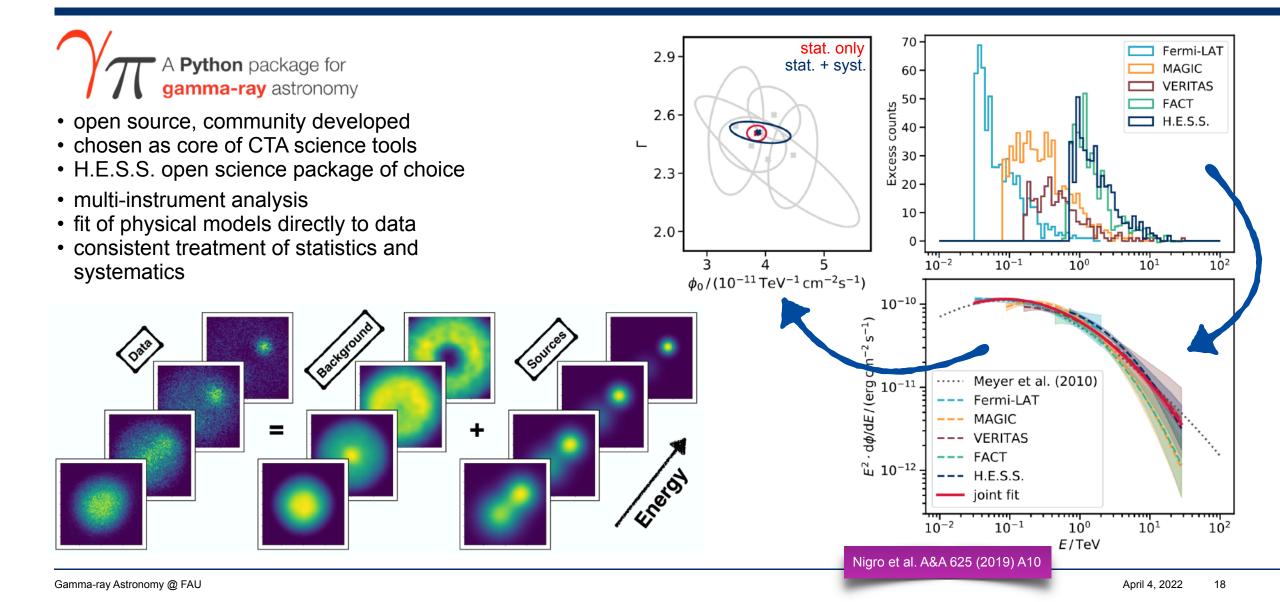
ECAP contributions to SWGO



Gammapy: Analysis Software for Gamma-ray Astronomy



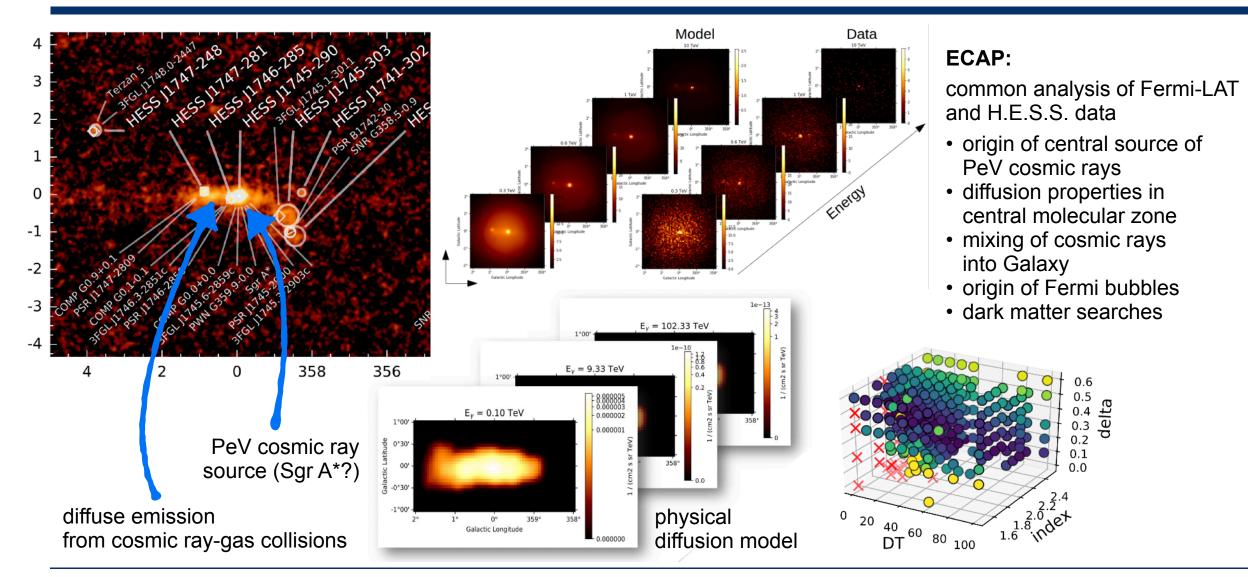
Towards joint instrument analysis



Analyses of complex regions

Galactic Centre with H.E.S.S.





Gamma-ray Astronomy @ FAU

ECAP

- strong involvement in leading running and up-coming gamma-ray observatories
- contributions to instrument building, calibration, reconstruction, analysis
- increasing focus on joint multiwavelength/multi-messenger analysis

Thanks for your attention.