

FRanconian Astronomy Neighborhood Collaboration Incentive: FRANCI 2021

RNO-G: Radio Neutrino Observatory in Greenland



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RNO-G Group: DESY / FAU Erlangen-Nürnberg



What do we do? Database Development; Software Development; Simulations; Energy/Direction Reconstruction; Sensitivity studies; Laboratory assembly and tests











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Ultra High Energies Neutrinos

lceCube / NASA

Astrophysical neutrinos: from Astrophysical sources (TDEs, GRBs, Blazars, NS-NS mergers, ...)

 $p + \gamma \rightarrow \Delta^{++} \rightarrow \pi^+ + n$ $\pi^{+} \rightarrow \mu^{+} + \nu_{\mu}$ $_{3} \mu^{+} \rightarrow e^{+} + \overline{\nu}_{\mu} + \nu_{e}$



UHE neutrinos probe the extragalactic sky and UHECR sources



Cosmogenic neutrinos: from interaction between **UHECR** and **CMB**





RNO-G addresses UHE neutrinos (> 10 PeV)

Radio detectors' scientific purpose is to observe ultra-high energy neutrinos, a regime out of reach via traditional optical methods.

At UHE the flux decreases at best as E^{-2}

We need a larger detector to improve the sensitivity



Towards the IceCube Gen2-Radio Array...

RNO-G is designed for **scalability**: Large-scale production and deployment



J. Phys. G: Nucl. Part. Phys. 48 (2021) 060501



RNO-G Sky coverage

 First neutrino detector of the Northern sky at UHE energies above 10 PeV

- At UHE is sensible to down going neutrinos and Earth-skimming neutrinos
- In the field of view there are interesting targets, as the sources seen by IceCube as the most significant sources in a point-source search.



The diurnally-averaged total field of view of the detector

Askaryan radiation

• Askaryan emission is caused by showers developing in a (dense) medium.

• The radiation is stronger at the Čerenkov angle but gets refracted in the ice.

• The attenuation length of radio waves in ice (order of kms) allows to detect higher energies.





RNO-G Planned Layout



RNO-G station arXiv: 2010.12279

Hybrid design combines advantages of **ARA** (deep) & **ARIANNA** (shallow):

Shallow component (~3 m)

- LPDAs facing upward and downward
- Cosmic rays and RFI veto

Deep component (~100 m)

- Hpols & Vpols
- Provides large fraction of effective volume
- Event reconstruction
- Includes a low-threshold trigger system



Work at ECAP Assembly and testing of the Amplifiers

Surface board









DRAB

IGLU

(Downhole Receiver and Amplifier Board)











Work at ECAP Assembly and testing of the Amplifiers: Spring 2021



Assembly





Function test





Stress test



Deployment From June to August (2021)













Events from the deep component: <u>Snowmobile</u>



First UHE radio neutrino observatory in the Northern hemisphere

- Designed for scalability
- 35 station planned, 3 deployed
- Three more field seasons planned (last deployment in 2024)
- Data are coming from the first stations...



