# Preserving Gauge Symmetries in Cosmology via Loop Quantum Gravity

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### Fiber bundle structure of Ashtekar variables A Yang-Mills approach

The natural language of Yang-Mills theories is the principal bundle formalism.

Principal G-bundle  $(P, \pi, M)$ 

- *P* a smooth manifold called «fiber bundle»,
- *M* a smooth manifold called «base manifold»,
- $\pi: P \to M$  a smooth projection,

#### Such that:

- G has a free and transitive action on  $\pi^{-1}(x)$ ,  $x \in M$ . Hence,  $\pi^{-1}(x) \cong G$ ,
- $P/G \cong M$ .

Gauge field encoded in a connection

 $\omega \in \Omega^1(P, g)$ 

Such that:

- $\omega(X^A) = A, A \in \mathfrak{g}$
- $R_a^* \omega = \operatorname{Ad}_{a^{-1}} \omega, \ a \in G$

#### Fiber bundle structure of Ashtekar variables Orthonormal frame bundle

A principal bundle appears from the tetrad formulation: orthonormal frame bundle  $P^{SO}(M)$ 

$$g \text{ metric on } M \iff P^{SO}(M)$$

 $P_x^{SO}(M) = \{h: \mathbb{R}^n \to T_x M \mid h \text{ oriented isometry}\} = \{\text{collection of orthonormal basis of } T_x M \}$ 

A choice of the tetrad is the choice of a section  $e: M \to P^{SO}(M)$ 

A Lorentz connection  $\omega$  is a connection on  $P^{SO}(M)$  (metric-compatible)

**E.g.** in n=4 the tetrad  $e_{\alpha}^{\mu}$  is  $h(\varepsilon_{\alpha}) \doteq e_{\alpha} = e_{\alpha}^{\mu}\partial_{\mu}$  and so  $g(e_{\alpha}, e_{\beta}) = \langle \varepsilon_{\alpha}, \varepsilon_{\beta} \rangle = \eta_{\alpha\beta}$  $\{\varepsilon_{1}, \varepsilon_{2}, \varepsilon_{3}, \varepsilon_{4}\}$  canonical basis of  $\mathbb{R}^{4}$  M

# Preserving Gauge Symmetries in Cosmology via LQG

#### Fiber bundle structure of Ashtekar variables The dimension 3 - Dreibein

**3-dimensional Riemannian** manifold  $(\Sigma, q)$ 

 $P^{SO}(\Sigma)$  is a principal SO(3)-bundle  $\Rightarrow$  Gauge group SO(3)

A choice of a triad  $e_i^a(x)$  is equivalent to the choice of a section  $e: \Sigma \to P^{SO}(\Sigma)$ , i.e.  $\pi(e_x) = x$ 

 $e_x(\varepsilon_i) = e_i^a(x)\partial_a$ 

 $P^{SO}(\Sigma)$ 

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3

#### **Fiber bundle structure of Ashtekar variables The dimension** 3 – SU(2) appears

**3-dimensional Riemannian manifold**  $(\Sigma, q)$ 

Spin Structure

 $P^{Spin}(\Sigma)$ 

 $P^{SO}(\Sigma)$ 

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• Principal SU(2)-bundle  $P^{Spin}(\Sigma)$ 

• **Double-covering**  $\rho: P^{Spin}(\Sigma) \to P^{SO}(\Sigma)$ 

Lift of the dreibein  $\overline{e}$  s.t.  $e = \rho \circ \overline{e}$  (It is not unique, but it does not matter)

$$P^{SO}(\Sigma) \\\downarrow \\\Sigma$$





#### Fiber bundle structure of Ashtekar variables The dimension 3 – Connection

**3-dimensional spin manifold** 
$$(\Sigma, q)$$
  
 $\downarrow$   
 $P^{SO}(\Sigma)$   
 $\downarrow$   
 $\Sigma$ 

A spin connection is represented by a 1-form  $\omega$  with values in  $\mathfrak{su}(2)$  on  $P^{Spin}(\Sigma)$ 

A metric-compatible connection is represented by a 1-form  $\varpi$  with values in  $\mathfrak{so}(3)$  on  $P^{SO}(\Sigma)$ 

Ashtekar connection: Local field of a spin connection  $A = \bar{e}^* \omega$  is a 1-form with values in  $\mathfrak{su}(2)$  on  $\Sigma$ 

### Fiber bundle structure of Ashtekar variables The dimension 3 – Electric field



The space of Ashtekar connection  $\mathcal{A}$  is independent on the choice of spin structure and metric. Canonically it is the space of SU(2) connections

Collect the metric data in the electric field E

The electric field E can be interpreted as a 2-form on  $\Sigma$  with values in  $ad^*P^{Spin}(\Sigma)$ . Described locally by:

 $E = \star e_a^i dx^a \tau_i$ 

#### Cosmological sector of GR in Ashtekar variables A Yang-Mills approach

[Brodbeck '96] [Bojowald, Kastrup '00] [MB '24] [MB '24]

 $P^{Spin}(\Sigma)$ Yang-Mills variables $\downarrow$ Connection  $\omega$  is a 1-form on  $P^{Spin}(\Sigma)$  with value in the Lie algebra of SU(2) $P^{SO}(\Sigma)$ Dreibein e is a section in  $P^{SO}(\Sigma)$  $\downarrow$ Ashtekar variables $\Sigma$ Connection A is the local field  $A = e^*\omega$ Electric field E is built from the dreibein  $E = * e_a^i dx^a \tau_i$ 

The problem is to find the cosmological sector of those variables

#### Cosmological sector of GR in Ashtekar variables Cosmological hypothesis

**Homogeneity:** a connected Lie group S acts transitively and freely on  $\Sigma \Rightarrow \Sigma \cong S$ 

Namely, for every two points x, y there exists a unique  $g \in S$  s.t.  $y = L_g x$ 

Homogeneous ADM variables (q, K): they are S-invariant tensors

$$L_g^* q = q$$
$$L_g^* K = K$$

Breaks diffeo-invariance  $Diff(\Sigma) \rightarrow S_R \rtimes Aut(S)$ 

#### Cosmological sector of GR in Ashtekar variables Cosmological hypothesis for gauge fields

 $P^{SO}(S)$  is homogeneous if it is S-invariant, i.e. there exists an action  $\phi$  of S s.t.  $\pi \circ \phi(g) = L_g$ 

There exists a unique(!) homogeneous spin structure on S

Homogeneity condition for connection from Wang's theorem  $\phi(g)^*\omega = \omega, \forall g \in S$ (classified by linear maps  $\Lambda: \mathfrak{s} \to \mathfrak{su}(2)$ )

Ashtekar connections as homogeneous spin connections

The request of homogeneity for Ashtekar connection yields a homogeneous geometry for  $\Sigma$ 

#### Cosmological sector of GR in Ashtekar variables Reduced Phase Space

**Configuration space**  $\mathcal{A}^{S} = \{A \mid A = e^{*}\omega, \omega \text{ homogeneous}\}$ 

Dreibein is a section in a homogeneous bundle, but it is not homogeneous itself

Phase space variables  $(A_a^i(x), E_i^a(x))$  such that, in some gauge  $(\phi_I^i \theta^I(x), p_i^I \xi_I(x))$   $\xi_I \in \mathfrak{s} \quad [\xi_I, \xi_J] = f_{IJ}^K \xi_K$  $\theta^I \in \mathfrak{s}^* \quad \theta^I + \frac{1}{2} f_{JK}^I \theta^J \theta^K$ 

The set of constraints are the same of LQG

### Cylindrical functions for Cosmology Gauge transformations

**Configuration space**  $\mathcal{A}^{S} = \{A \mid A = e^{*}\omega, \omega \text{ homogeneous}\}$ 

[Brunnemann, Fleischhack '12]

Cylindrical functions  $Cyl(\mathcal{A}^{S})$   $f_{\gamma}: \mathcal{A}^{S} \to SU(2)^{E(\gamma)} \to \mathbb{C}$ 

#### Infinitesimal transformations

Recover gauge symmetry 
$$\begin{array}{lll} \mathcal{G} \times \mathcal{A}^S & \to & \mathcal{A}^S \\ (u, e^* \omega) & \mapsto & e^* u^* \omega &= (u \circ e)^* \omega = e'^* \omega \end{array} \qquad A \mapsto A - D\Lambda$$

Residual diffeo symmetry
$$\operatorname{Aut}(S) \times \operatorname{Cyl}(\mathcal{A}^S) \rightarrow \operatorname{Cyl}(\mathcal{A}^S)$$
  
 $(\varphi, f_{\gamma}) \mapsto f_{\varphi(\gamma)}$  $A \mapsto A + \mathcal{L}_X A$ 

### Cylindrical functions for Cosmology Classical Constraints Algebra

**Configuration space**  $\mathcal{A}^{S} = \{A \mid A = e^{*}\omega, \omega \text{ homogeneous}\}$ 

Smearing functions must generate the correct gauge transformations

$$\in C_c^{\infty}(\Sigma, \mathfrak{su}(2))$$
  $X \in Lie(S_R \rtimes Aut(S)) \subset \mathfrak{X}(\Sigma)$   $N = const.$ 

 $\{G(\Lambda), G(\Lambda')\} \propto G([\Lambda, \Lambda'])$  $\{V(X), V(X')\} \propto V(\mathcal{L}_X X')$  $\{V(X), H(N)\} = 0$  $\{H(M), H(N)\} = 0$ 

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#### Cylindrical functions for Cosmology Integral curves and homogeneous graphs

An important property of curves on S:

every curve can be approximated by piecewise integral curves of invariant vector field  $\xi \in \mathfrak{s}$ 



Homogeneous graphs as suitable collections of piecewise integral curves

The homogeneous graphs are dense in the set of graphs

Distinguished subspace of cylindrical functions over homogeneous graph  $Cyl_{S}(\mathcal{A}^{S})$ 

# **Conclusions**

- Ashtekar-Barbero-Immirzi formulation has a rigourous and clear geometric interpretation, in which the data are encoded in a spin connection A and a section (gauge) e.
- In this formulation we can find a cosmological sector using the Wang's theorem preserving local gauge degrees of freedom.
- We are able to provide cylindrical functions, and a distinguished invariant subspace based on homogeneous graphs.
- We have constraints associated to gauge transformations, but with a simplified algebra.

#### Outlook

- Implement the algebra at the quantum level. Solving the Hamiltonian constraint restricting the ambiguities by comparison with LQC.
- Recovering the LQC dynamics in some suitable limit (single-vertex model,  $j \gg 1$ ).

# Thank you for your attention

#### **Preserving Gauge Symmetries in Cosmology via Loop Quantum Gravity**

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# Symmetry-reduced phase space

Configuration space  $\mathcal{E} = \{$ electric field gauge-equivalent to a homogeneous one $\}$ 

Fixing a point  $E \in \mathcal{E}$  determine a metric on  $\Sigma$ , and we can build  $\mathcal{A}_E^S \cong \mathcal{A}^S$ 

 $\mathcal{A}^S \hookrightarrow \mathcal{P}$ Phase space  $\mathcal{P}$  as fiber bundle  $\mathcal{P} = \bigsqcup_{E \in \mathcal{E}} \mathcal{A}_E^S$  $\mathcal{A}^S \hookrightarrow \mathcal{P}$  $\mathcal{E}$  $\mathcal{E}$ 

On this phase space, the contraints have the same functional forms of the constraints in usual LQG

# Point holonomies and invariant spin-network states

Equation of parallel transport along a homogeneous curve in a homogeneous gauge

 $\dot{u}(t) = \Lambda(v)u(t)$ 



**Holonomy**  $h_c(A) = u(1)^{-1} = \exp(\Lambda(v))$ 

The holonomies brought by invariant homogeneous spin-network states are point holonomies

# Topology of $\mathcal{A}/\mathcal{G}$

 $\mathcal{A}^{S}/\mathcal{G}$  spin (metric-compatible) homogeneous connections modulo gauge transformations

 $\mathcal{M}$  homogeneous SU(2) connections modulo gauge transformations

#### **Relevant Homeomorphisms**

	$\mathcal{A}^{S}/\mathcal{G}$	${\mathcal M}$
Bianchi universes ( $H = \{1\}$ )	$\mathbb{R}^{6}$	$\mathbb{R}^{6}$
Axial symmetry (H = U(1))	$\mathbb{R} \times \mathbb{R}_+$	$\bigsqcup_{n\in\mathbb{N}}\mathbb{R}$
FLRVV (H = SO(3))	$\mathbb{R}$	{0}