



Probing the physics of jet launching and propagation through numerical simulations

Ainara Saiz-Pérez

Christian M. Fromm, Felix Glaser, Raoul Kinadeter, Wladislaw Schulga, Eric Nanowsky, Florian Maul, Luca Ricci

Overview and Scales

Horizon Scale $< 1000r_g$

BH accretion & jet launching

Propagation Scale $> 10000r_g$

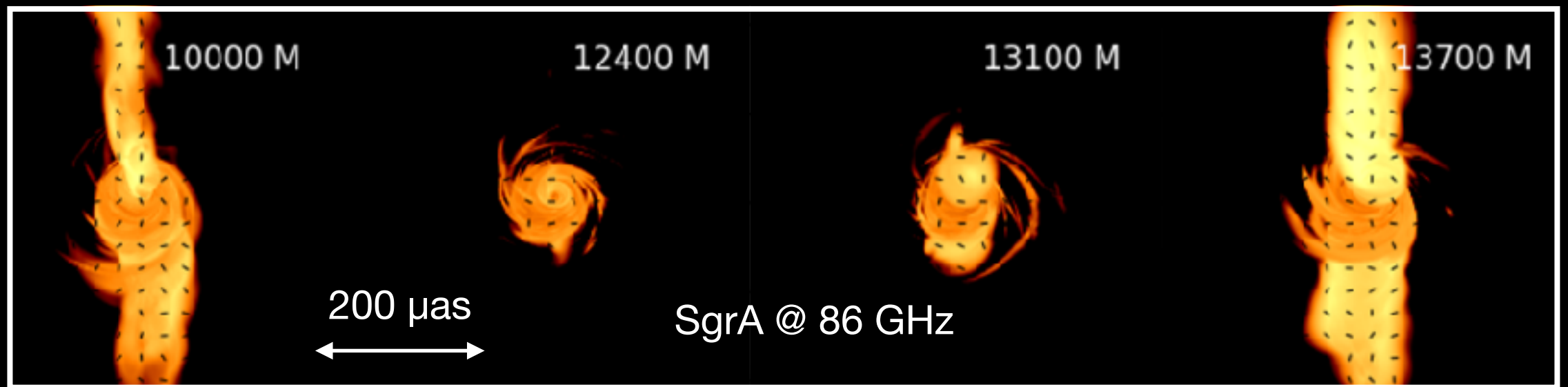
KH & RT Instabilities

Acceleration Scale $< 10000r_g$

Jet acceleration & CD Instabilities

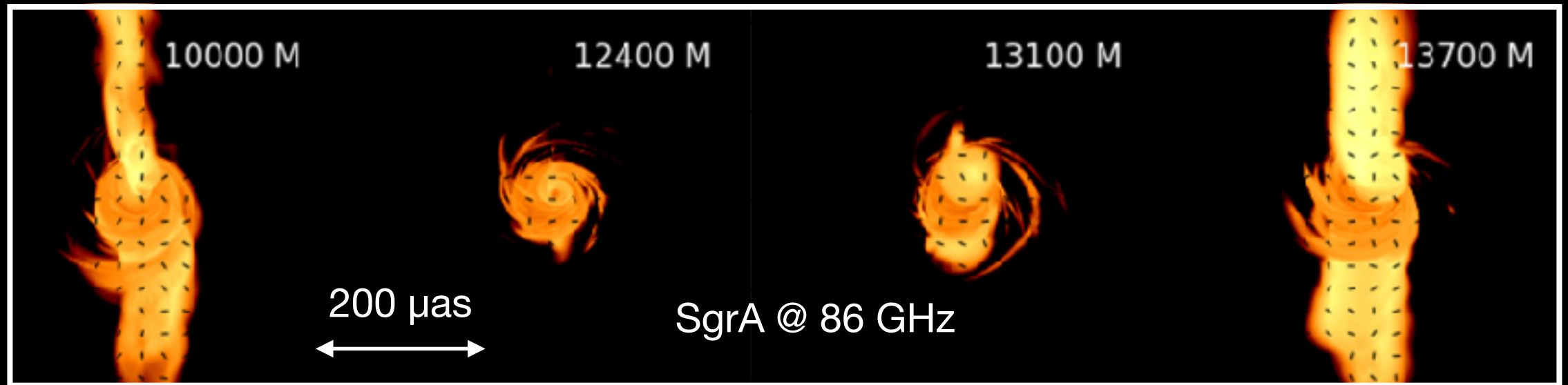
3D GRMHD & Rendering: C. M. Fromm

Black Hole Accretion and Flares



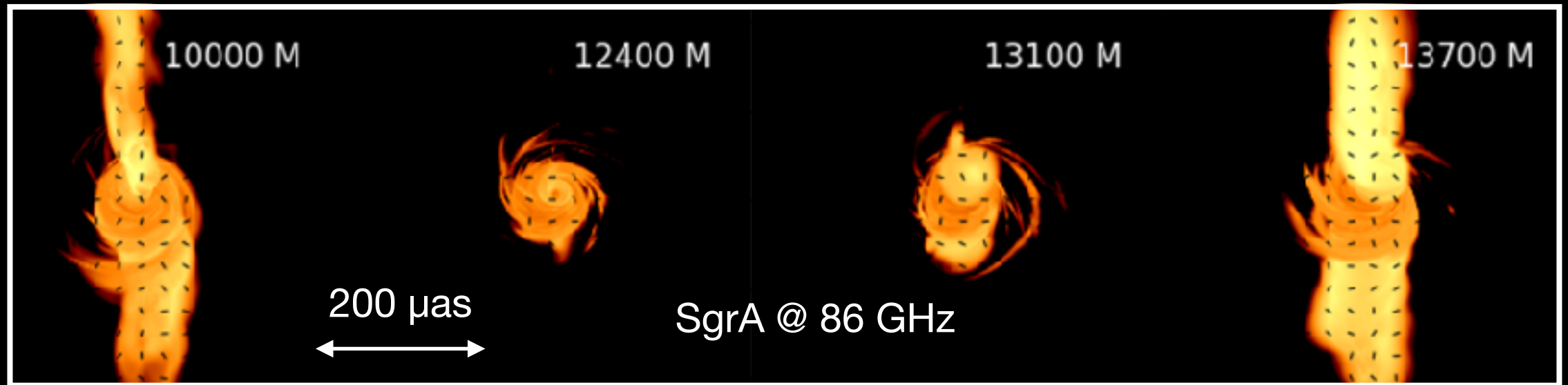
Black Hole Accretion and Flares

Transient jets during accretion of mag. loops with alternating polarity



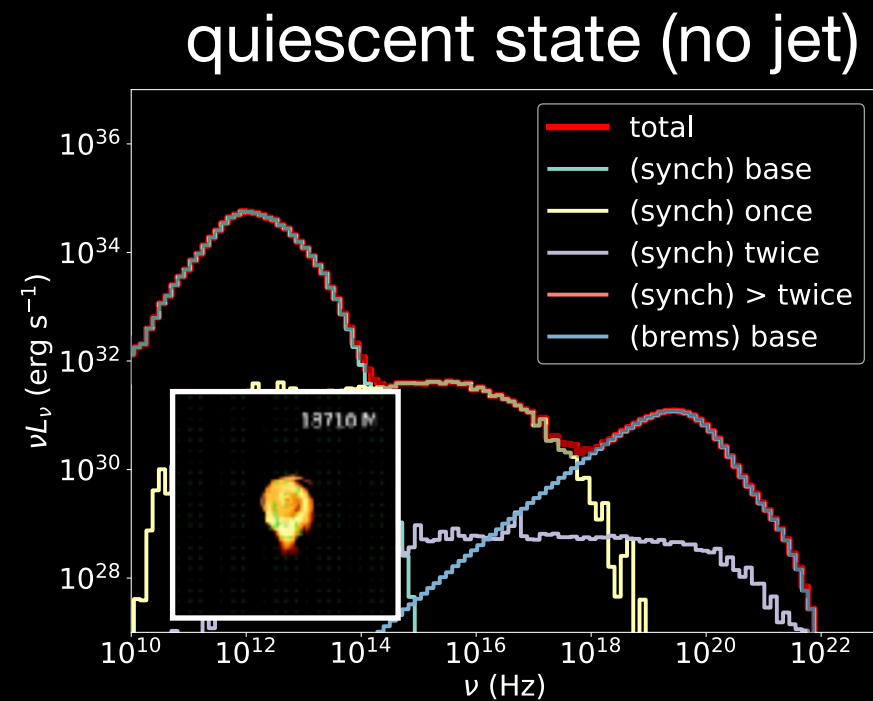
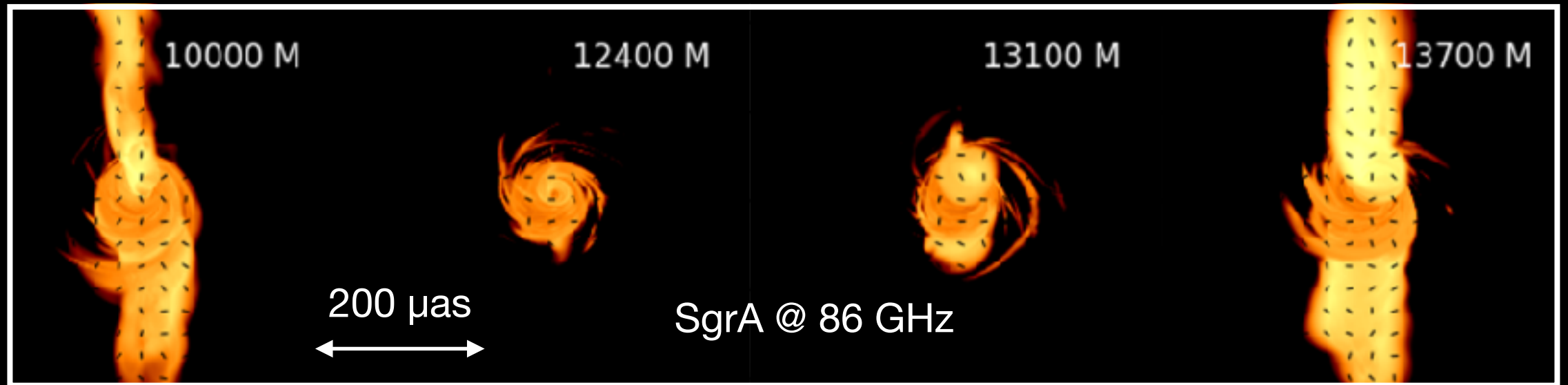
Black Hole Accretion and Flares

Transient jets during accretion of mag. loops with alternating polarity
—> modelling the flares of SgrA*



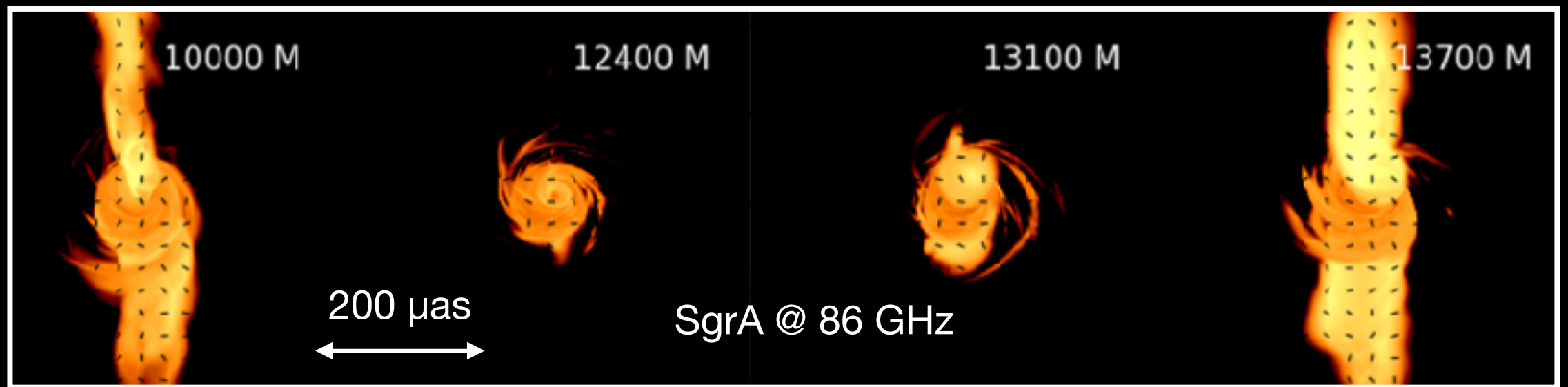
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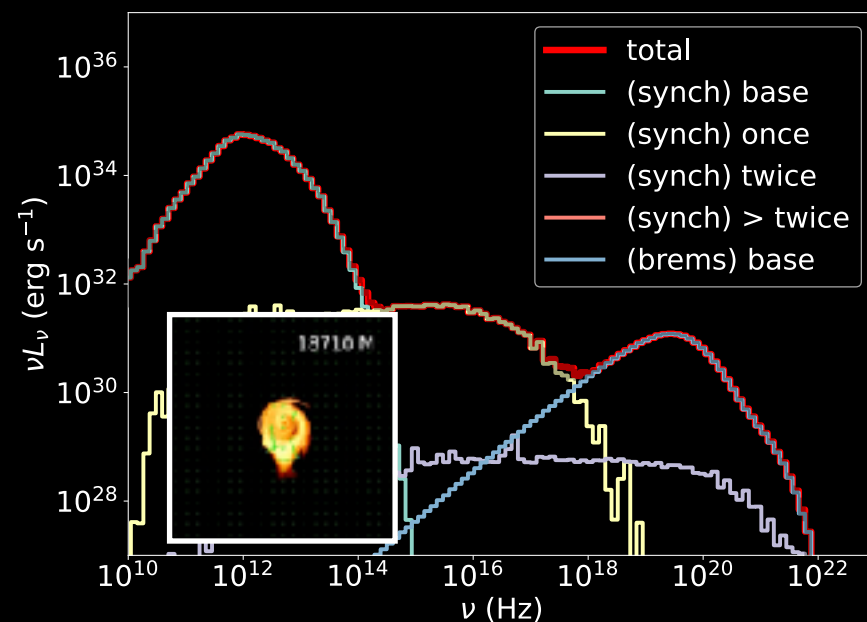


Black Hole Accretion and Flares

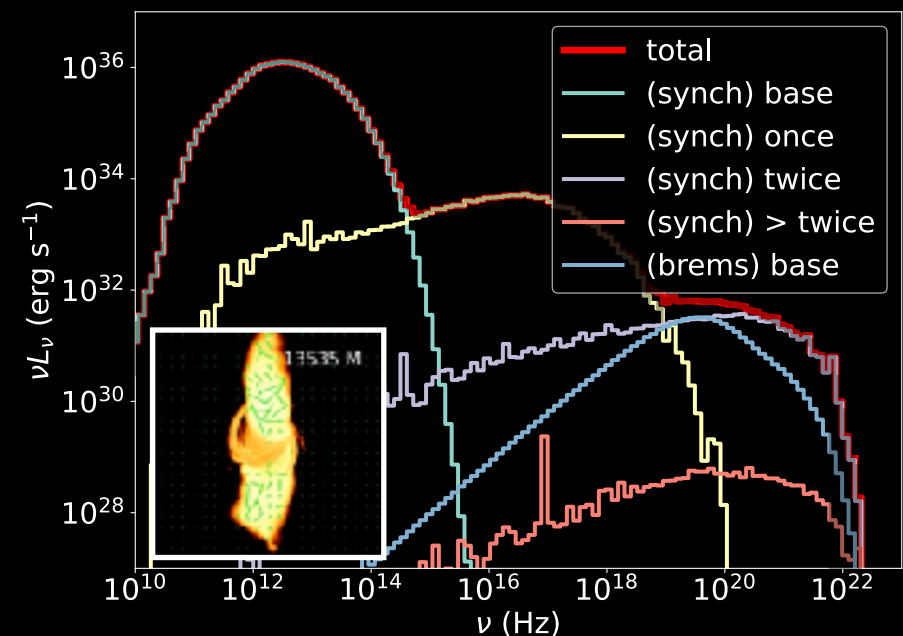
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quiescent state (no jet)

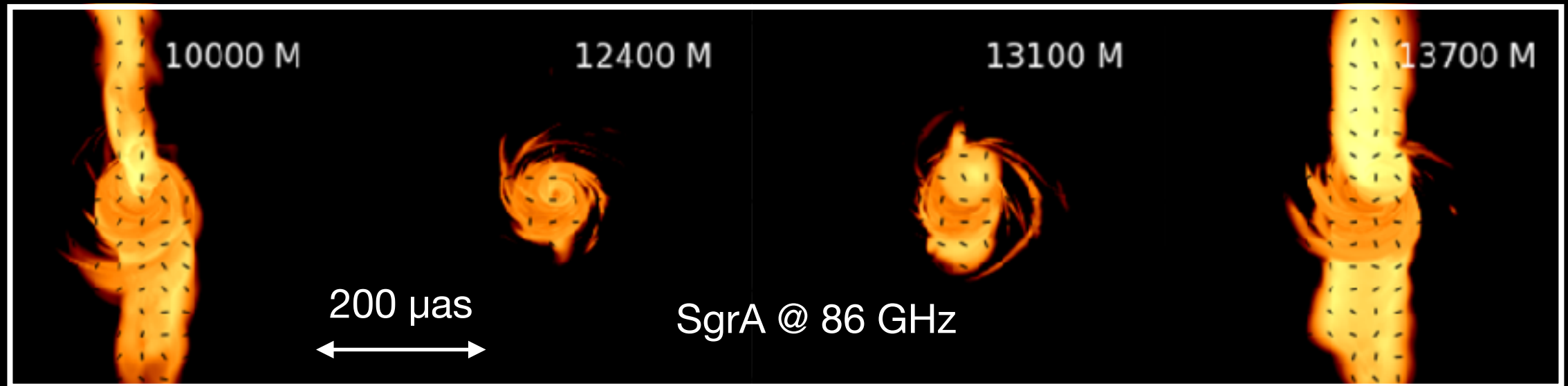


active state (strong jet)

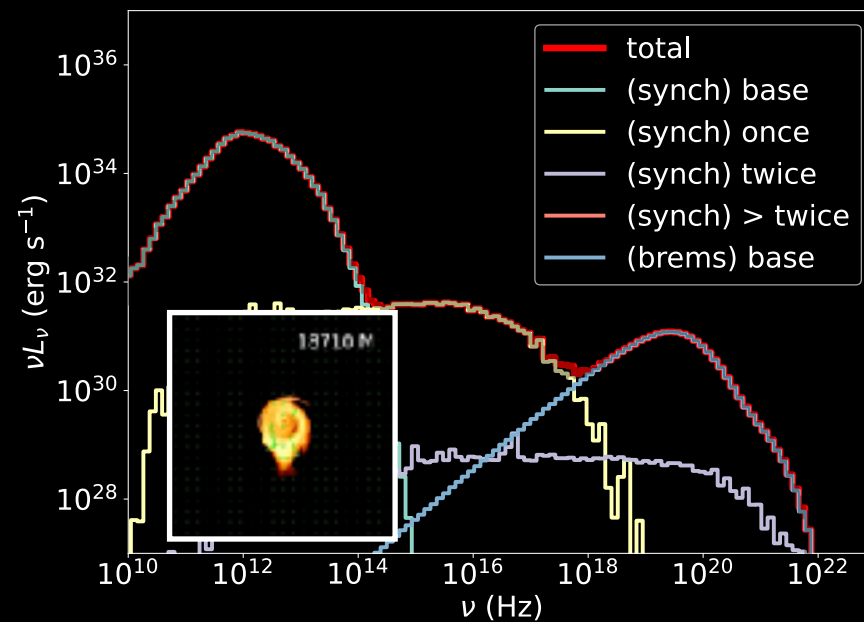


Black Hole Accretion and Flares

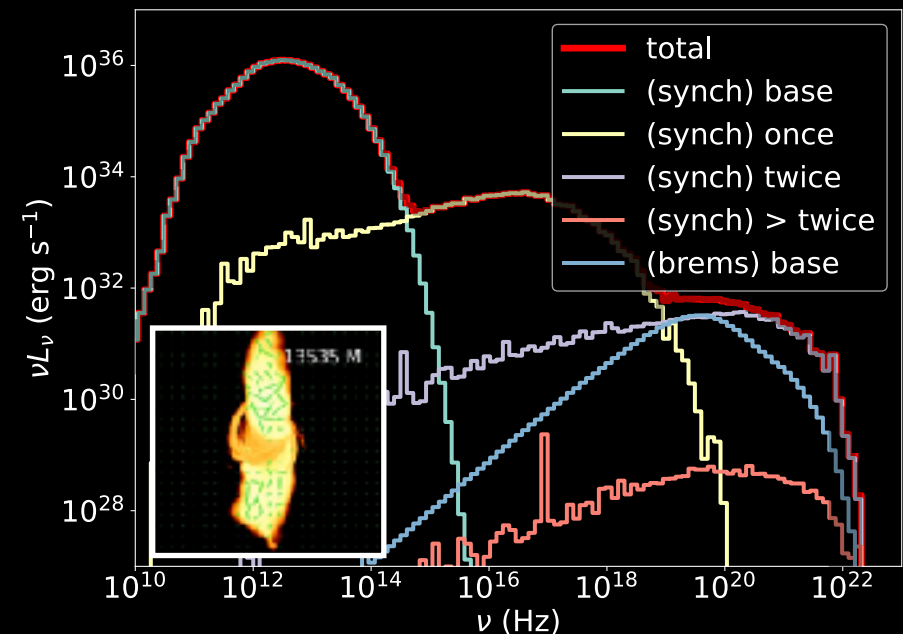
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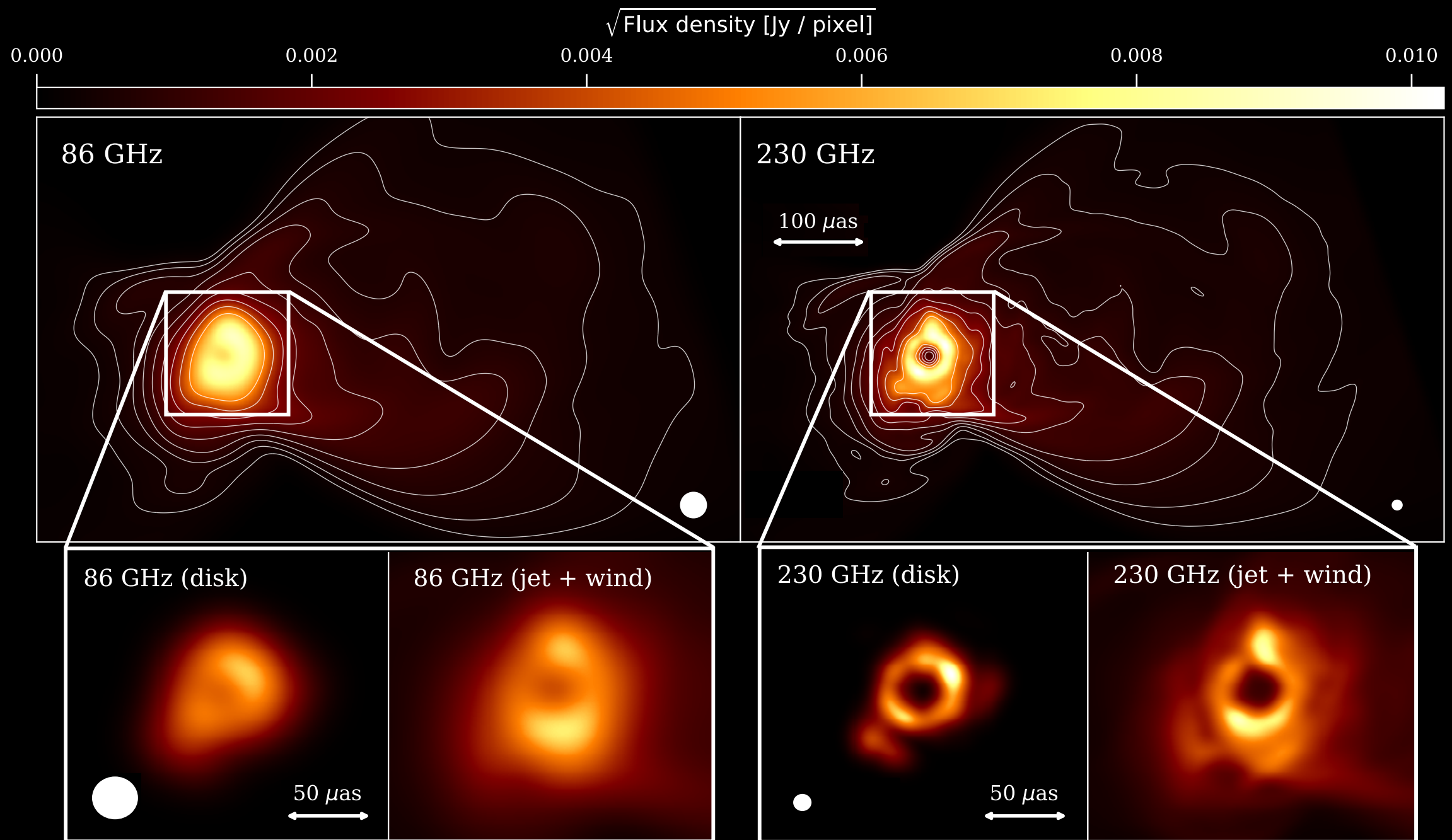
active state (strong jet)



—> jet formation/propagation connect to high-energy flares

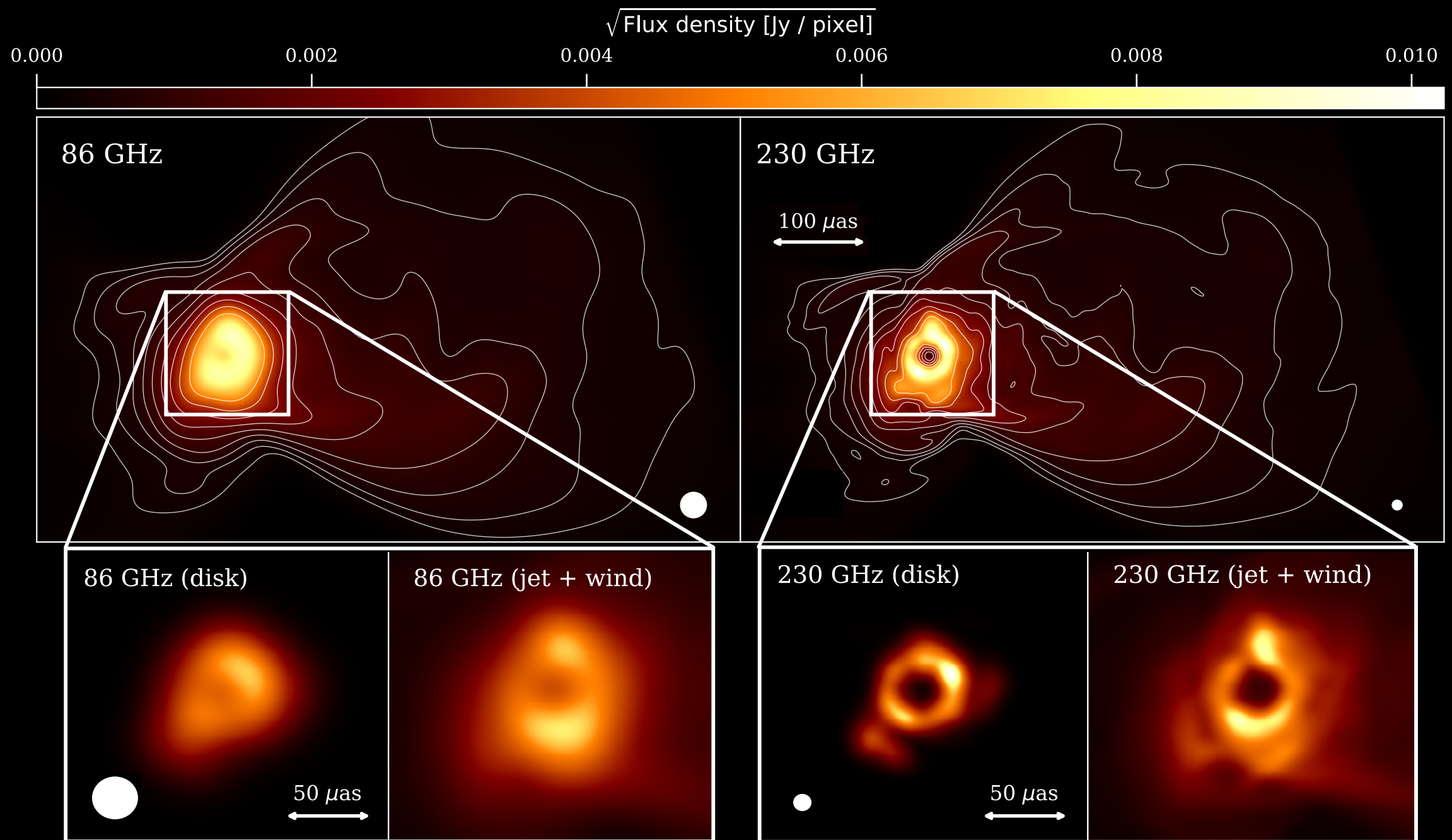
Jet-Disk Coupling

Modelling the observed ring size of M87 at different mm frequencies



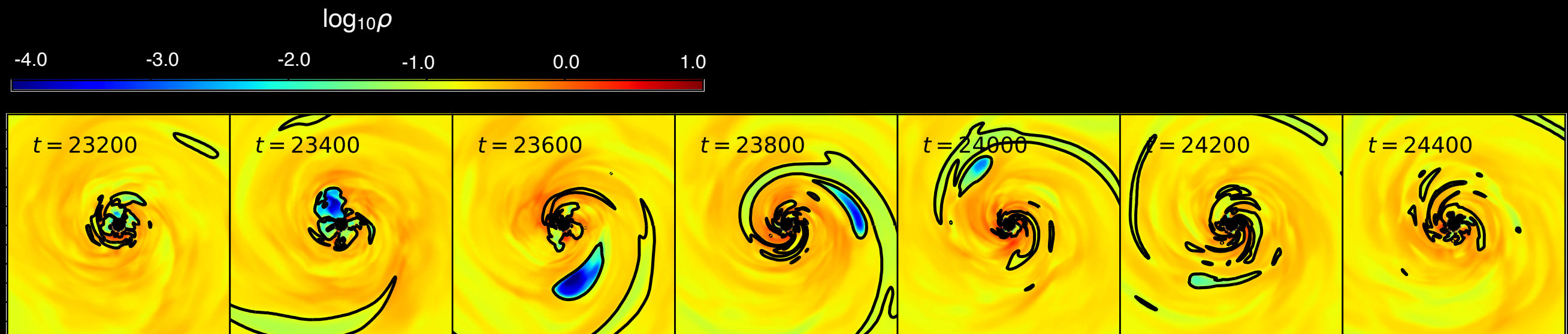
Jet-Disk Coupling

Modelling the observed ring size of M87 at different mm frequencies
—> role of non-thermal emission (electron acceleration)



Jet-Disk Coupling

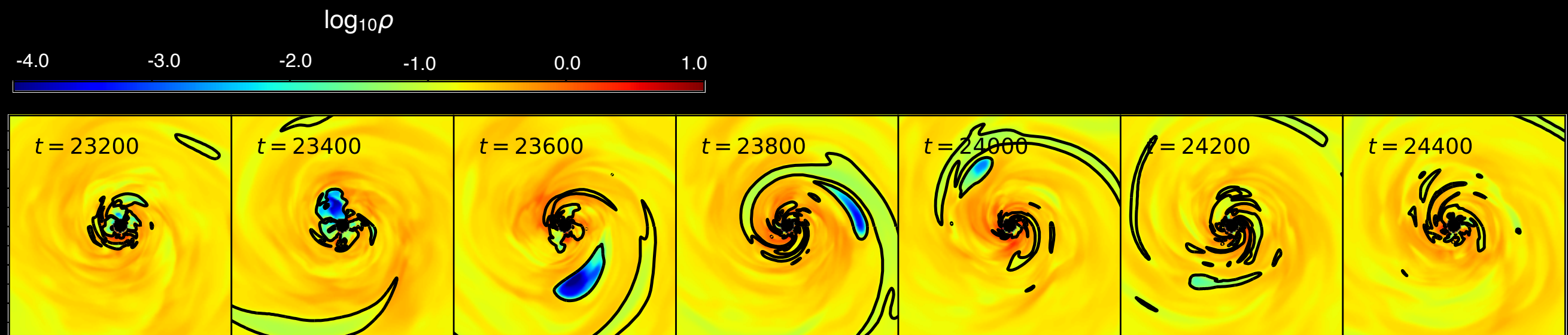
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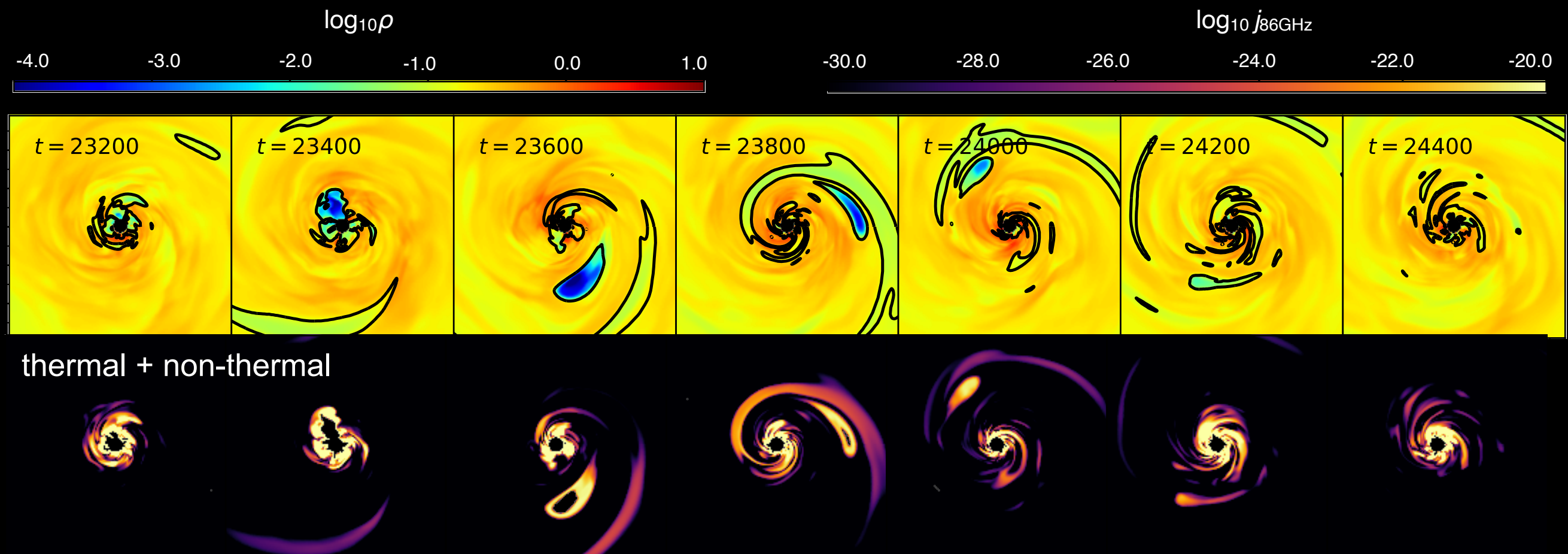
—> role of MAD flares (ejection of material)



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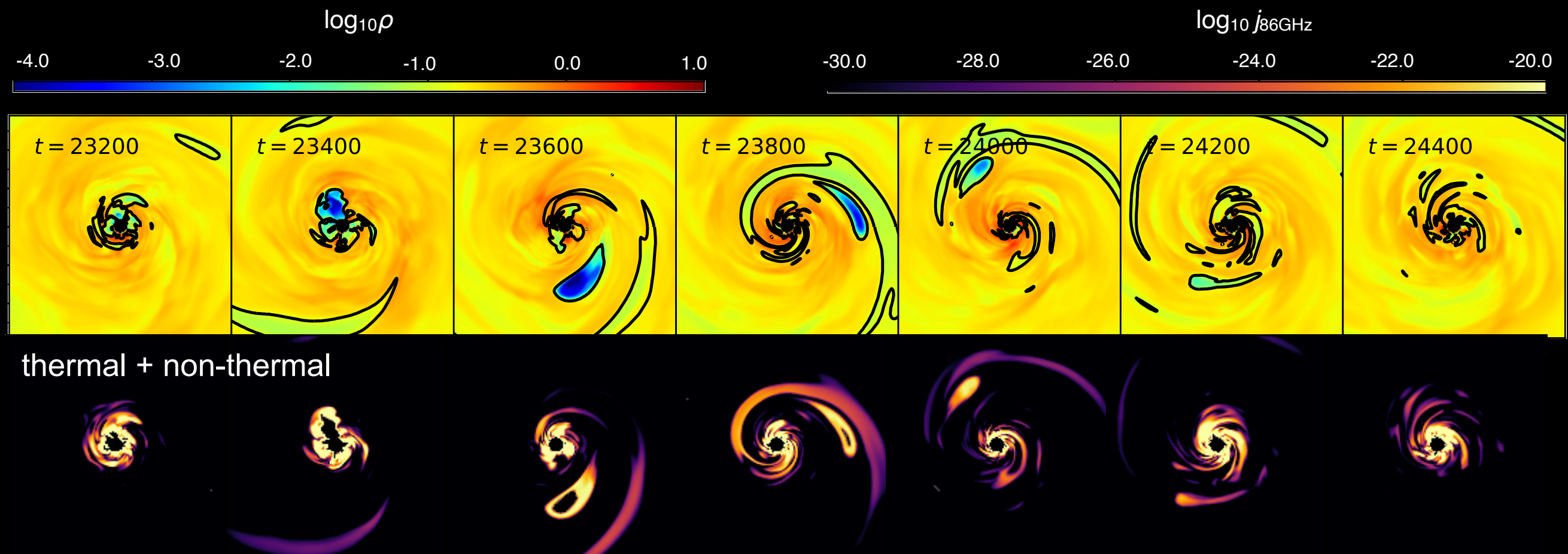


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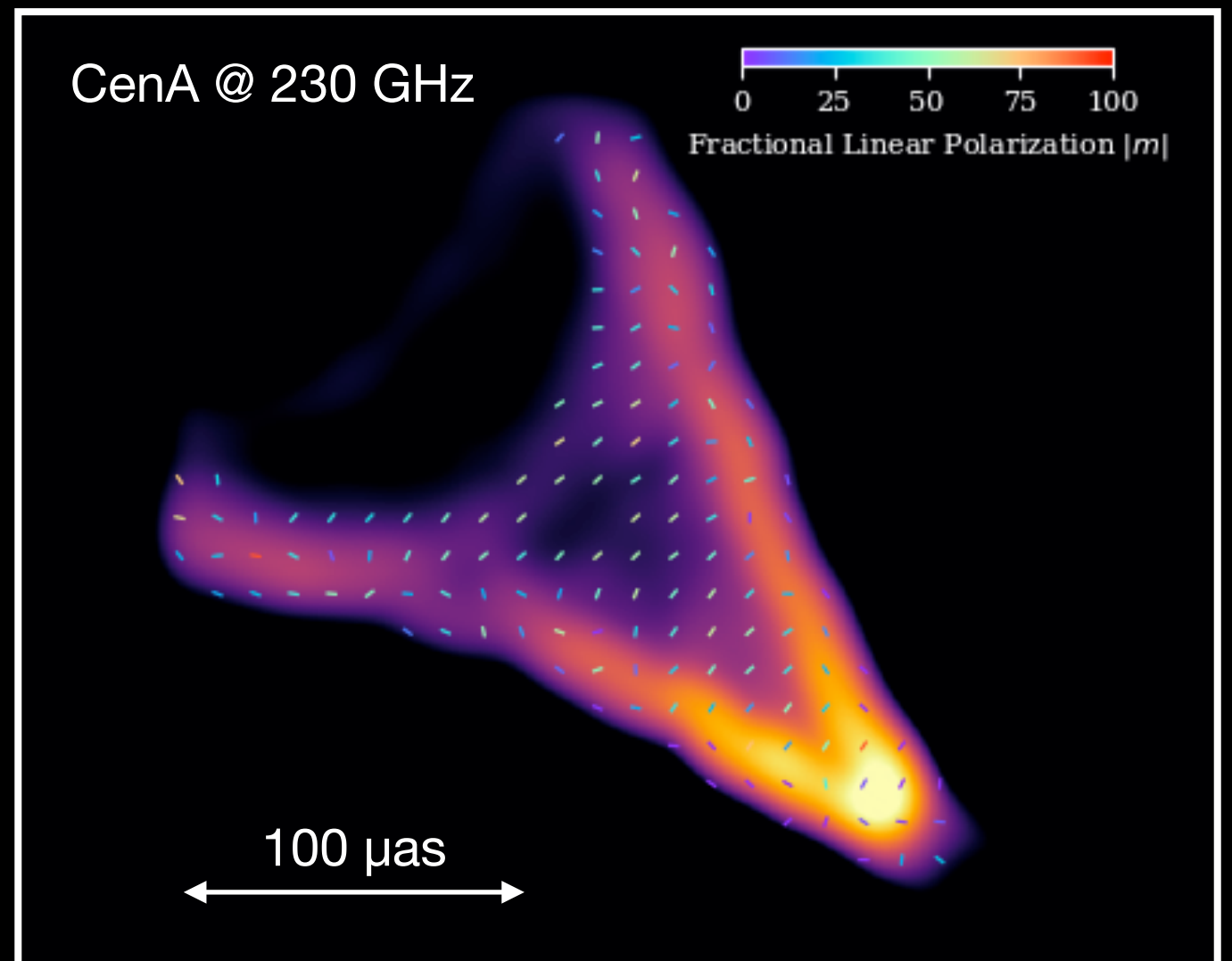
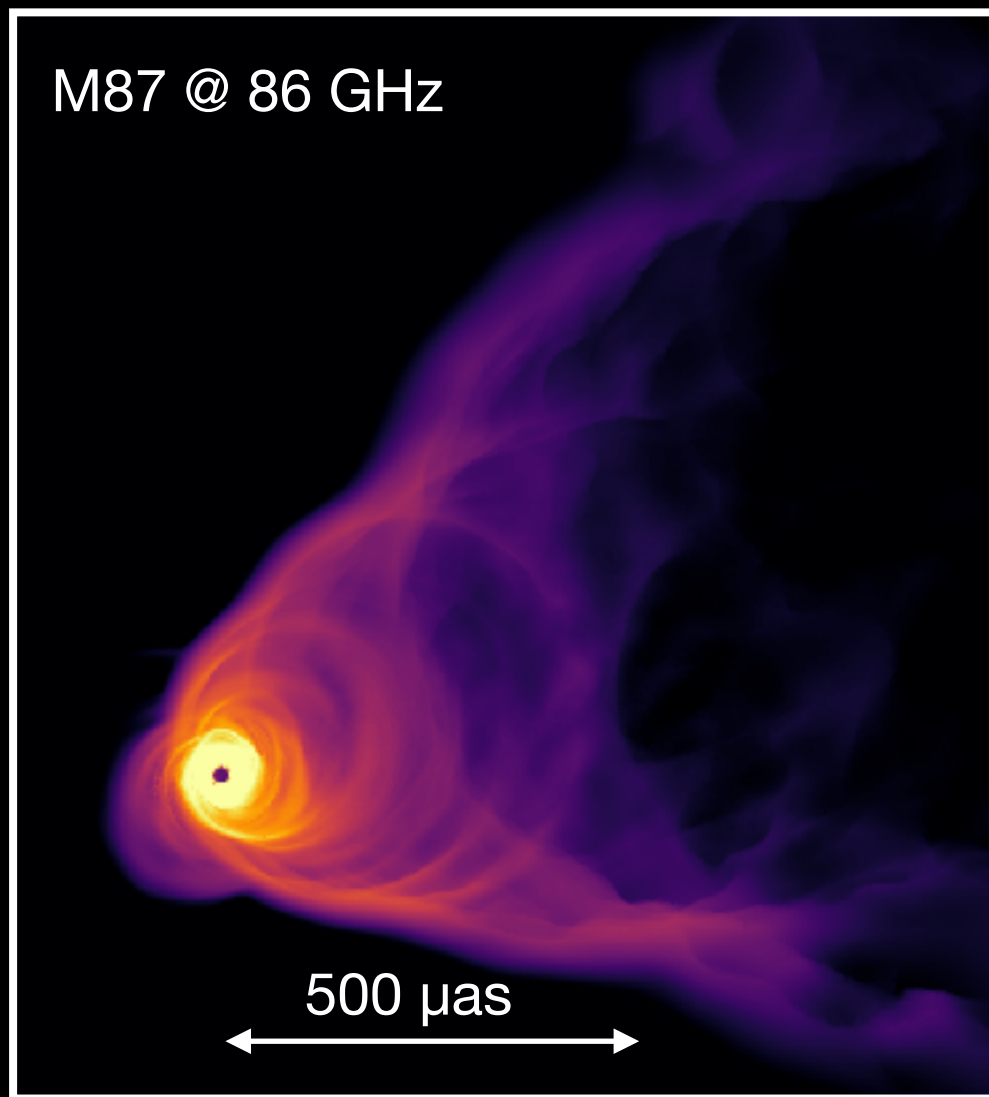
—> better understanding of the physical parameters of M87 (spin, etc)



Radiation Microphysics and PIC simulations

$$f(\gamma, \xi) = \phi(\xi) f_{\text{iso}}(\gamma)$$

$$\phi(\xi) \propto [1 + (\eta - 1) \cos^2 \xi]^{-p/2}$$

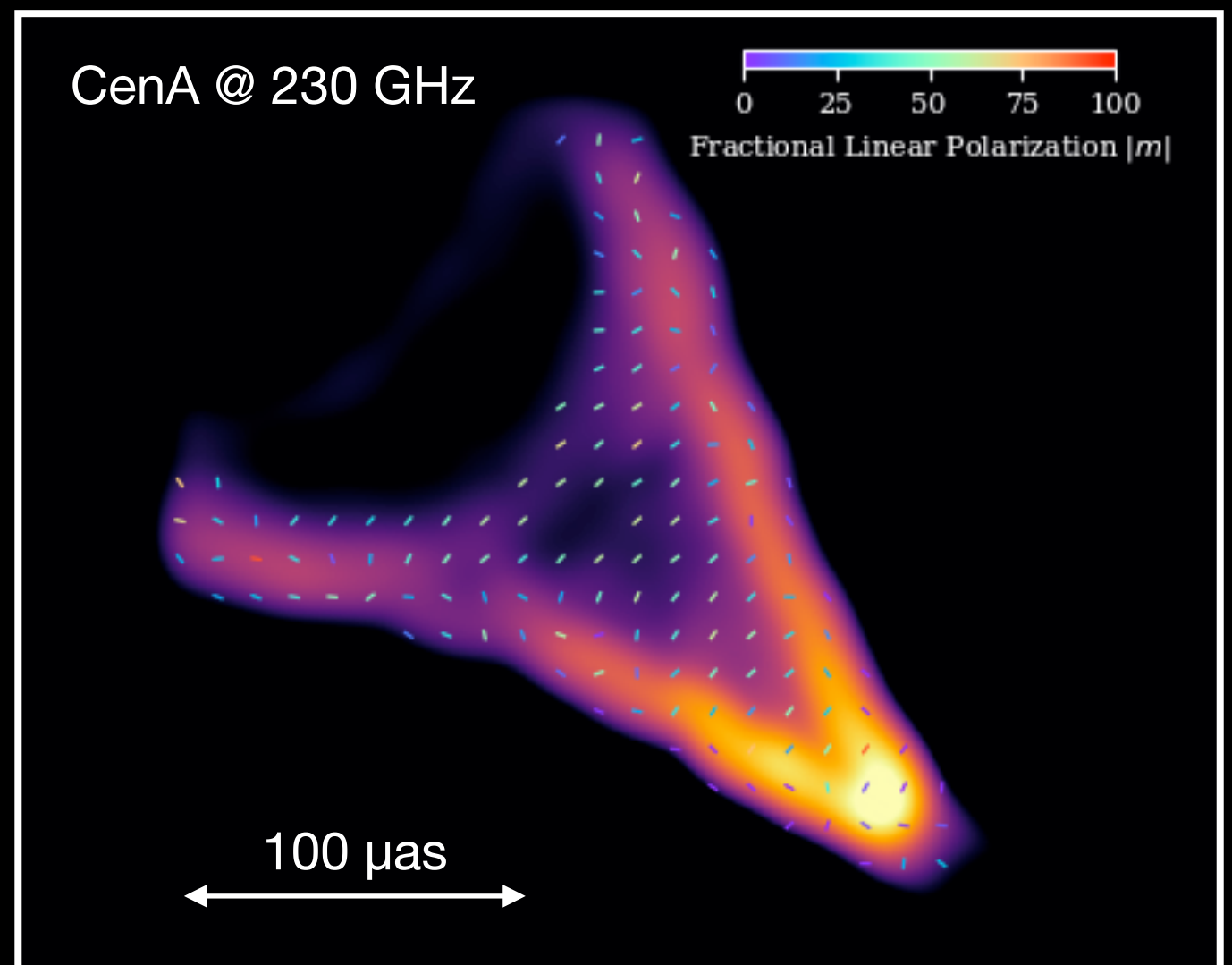
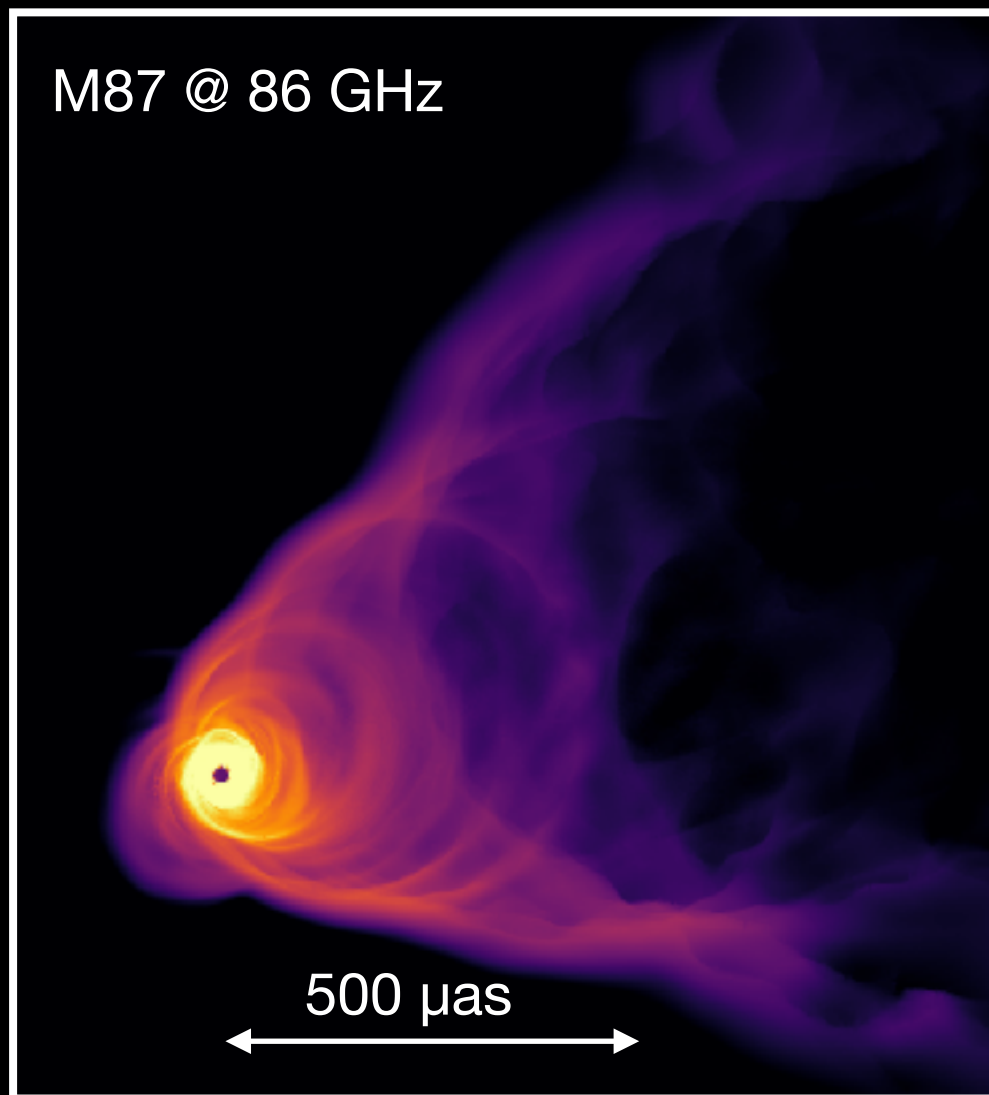


Radiation Microphysics and PIC simulations

mag. reconnection leads to anisotropic distribution of emitting particles

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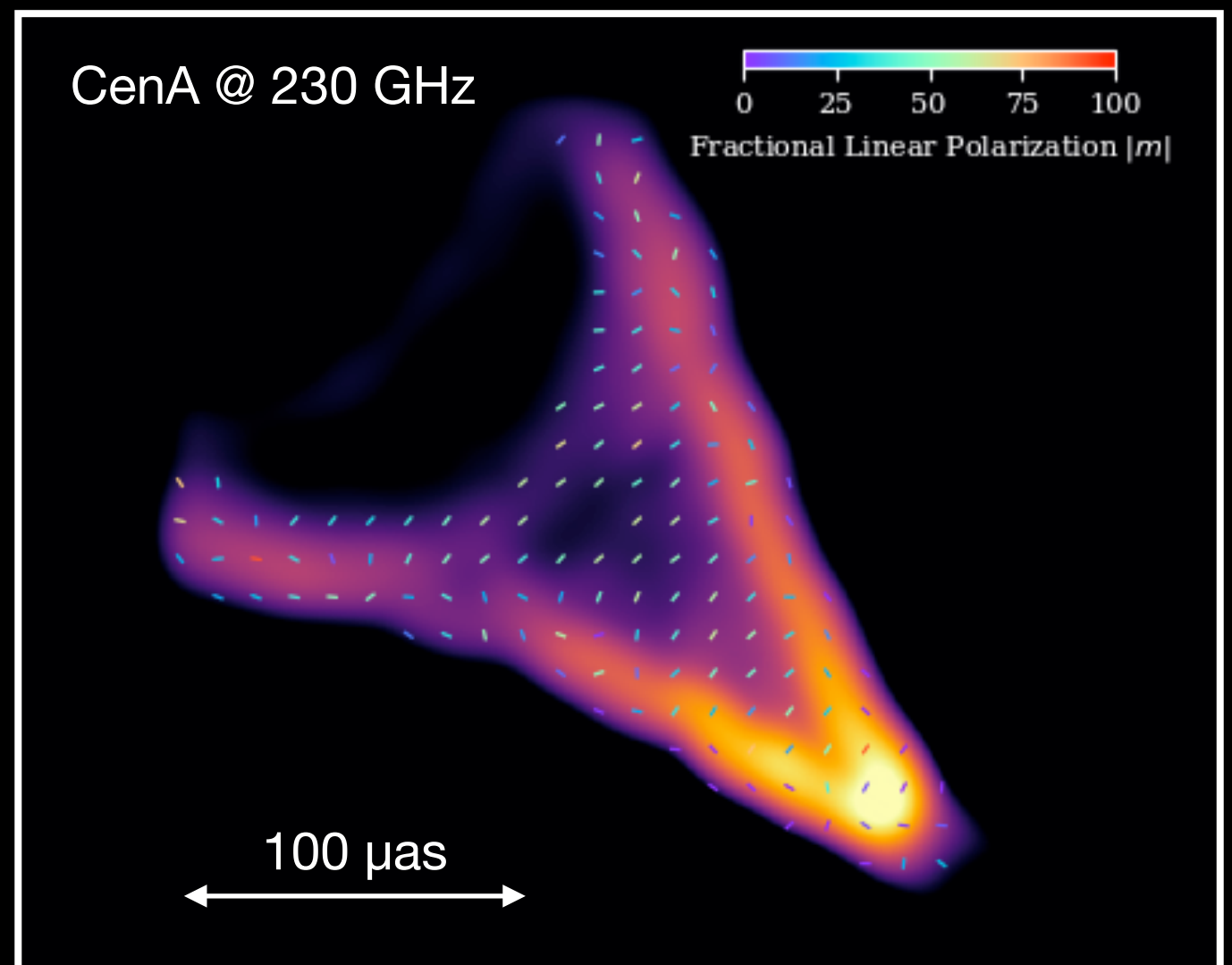
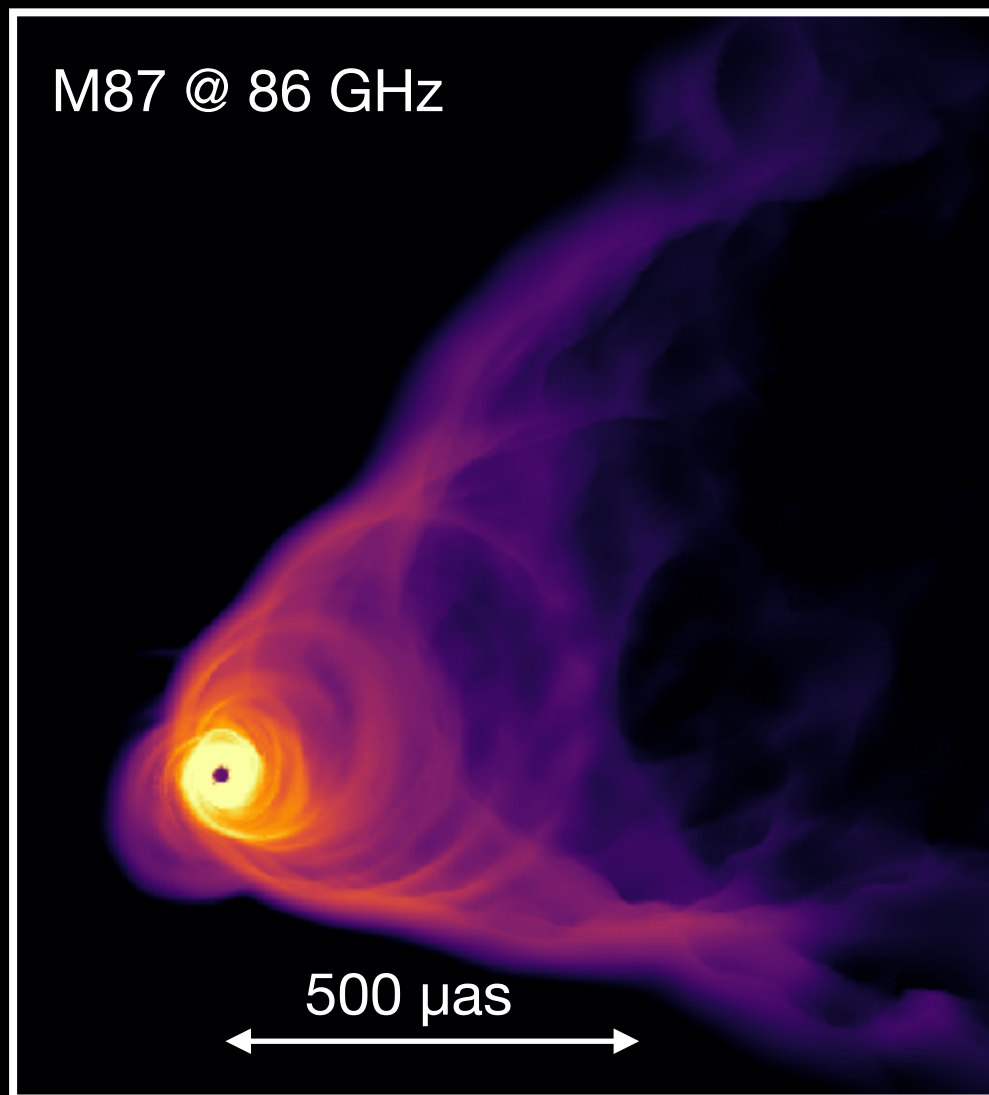
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Radiation Microphysics and PIC simulations

mag. reconnection leads to anisotropic distribution of emitting particles
—> strong edge-brightening of jets and “empty” jet spines

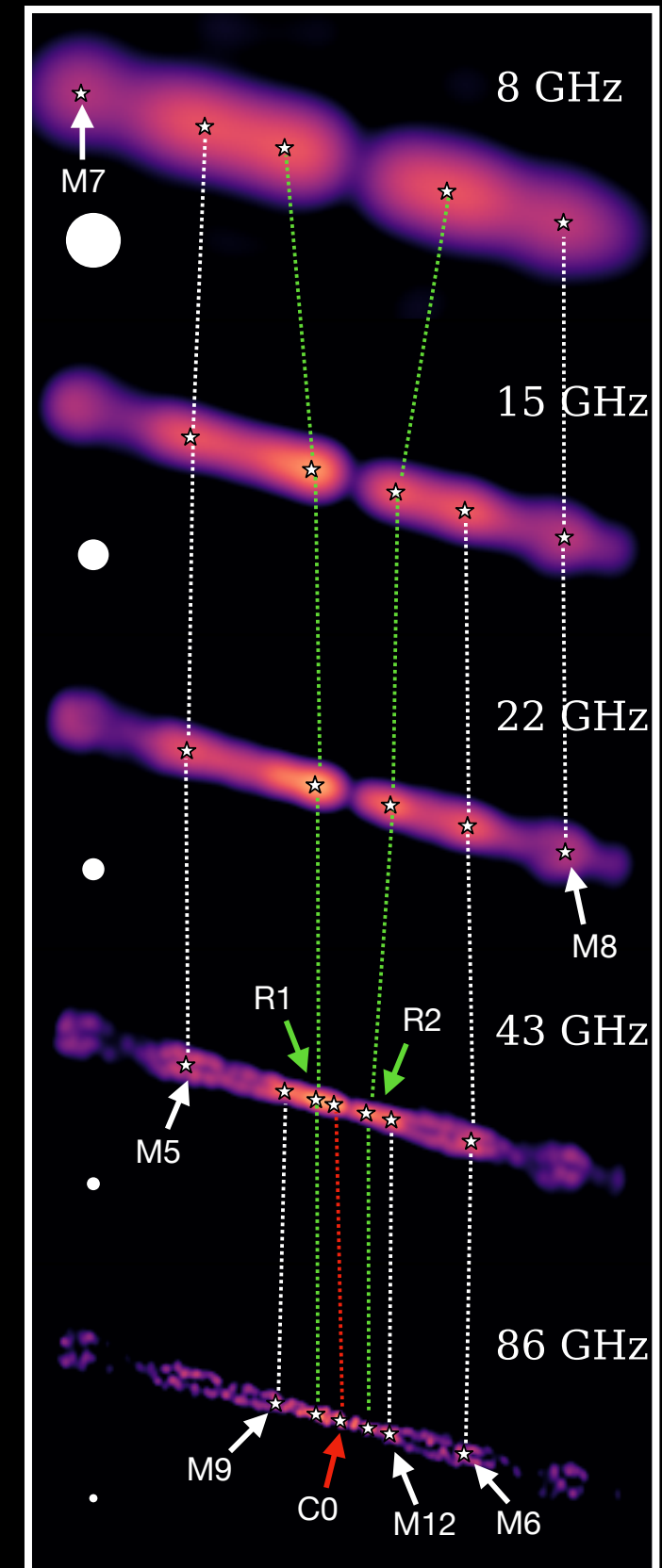
$$f(\gamma, \xi) = \phi(\xi) f_{\text{iso}}(\gamma) \quad \phi(\xi) \propto [1 + (\eta - 1) \cos^2 \xi]^{-p/2}$$



Large-Scale Jets and Instabilities

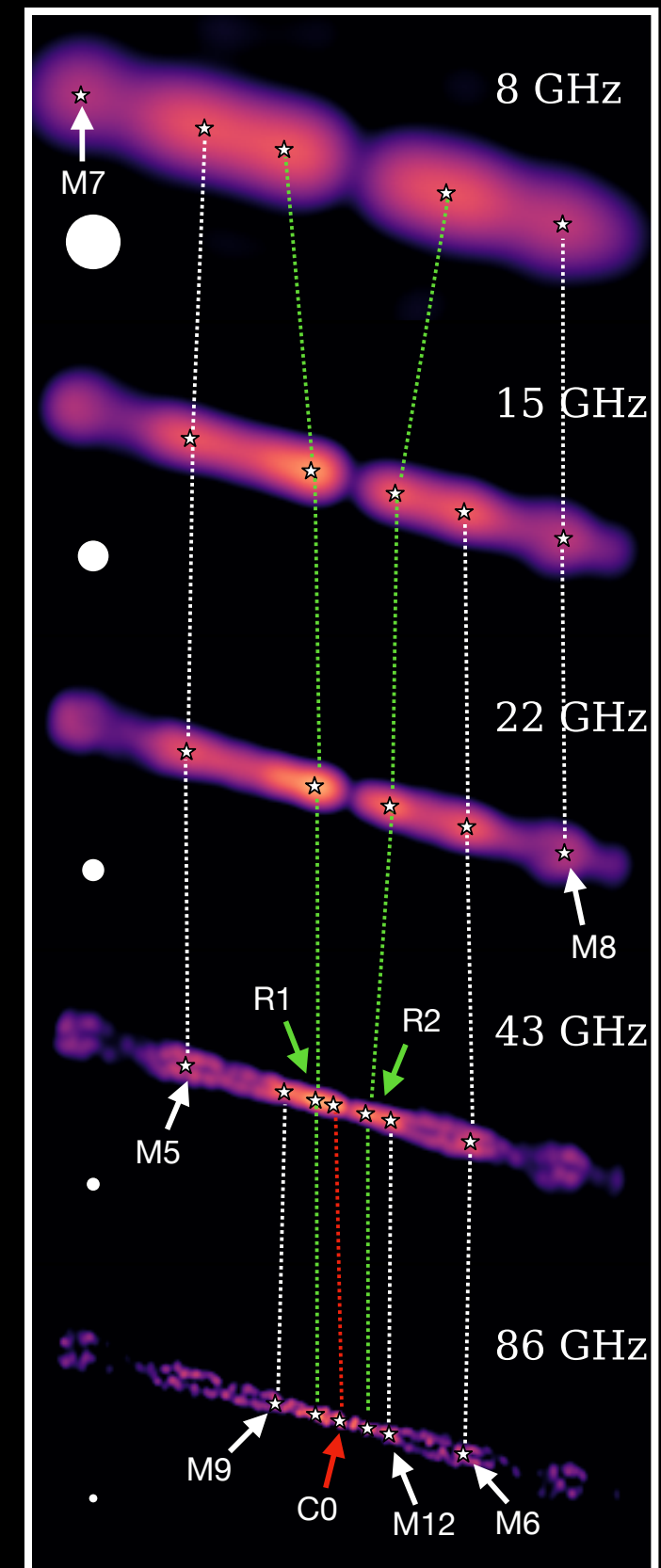
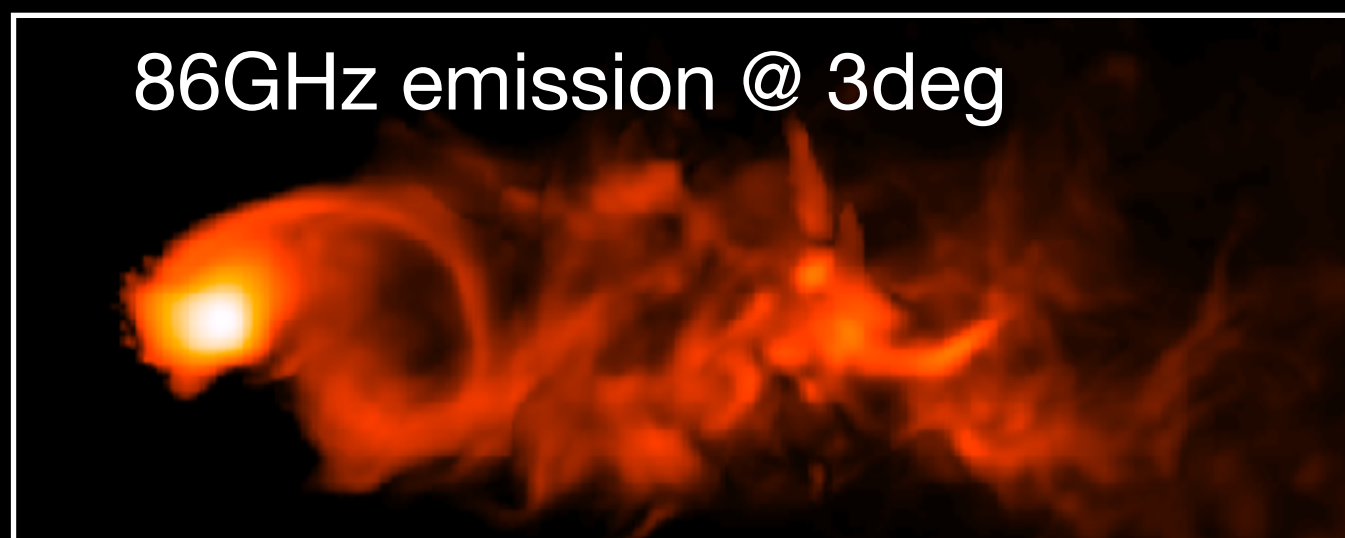
Large-Scale Jets and Instabilities

- **Jet asymmetry ($>1000r_g$)**
Developed End-to-End pipeline
Propagation of perturbations



Large-Scale Jets and Instabilities

- **Jet asymmetry ($>1000r_g$)**
Developed End-to-End pipeline
Propagation of perturbations
- **Jet asymmetry in 3D ($>1000r_g$)**
3D library of jet propagation
—> KH, RT, Richtmyer-Meshkov, Kink



Non-ideal GRMHD

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Include finite (global) resistivity

Non-ideal GRMHD

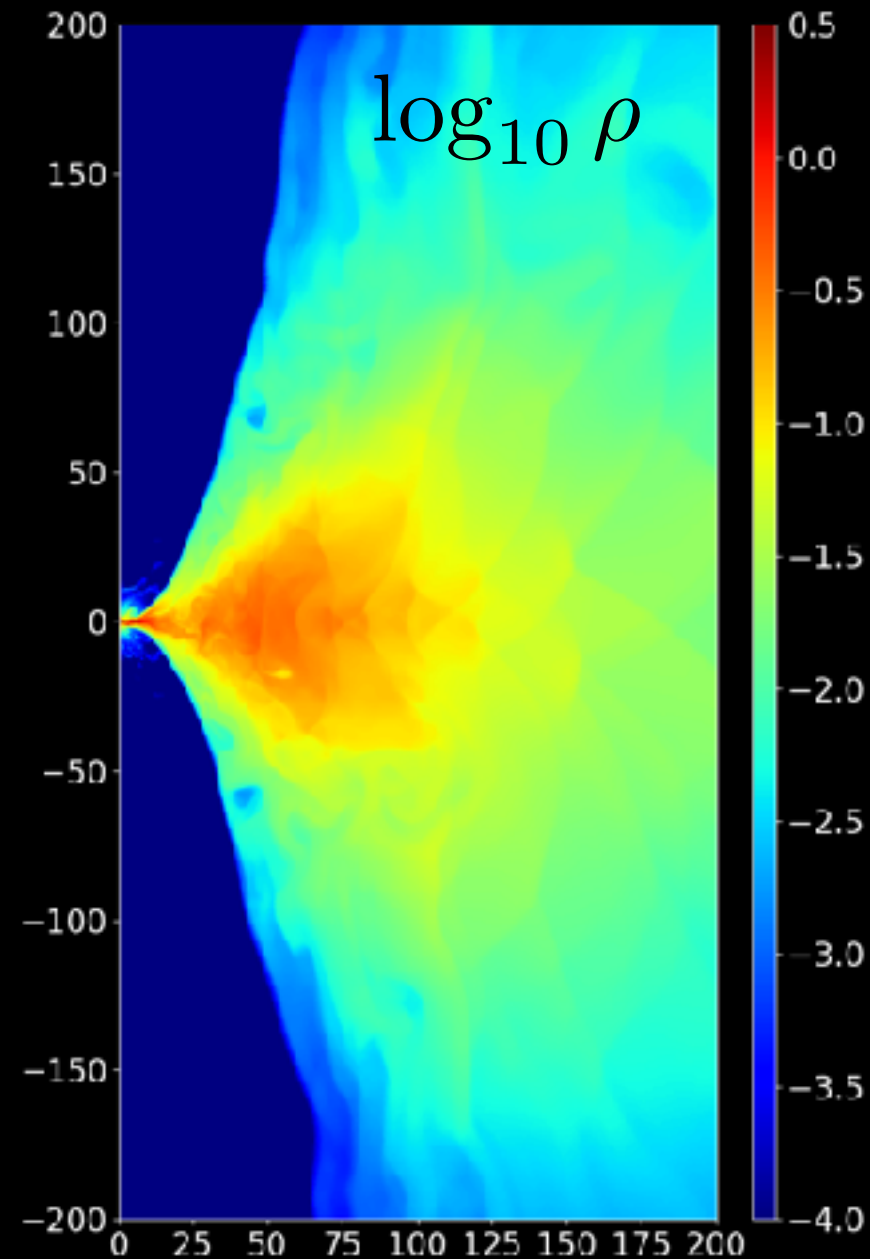
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physical magnetic field diffusion and mag. reconnection

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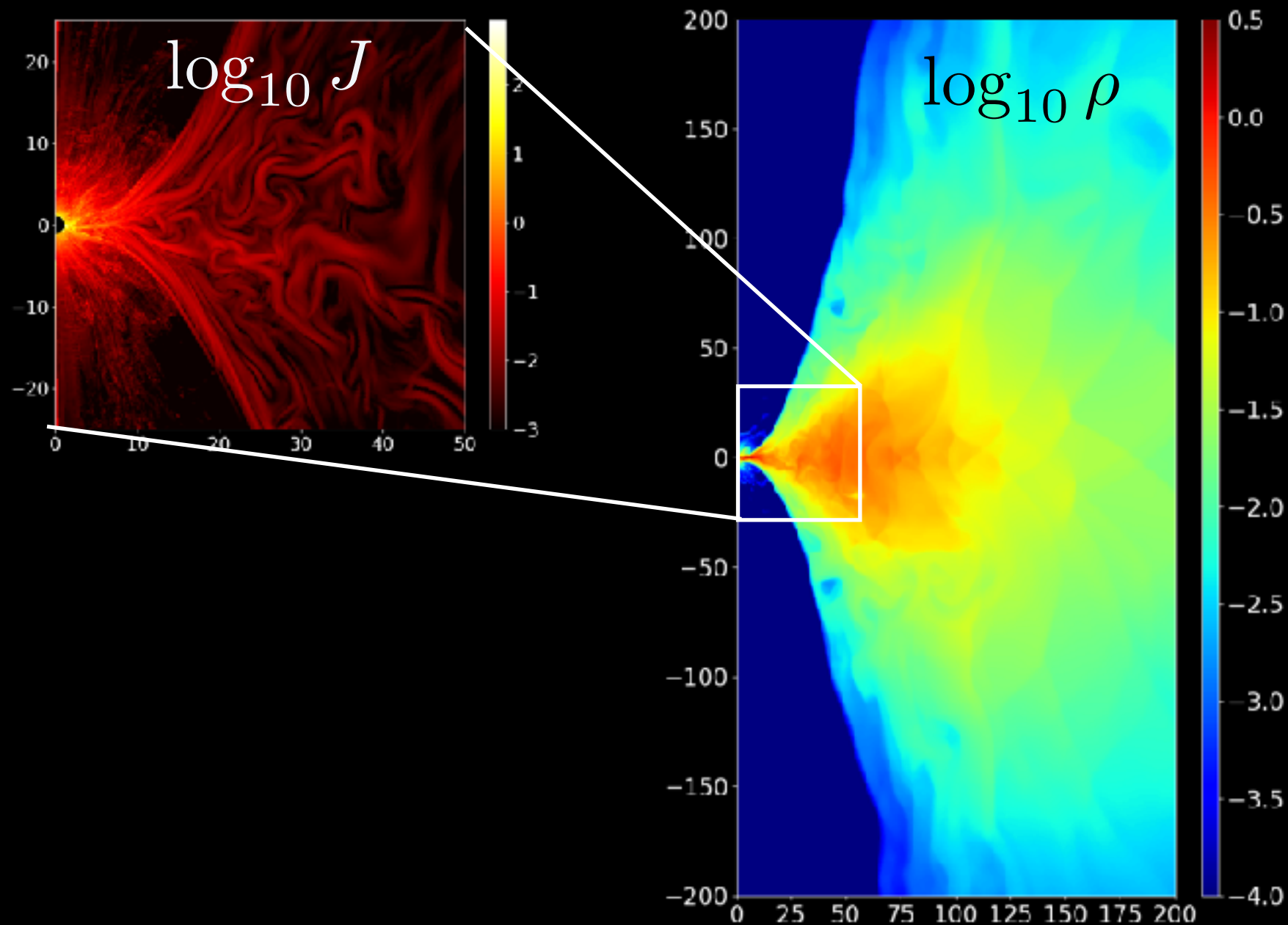
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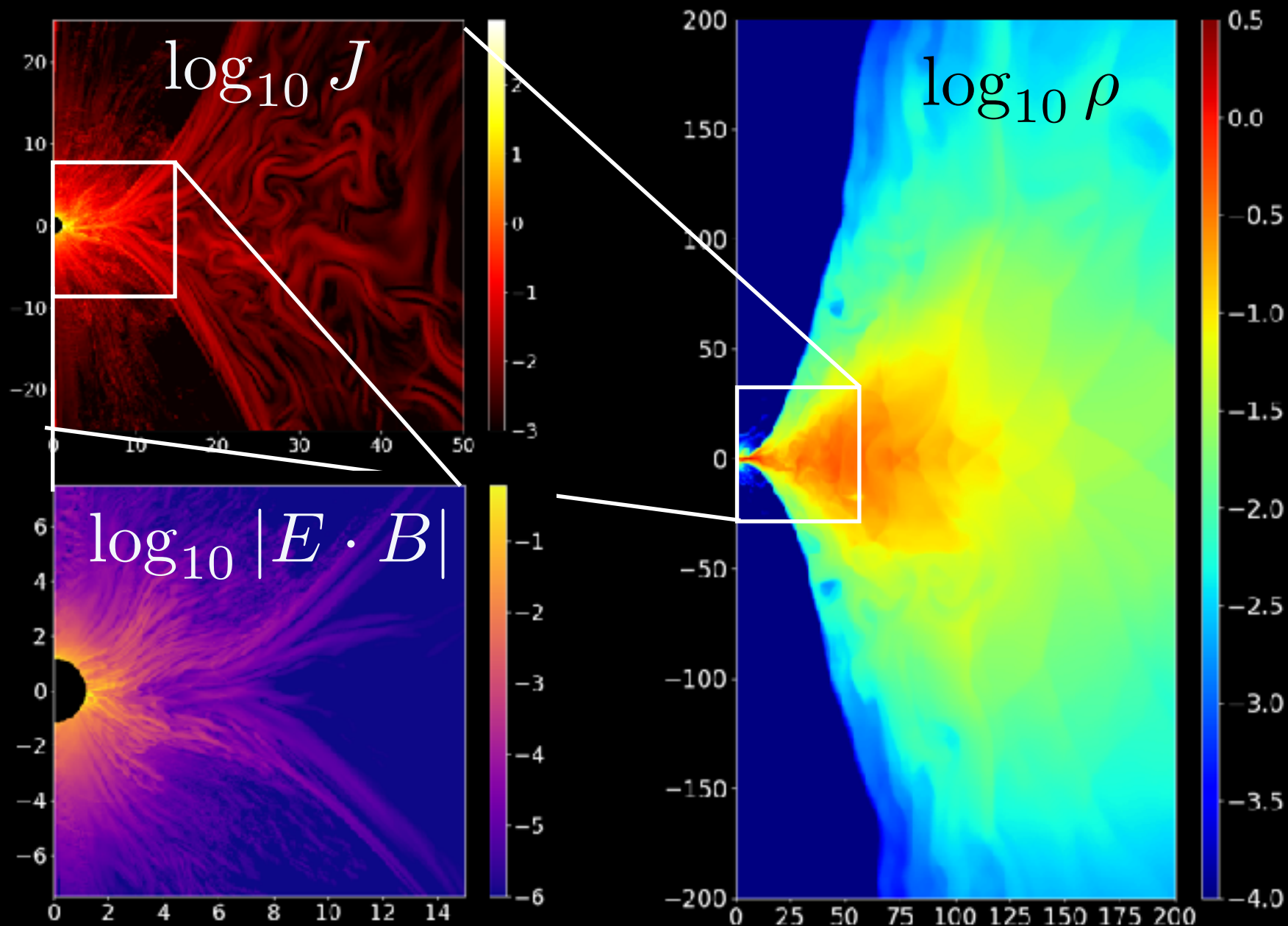
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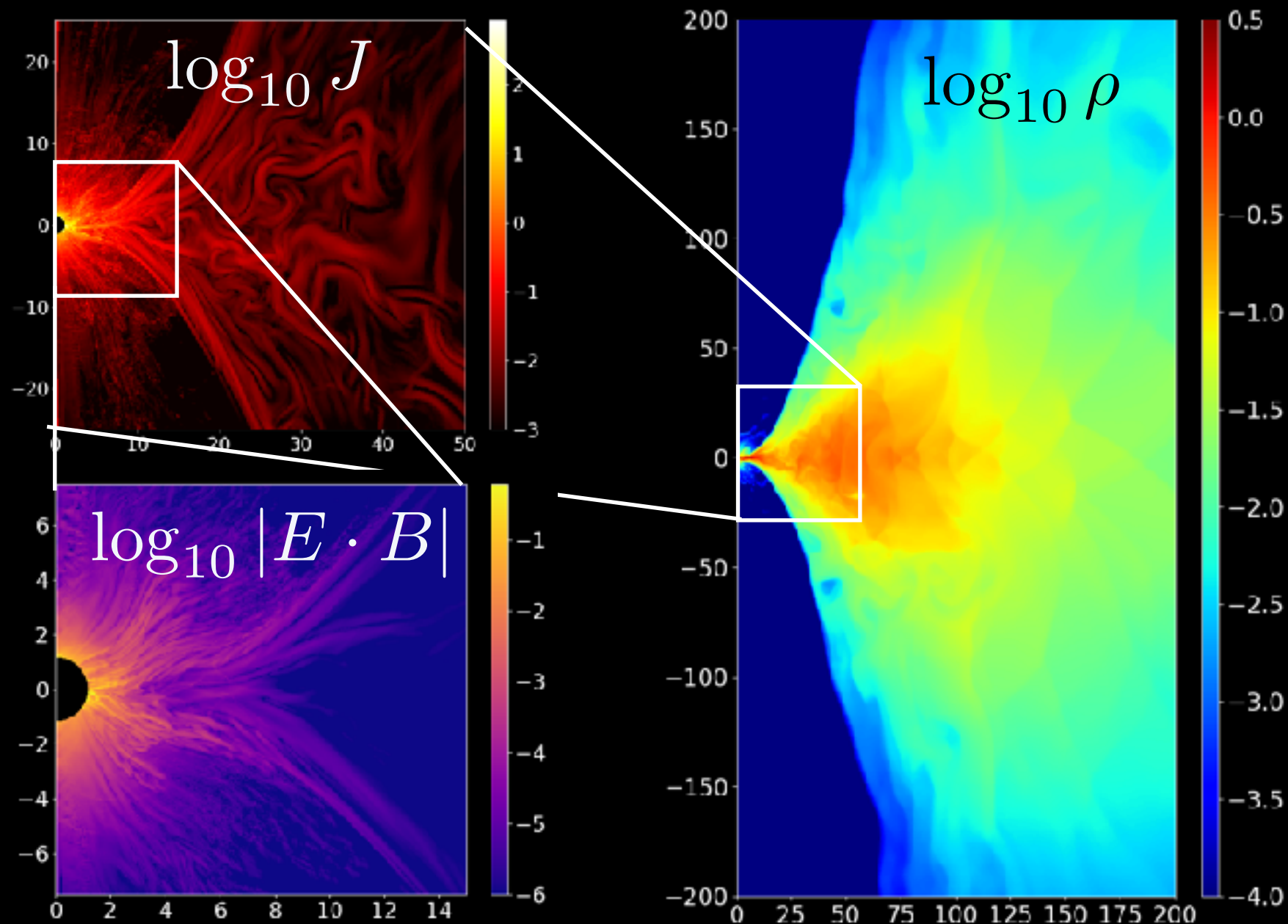
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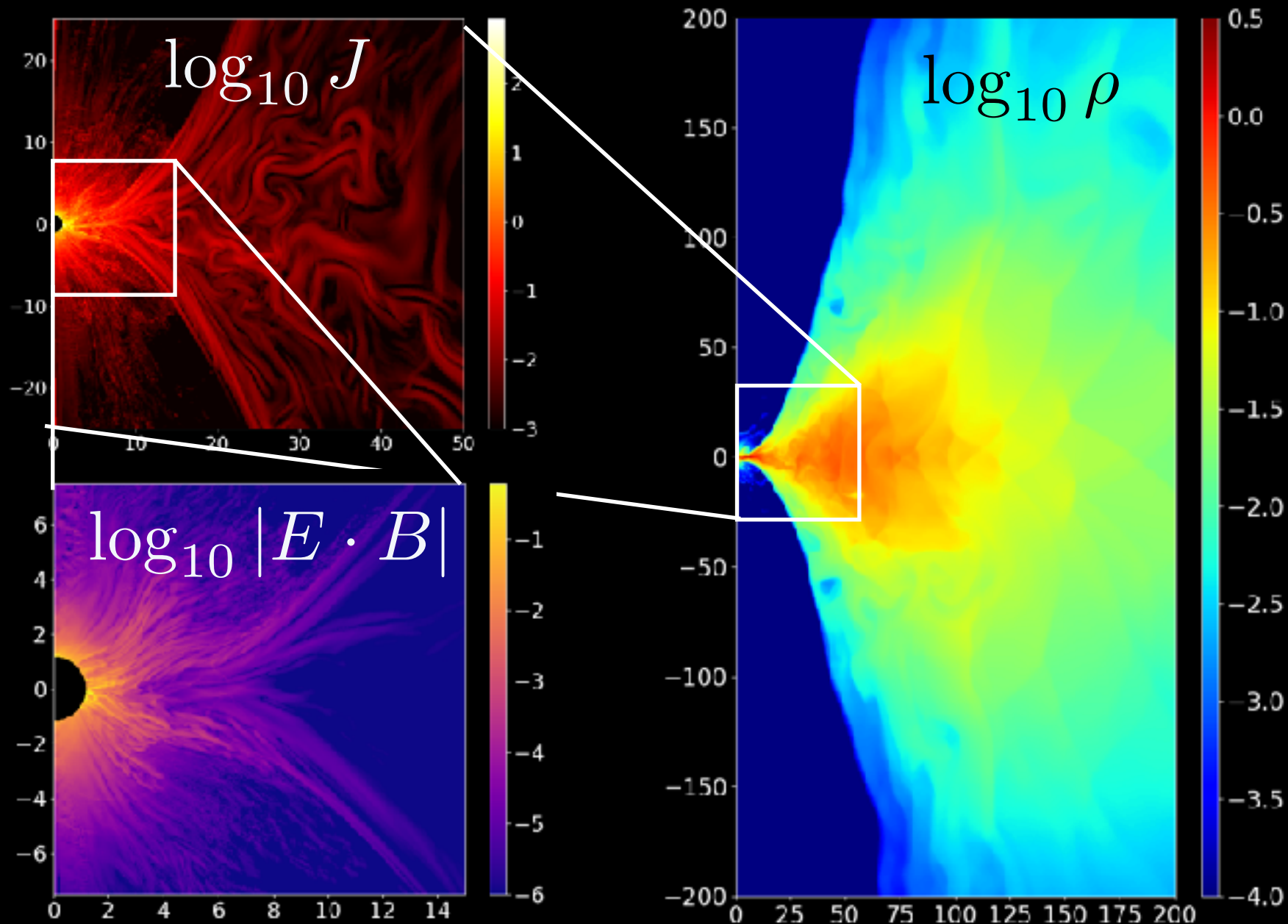


—> formation of current sheets, plasmoids, strong winds —>still MAD?

Non-ideal GRMHD

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physical magnetic field diffusion and mag. reconnection



- > formation of current sheets, plasmoids, strong winds —> still MAD?
- > acceleration of particles —> origin of neutrinos?



Thanks for your attention!

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