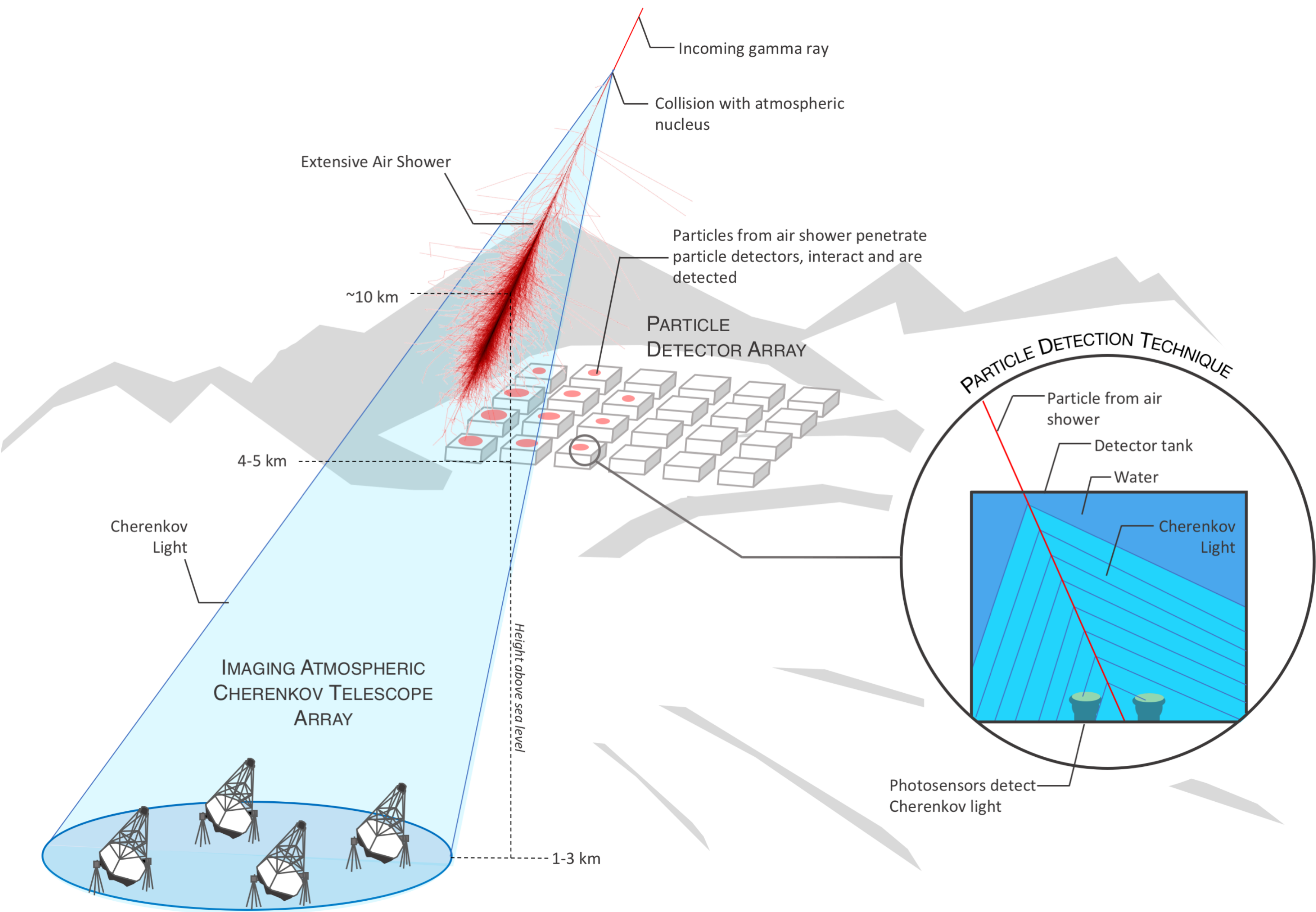


# An alternative way to sky maps

Gamma-ray astronomy × statistical physics.

# Detection mechanism at the TeV scales

*Use atmosphere as a calorimeter.*



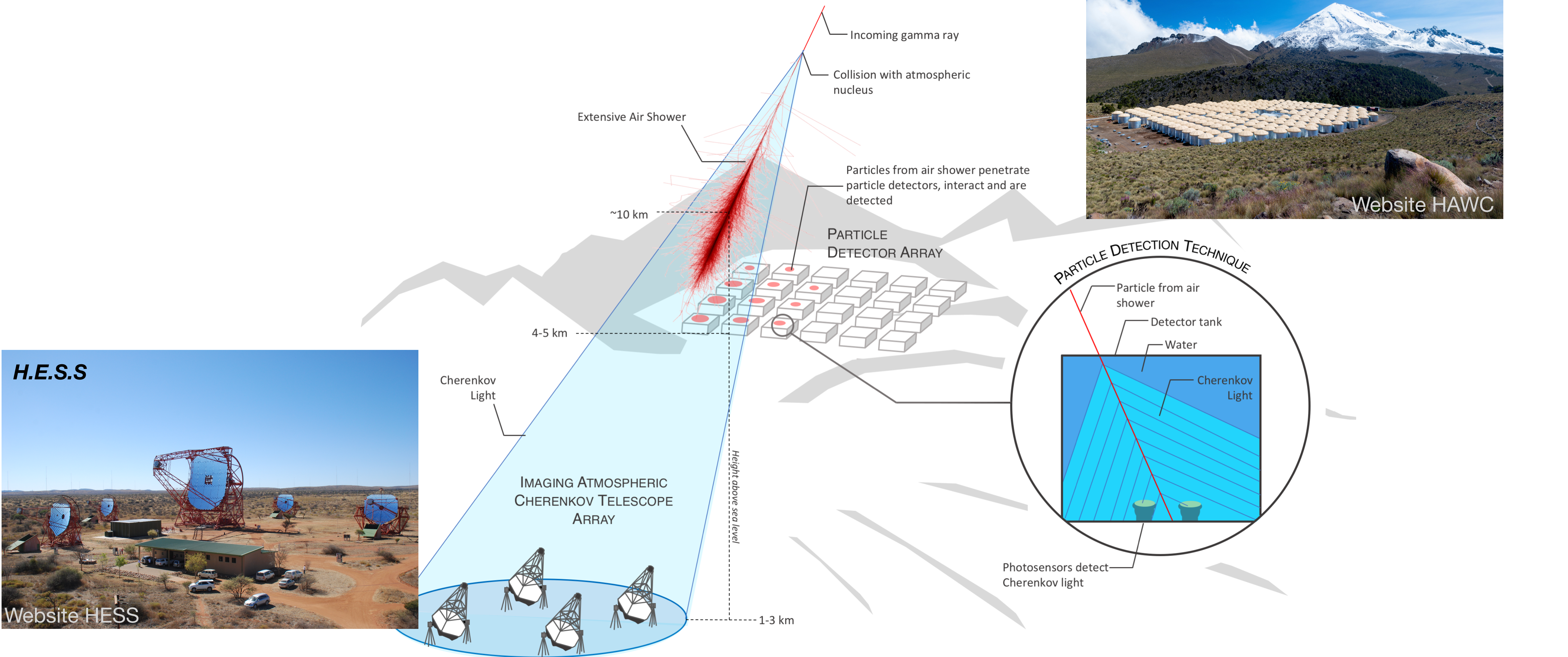
Shower image, 100 GeV  $\gamma$ -ray adapted from: F. Schmidt, J. Knapp, "CORSIKA Shower Images", 2005, <https://www-zeuthen.desy.de/~jknapp/fs/showerimages.html>

*Not to scale*



# Detection mechanism at the TeV scales

Use atmosphere as a calorimeter.

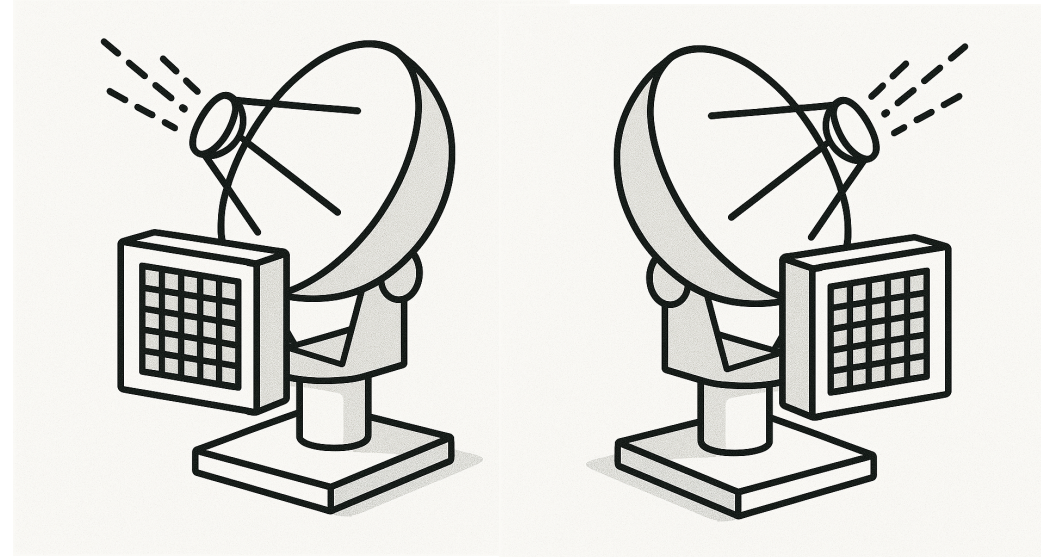


Not to scale



# Background modeling

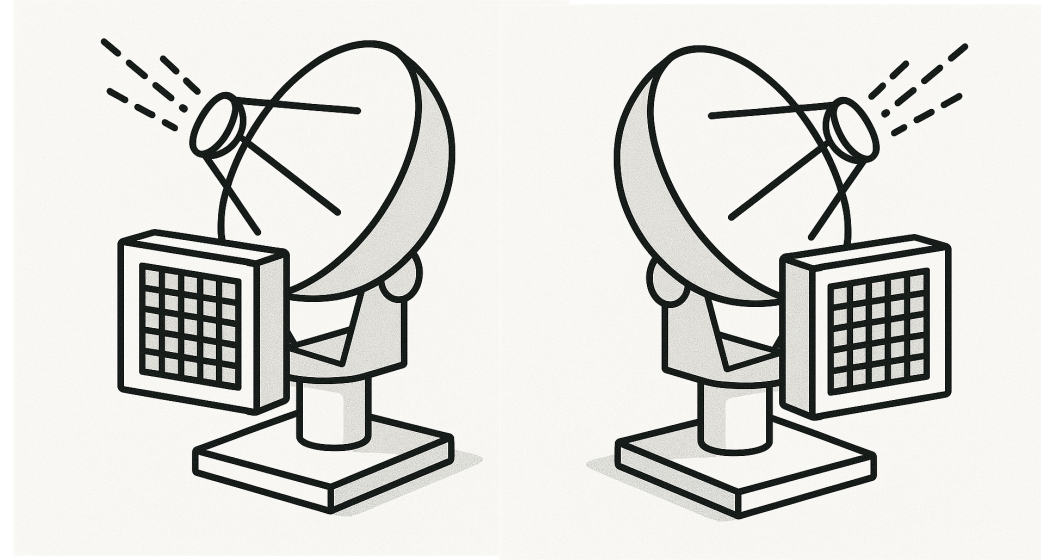
- Gamma/hadron separation is not perfect → hadronic background
- We need to estimate the number of background events in the region of interest



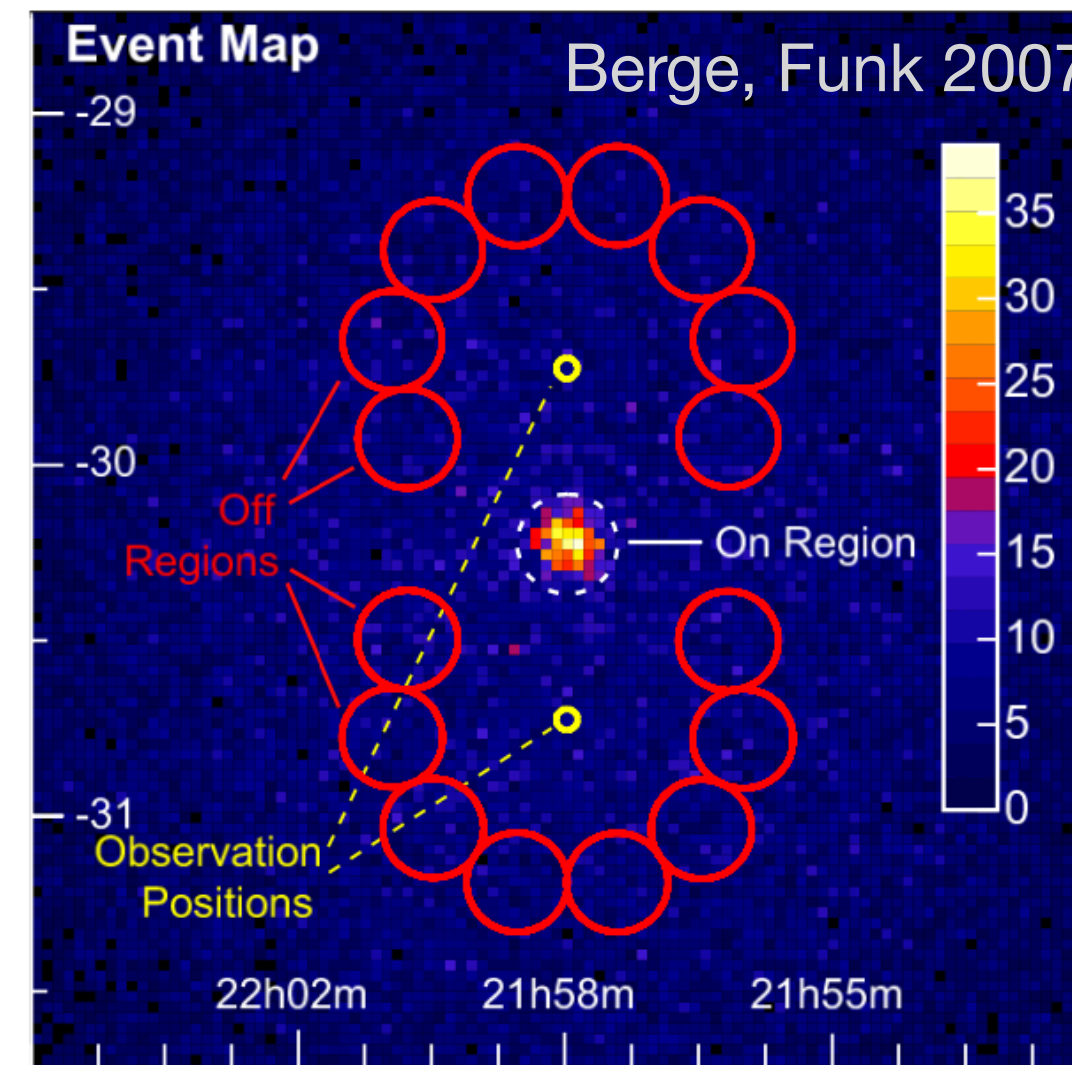


# Background modeling

- Gamma/hadron separation is not perfect  $\rightarrow$  hadronic background
- We need to estimate the number of background events in the region of interest



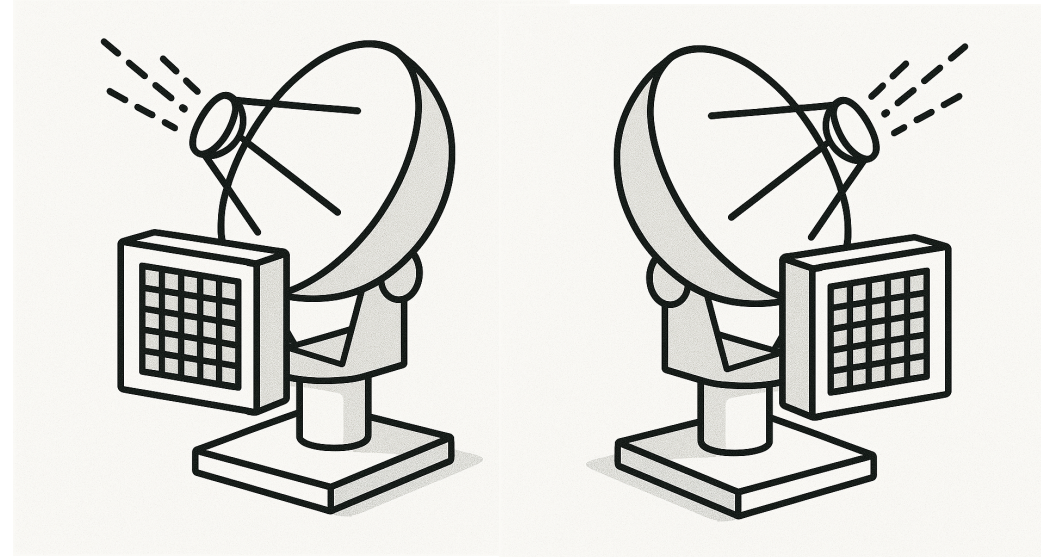
## Reflected Background



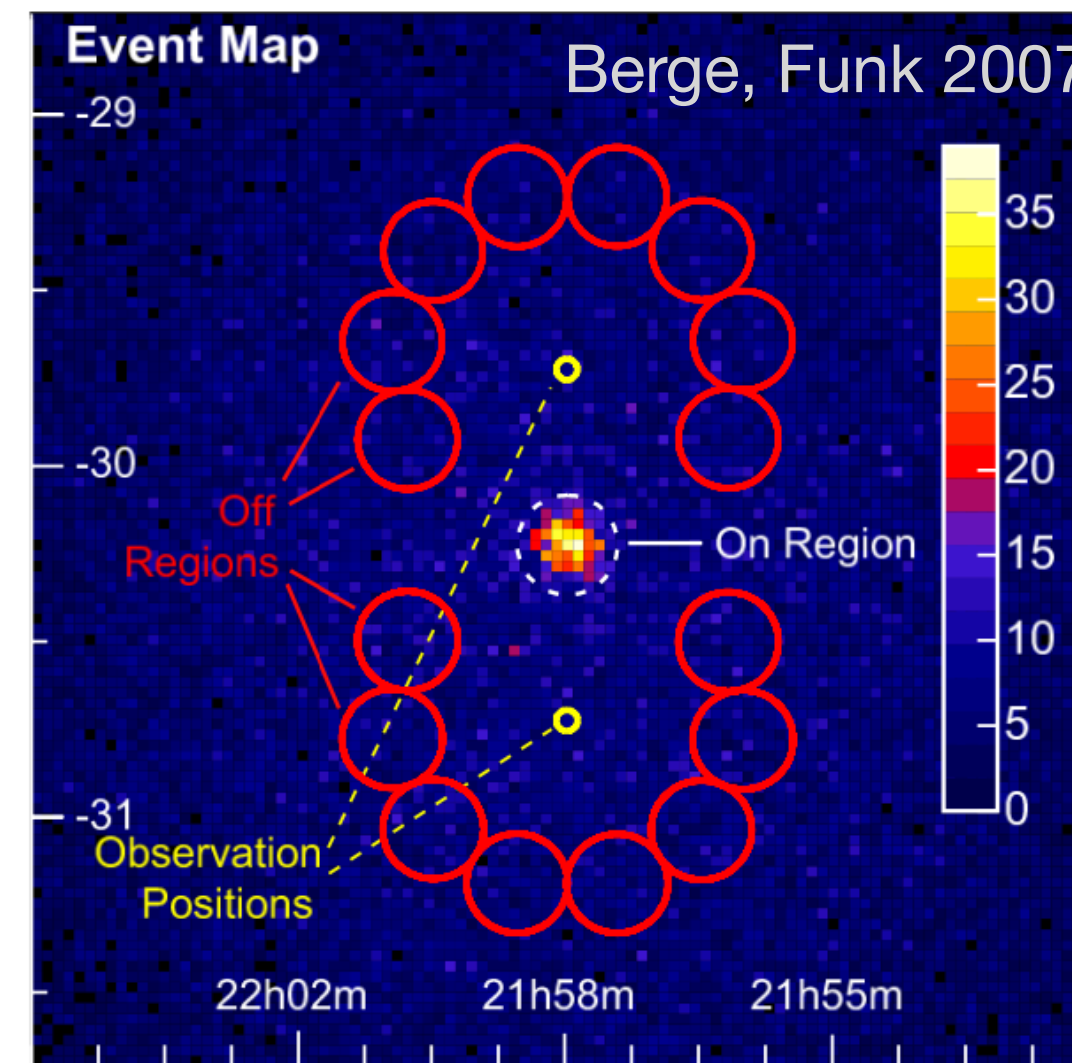


# Background modeling

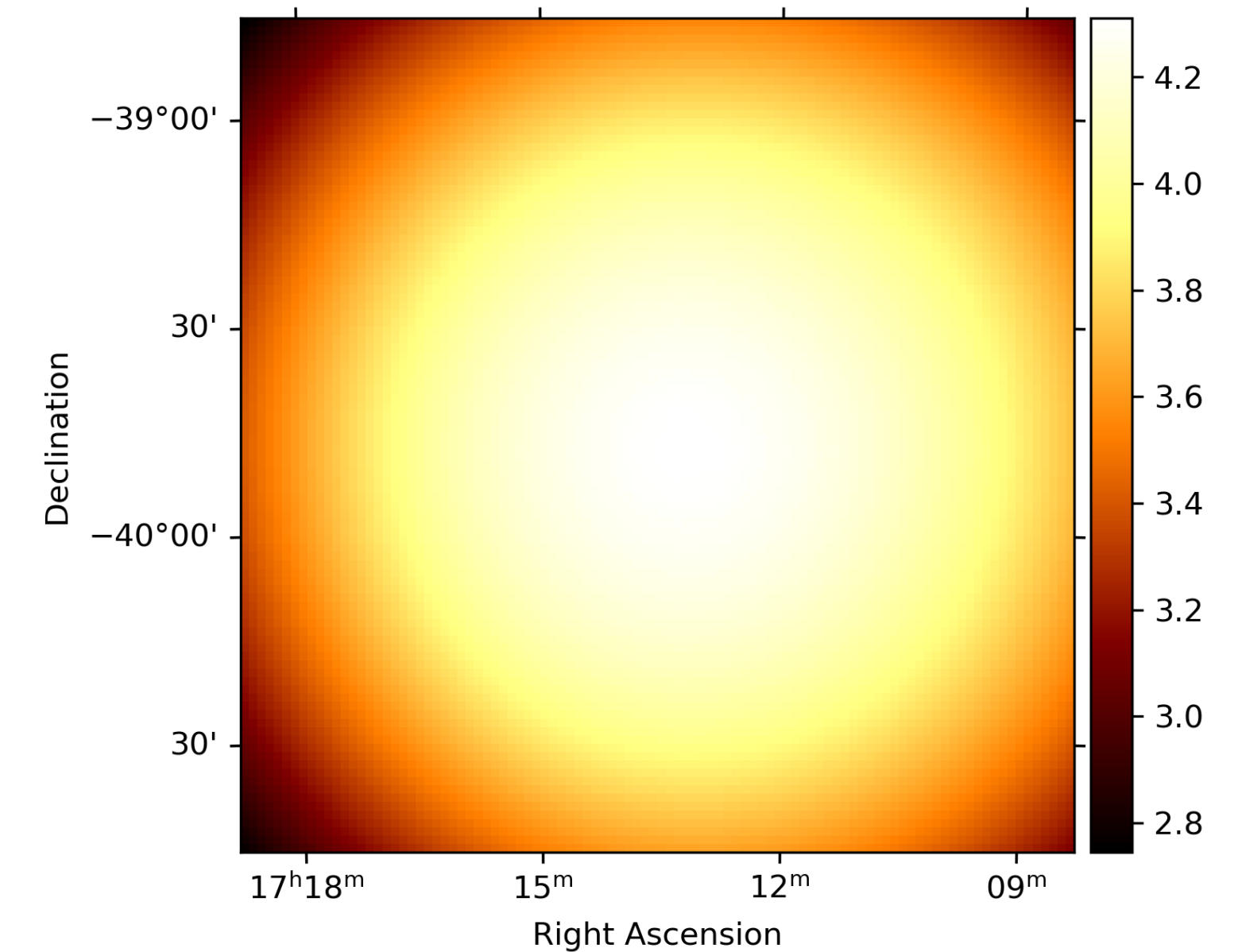
- Gamma/hadron separation is not perfect → hadronic background
- We need to estimate the number of background events in the region of interest



Reflected Background



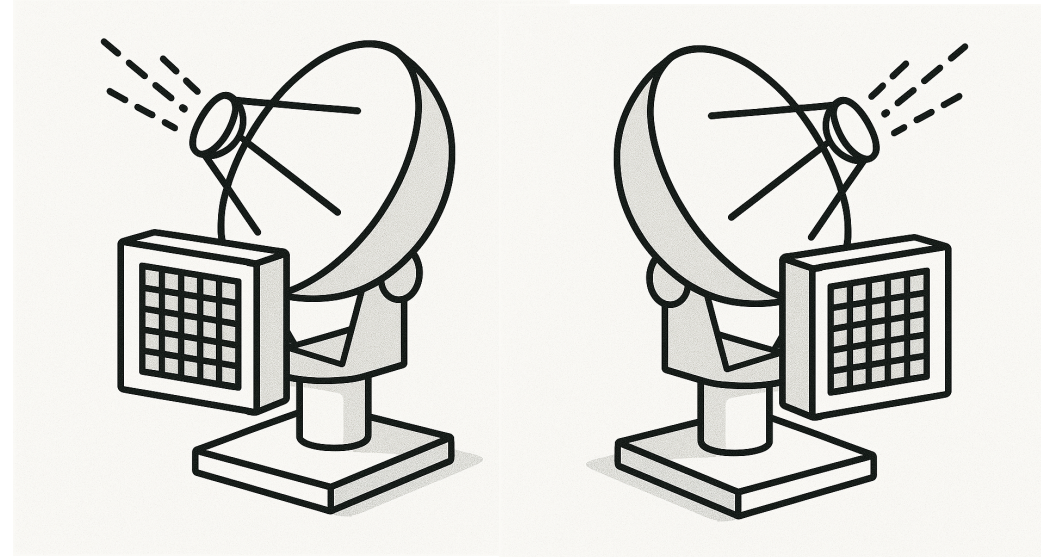
Background template



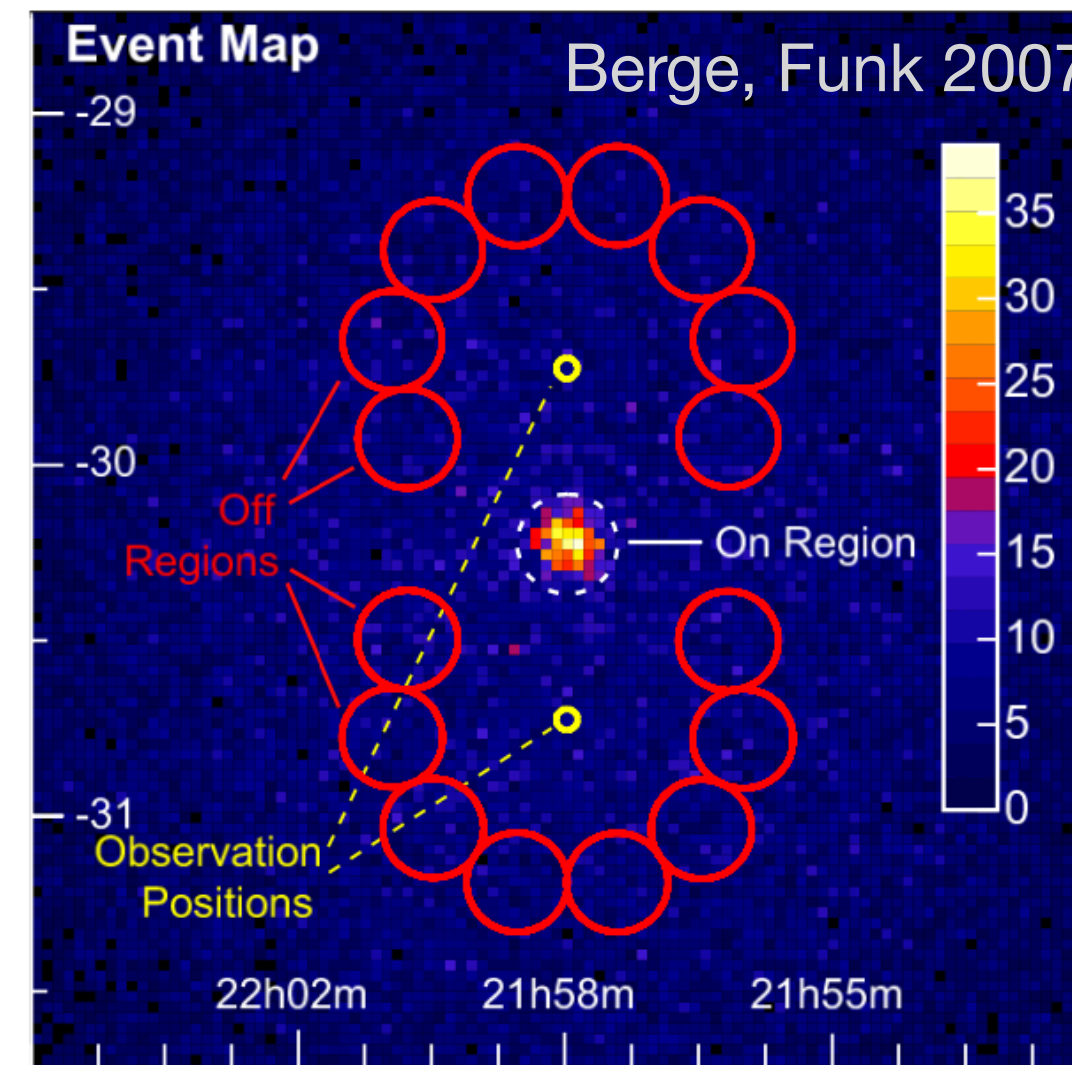


# Background modeling

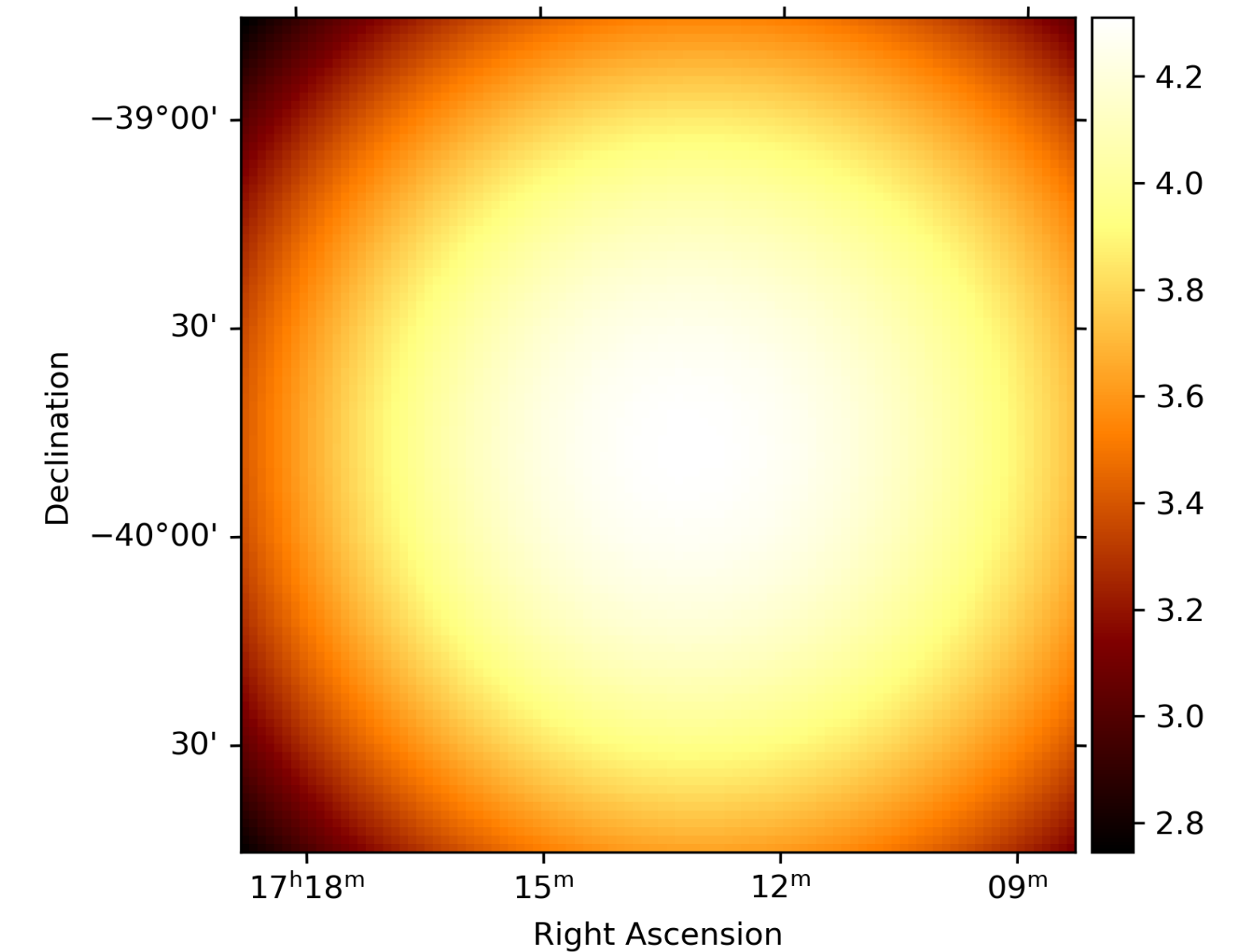
- Gamma/hadron separation is not perfect → hadronic background
- We need to estimate the number of background events in the region of interest



Reflected Background



Background template



**Lima Significance (1983)**

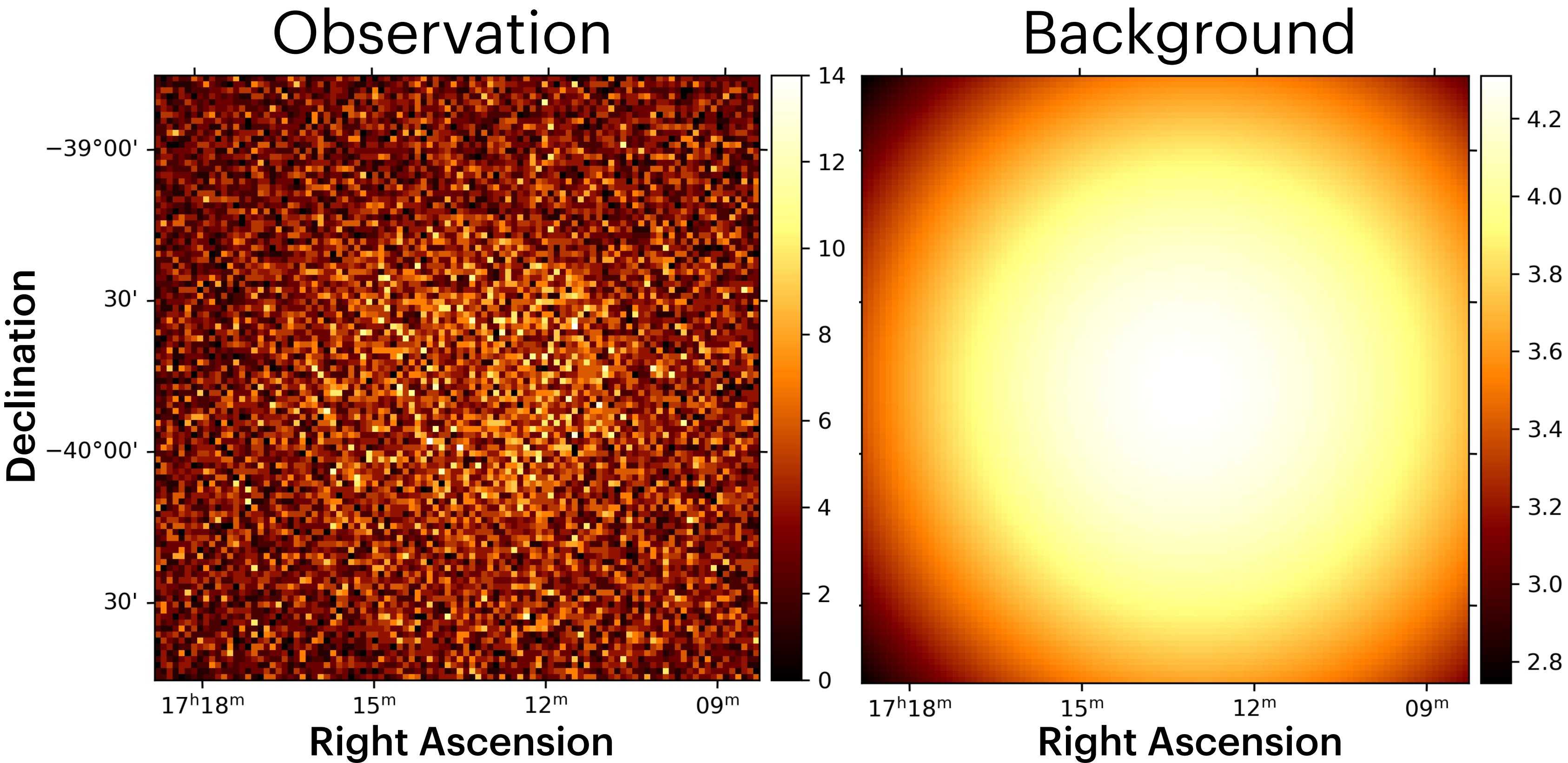
$$S(N_{\text{on}}, N_{\text{off}}, \alpha) = \sqrt{-2 \ln \Lambda} = \sqrt{2} \left\{ N_{\text{on}} \ln \left[ \frac{1 + \alpha}{\alpha} \frac{N_{\text{on}}}{N_{\text{on}} + N_{\text{off}}} \right] + N_{\text{off}} \ln \left[ (1 + \alpha) \frac{N_{\text{off}}}{N_{\text{on}} + N_{\text{off}}} \right] \right\}^{1/2}$$

*“S quantifies how likely it is that all counts from the “on region” were only due to background”*



# Example sky maps: RX J1713.7 3946

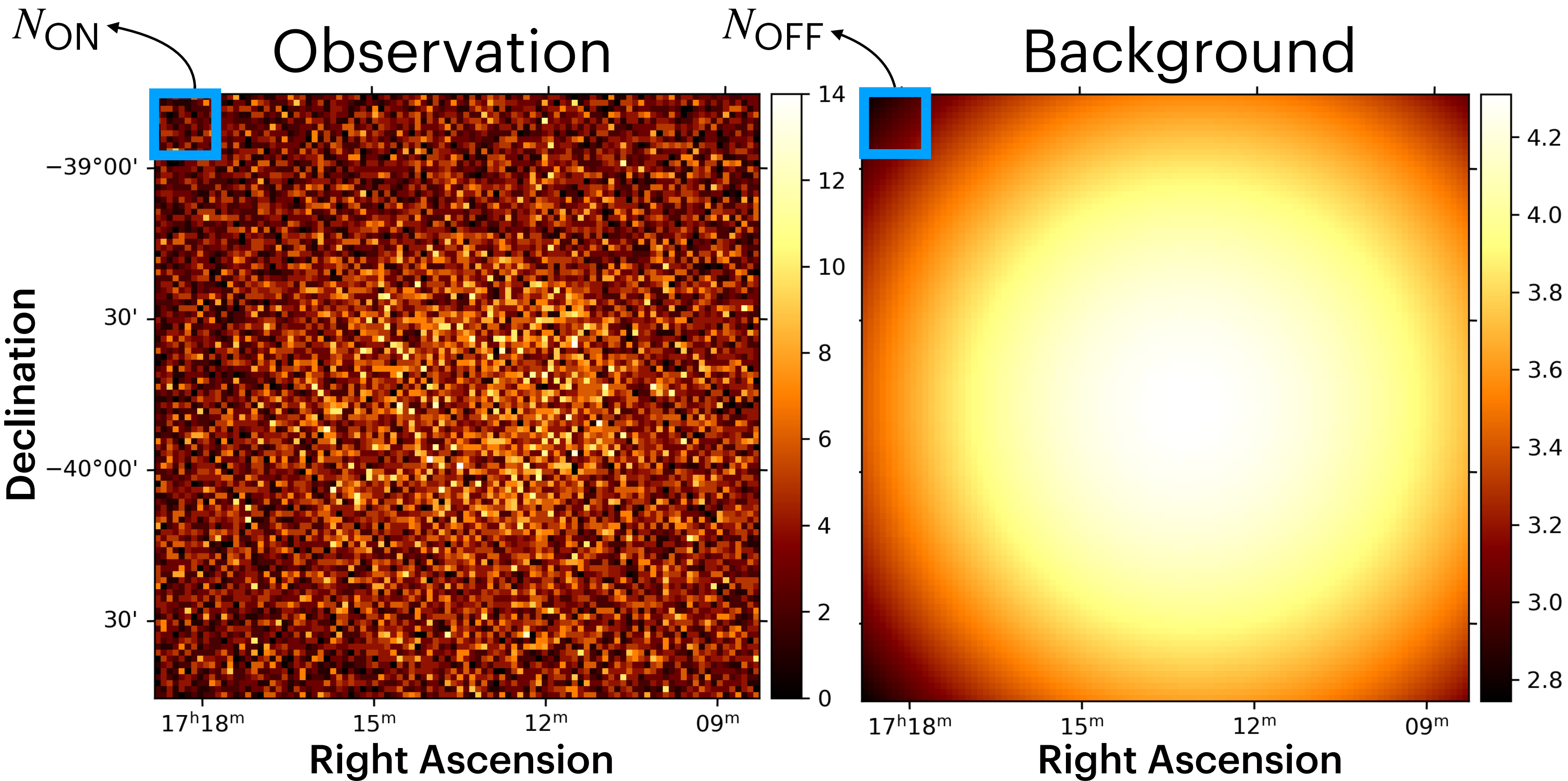
- 15 observations from 2004 released in HESS open data 2018





# Example sky maps: RX J1713.7 3946

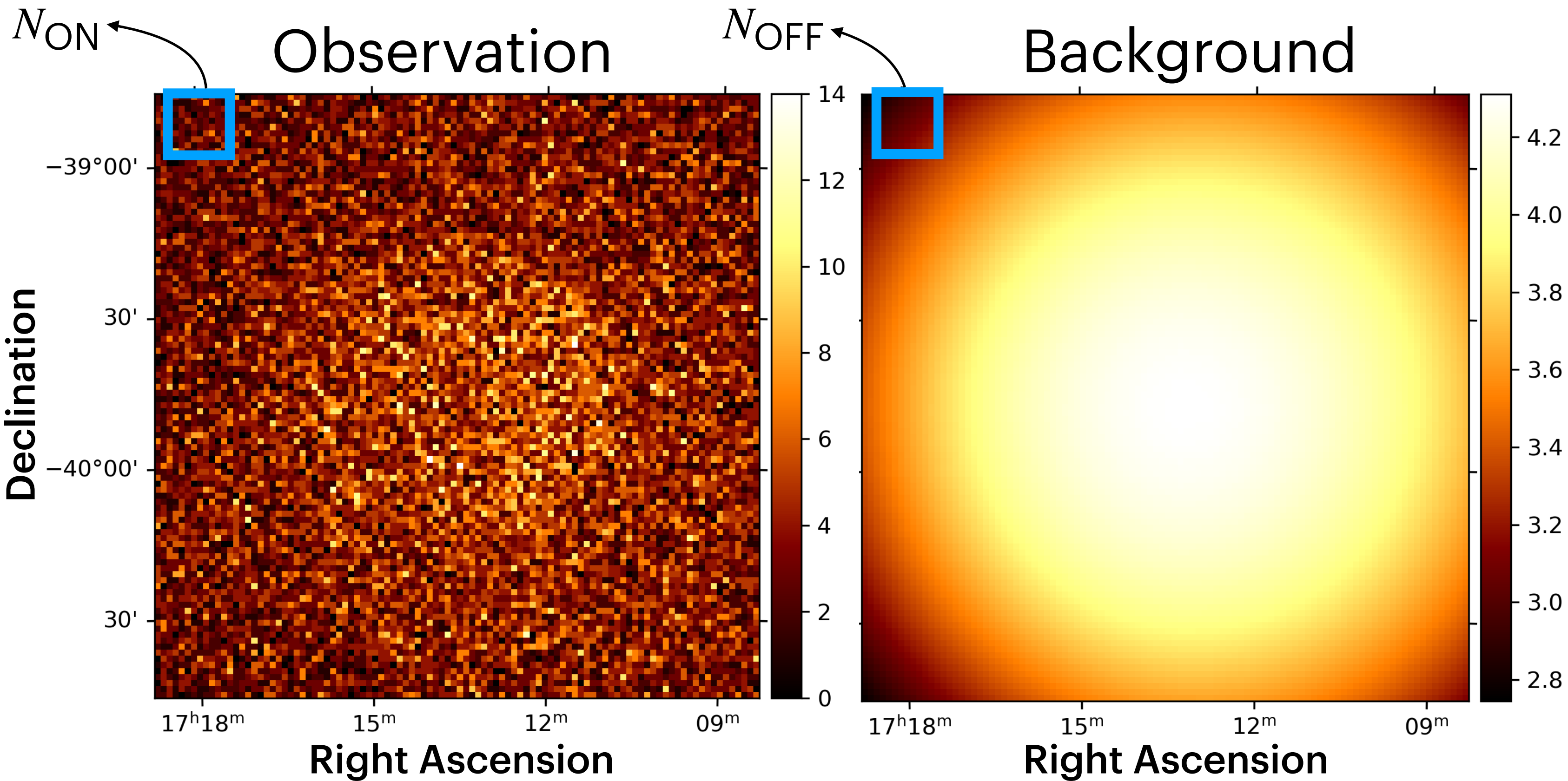
- 15 observations from 2004 released in HESS open data 2018





# Example sky maps: RX J1713.7 3946

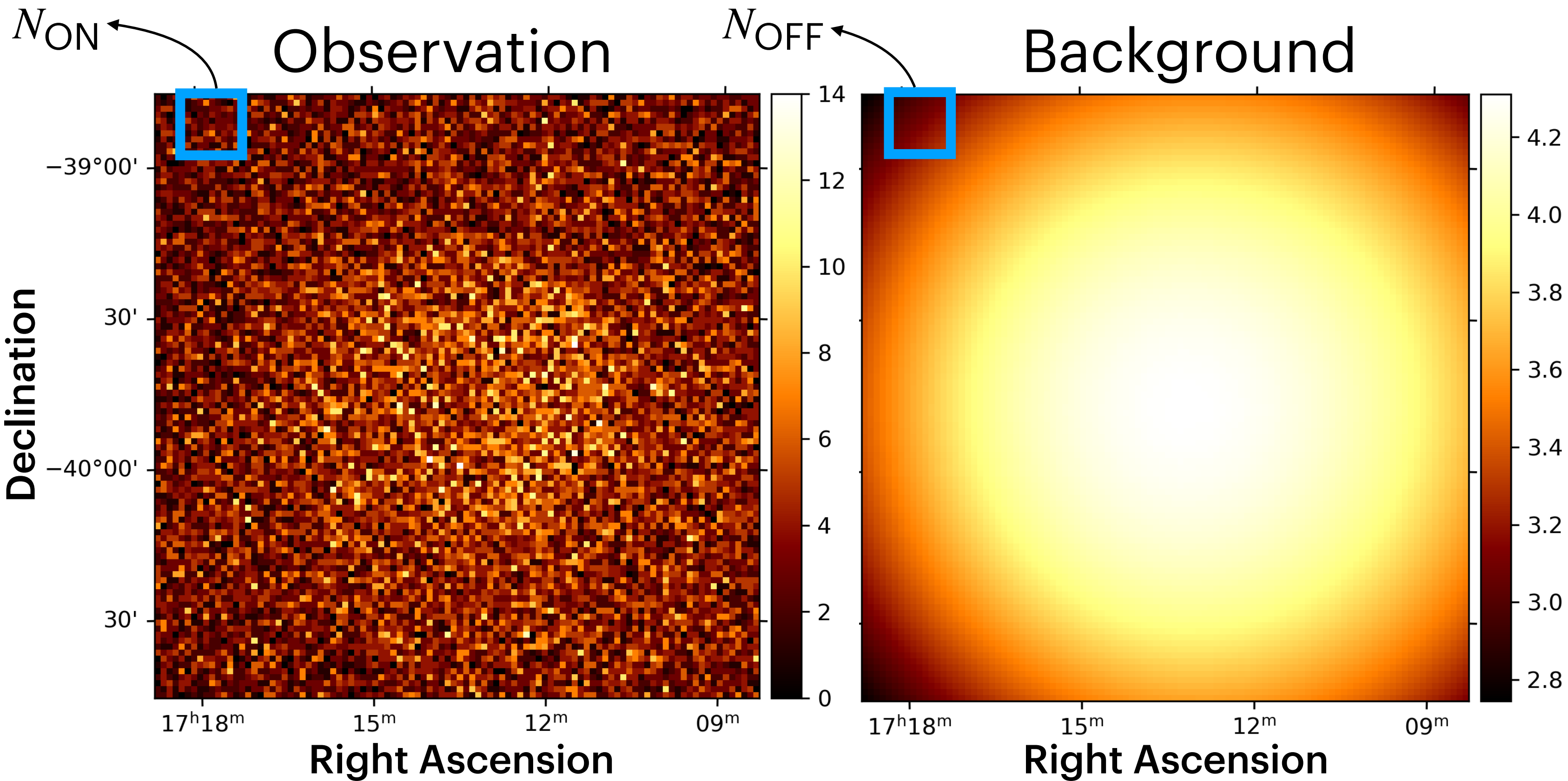
- 15 observations from 2004 released in HESS open data 2018





# Example sky maps: RX J1713.7 3946

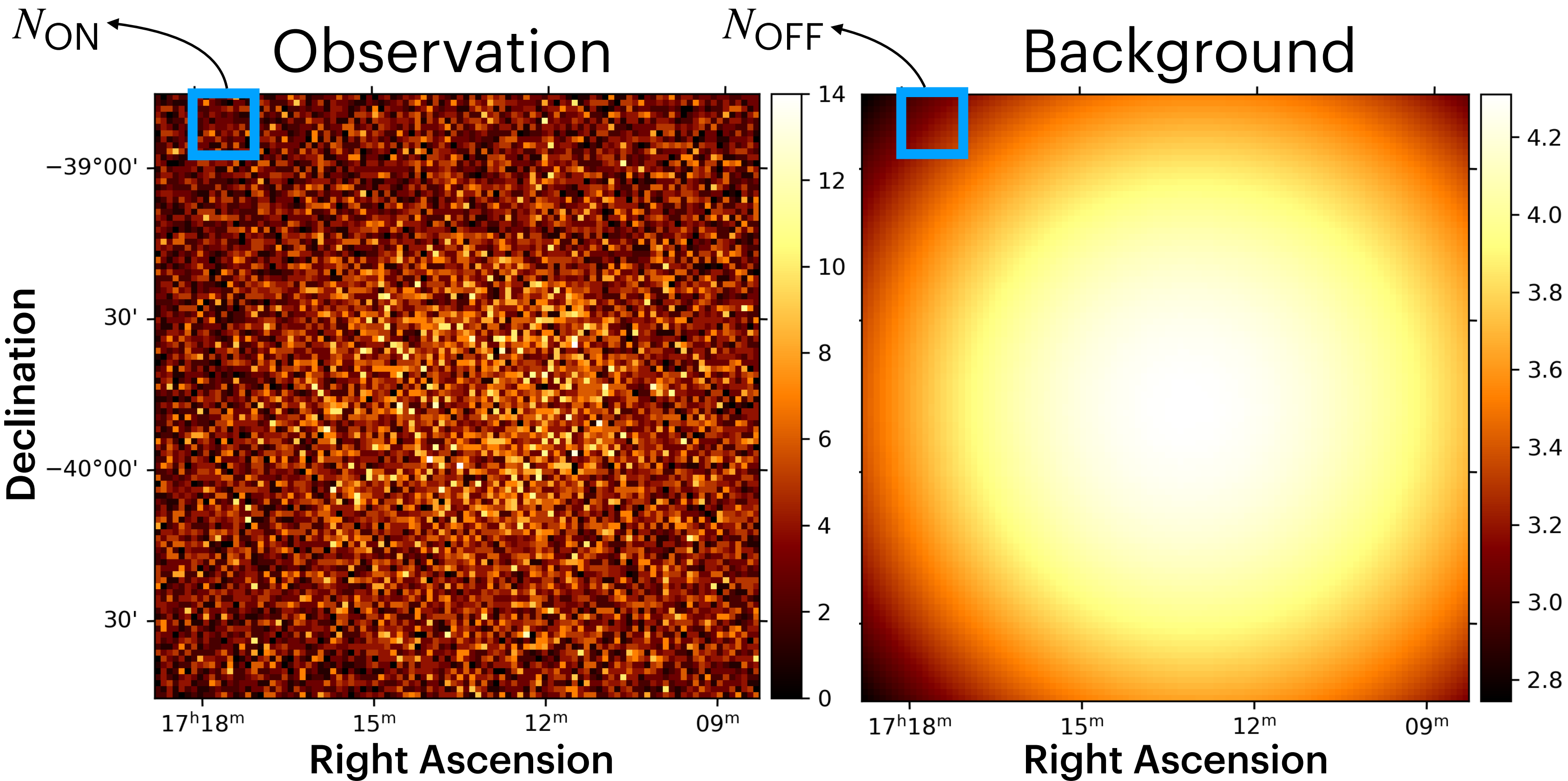
- 15 observations from 2004 released in HESS open data 2018





# Example sky maps: RX J1713.7 3946

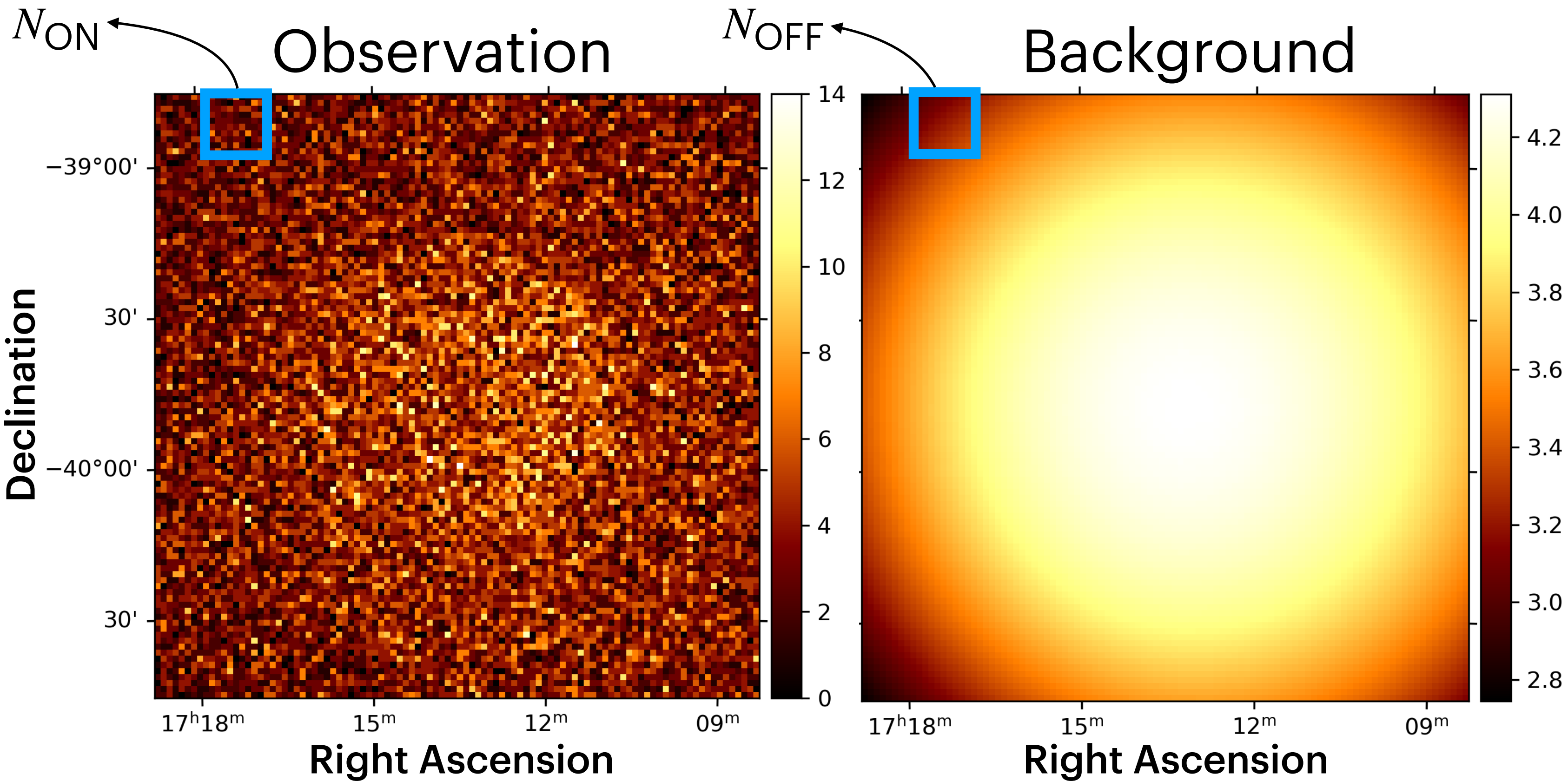
- 15 observations from 2004 released in HESS open data 2018





# Example sky maps: RX J1713.7 3946

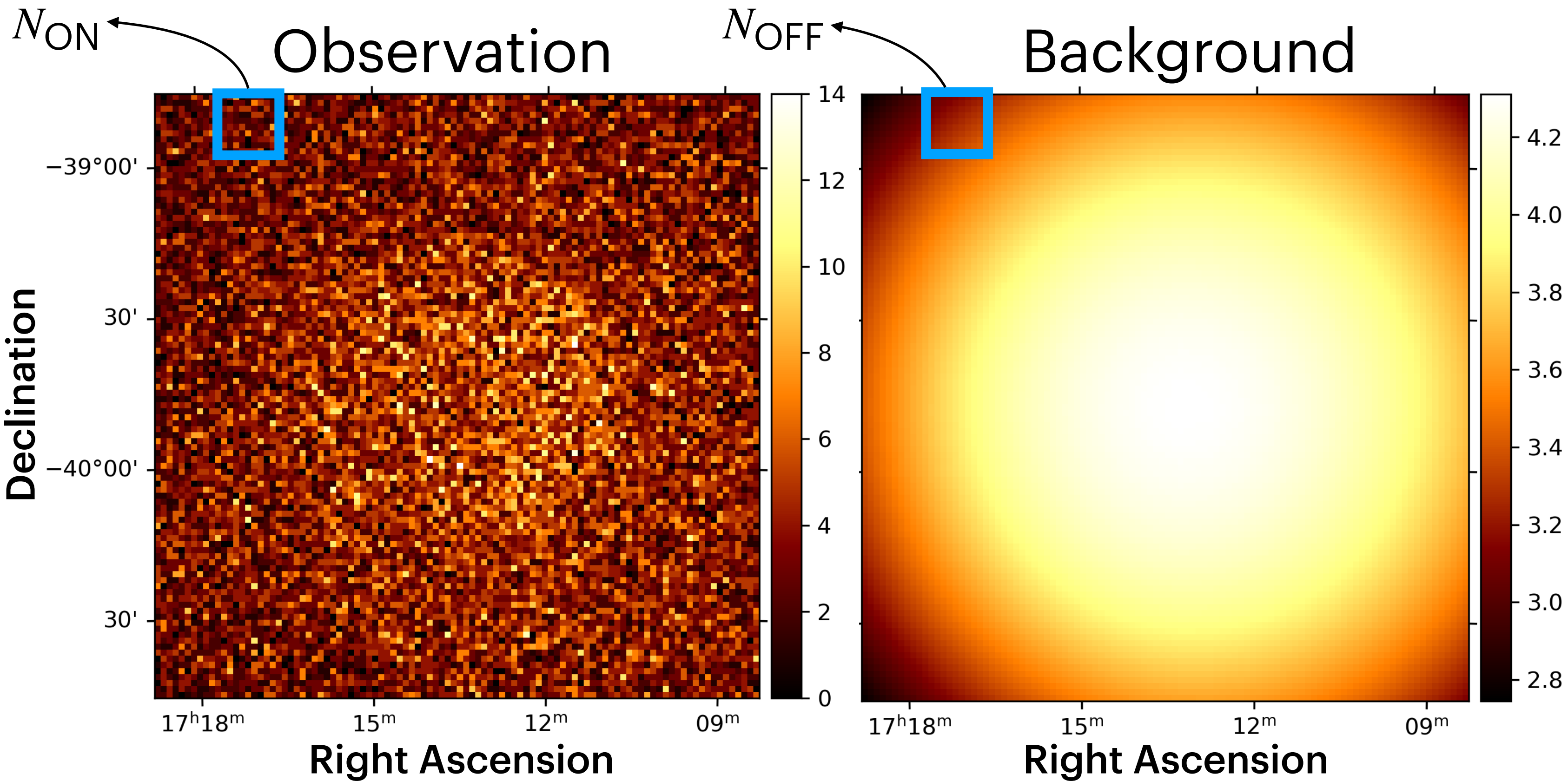
- 15 observations from 2004 released in HESS open data 2018





# Example sky maps: RX J1713.7 3946

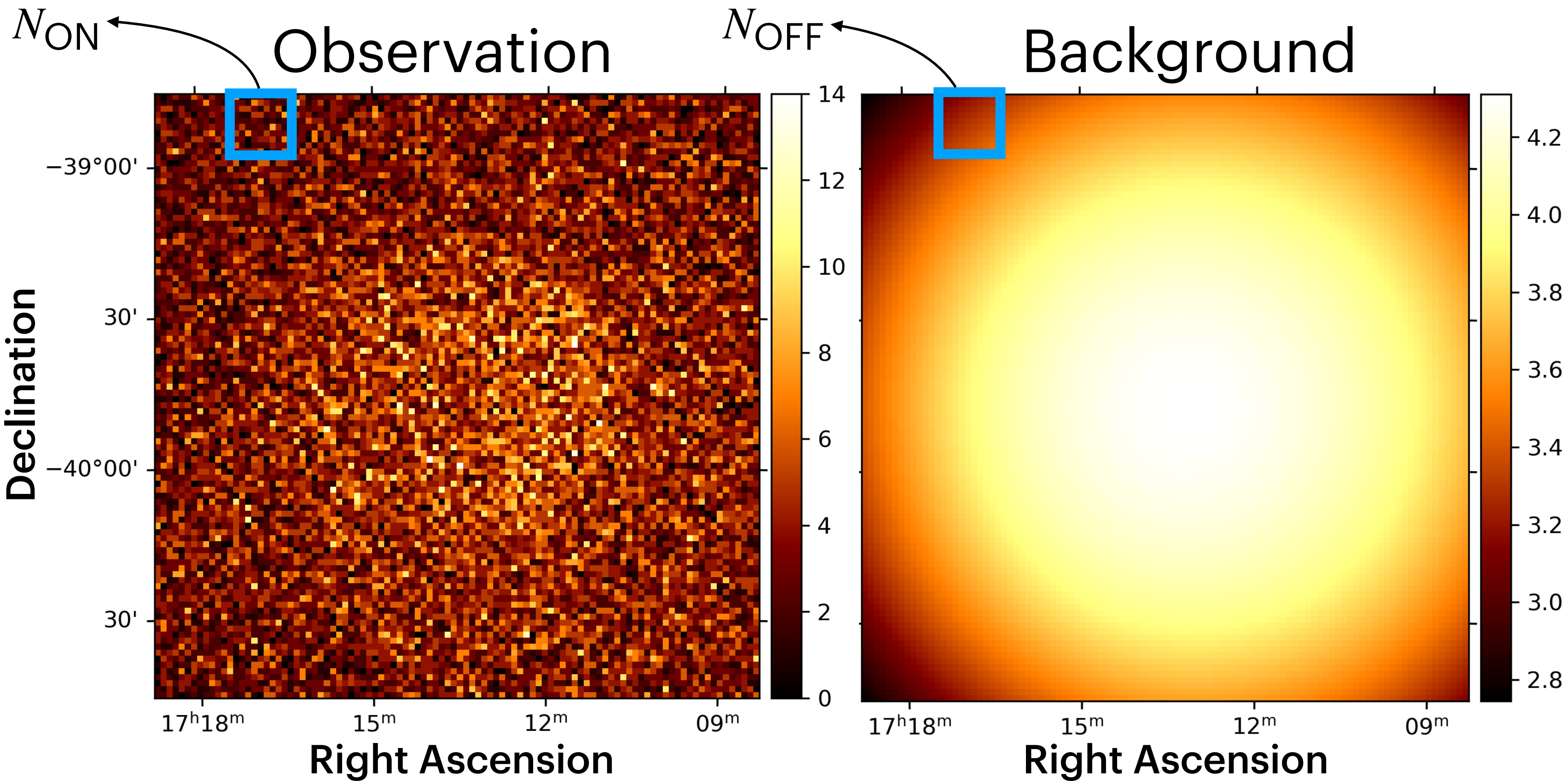
- 15 observations from 2004 released in HESS open data 2018





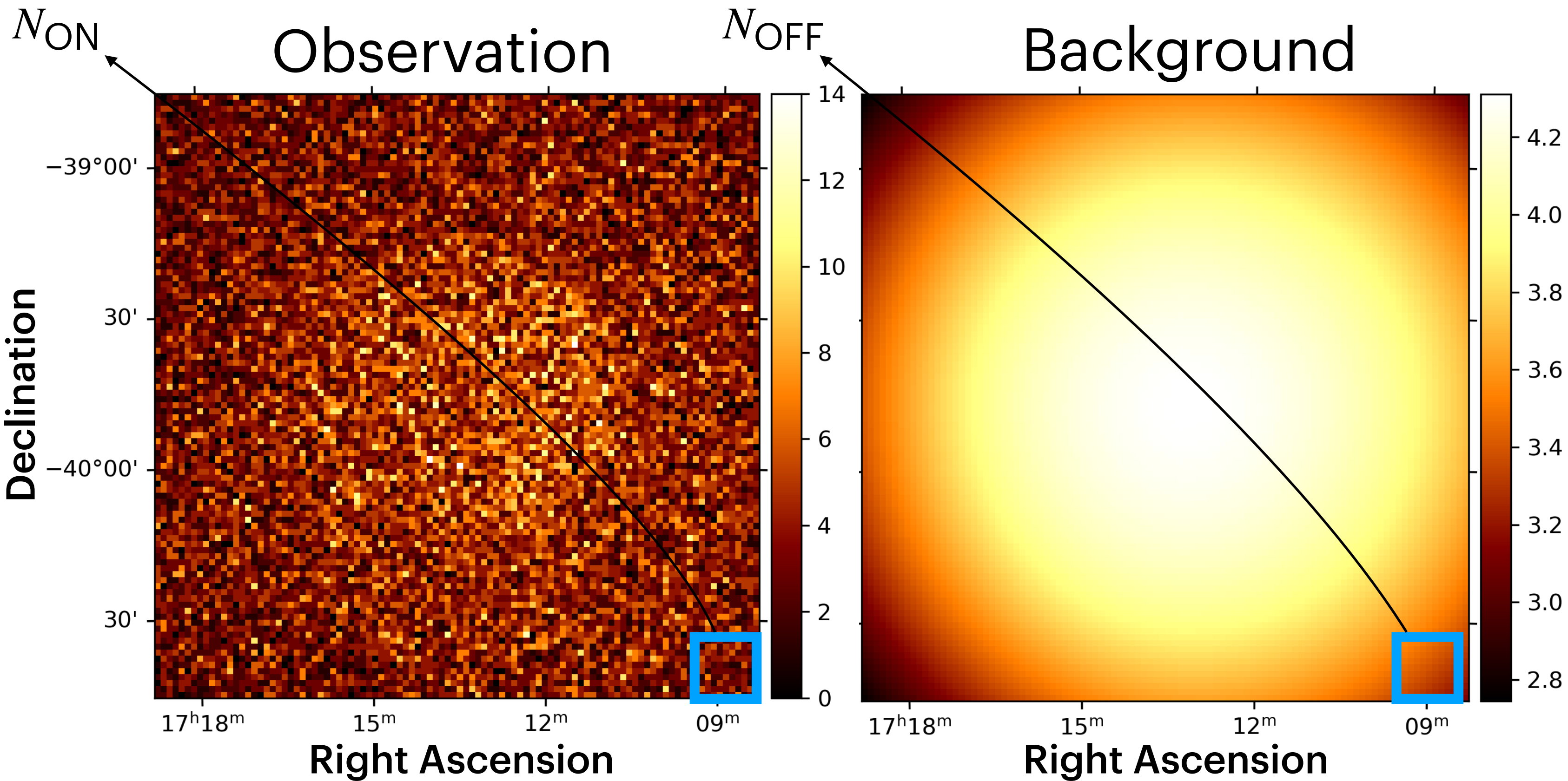
# Example sky maps: RX J1713.7 3946

- 15 observations from 2004 released in HESS open data 2018



# Example sky maps: RX J1713.7 3946

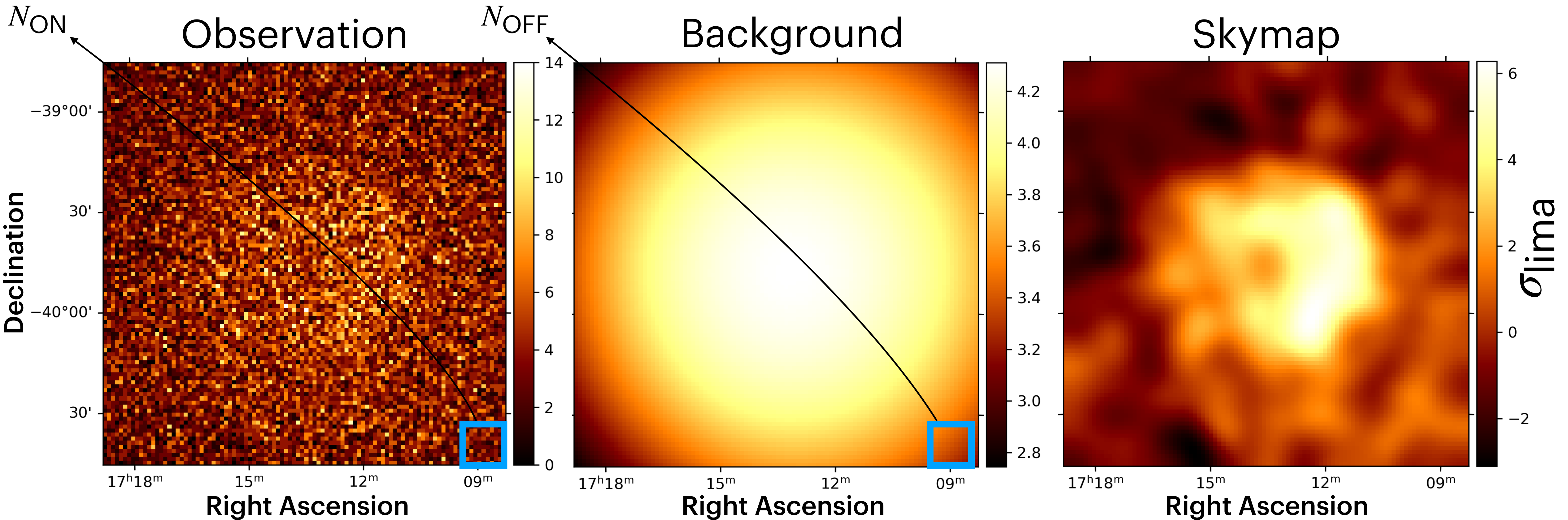
- 15 observations from 2004 released in HESS open data 2018





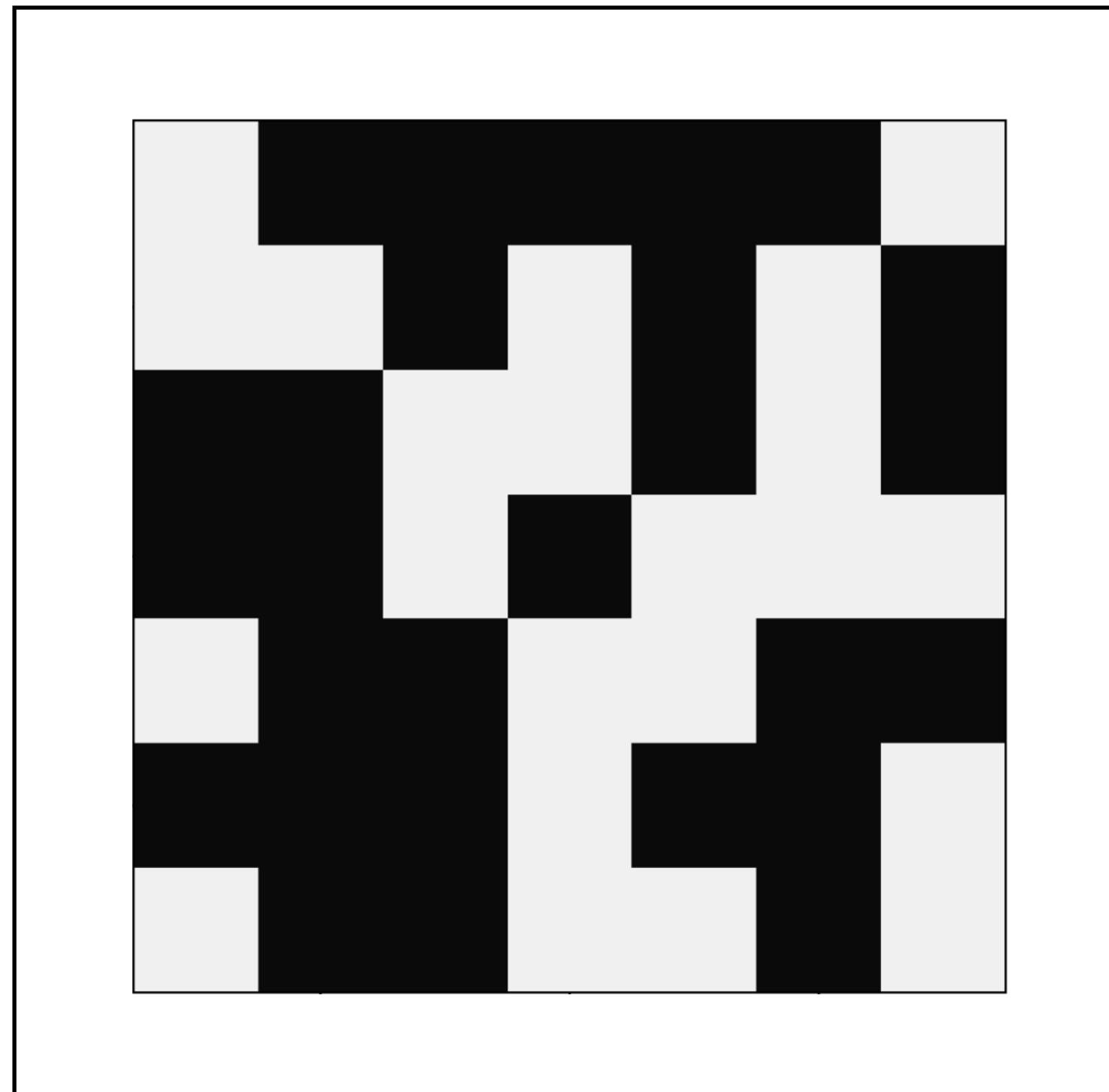
# Example sky maps: RX J1713.7 3946

- 15 observations from 2004 released in HESS open data 2018

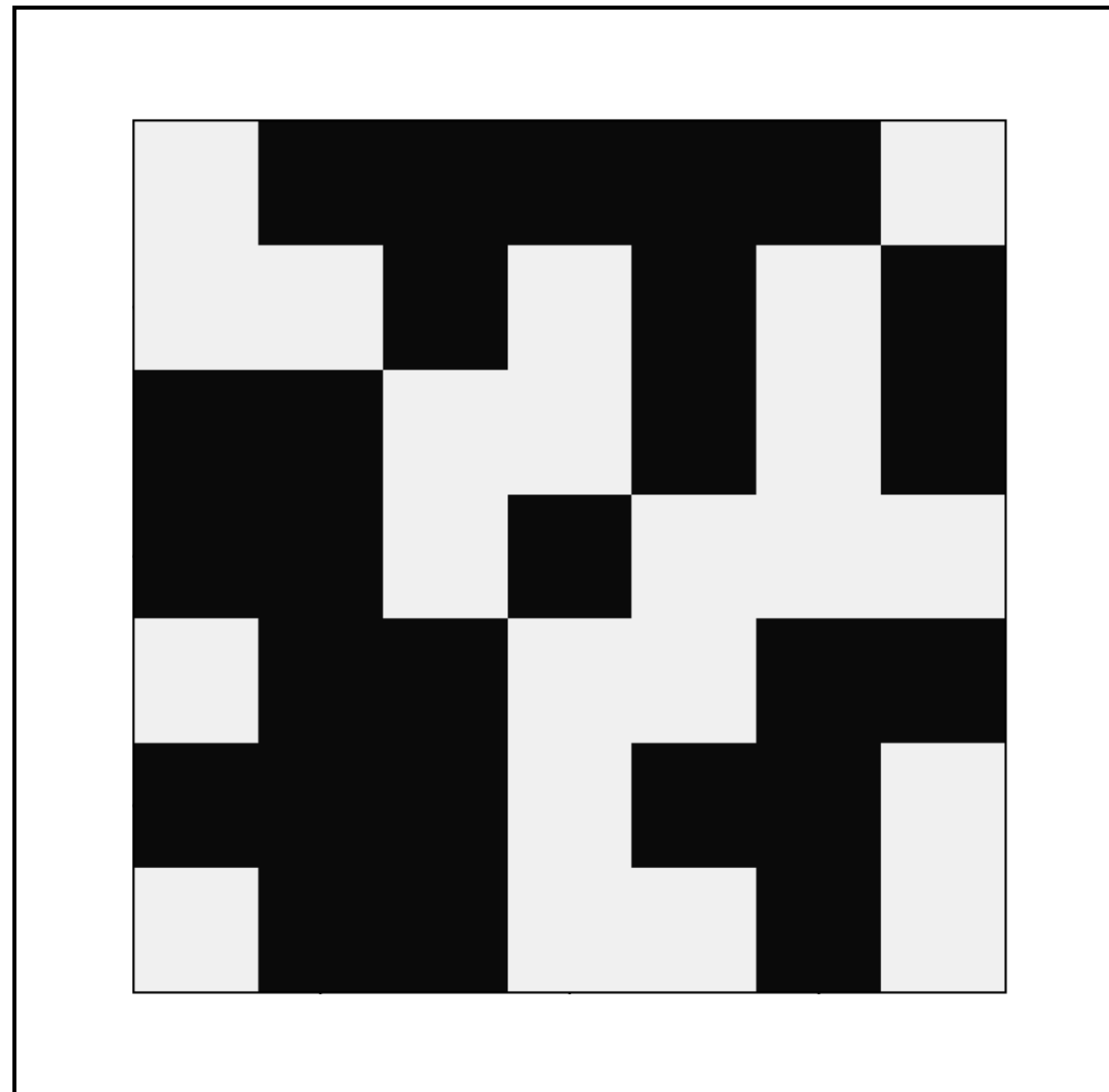


# Statistical physics





How can we quantify the structure of the black and white image?



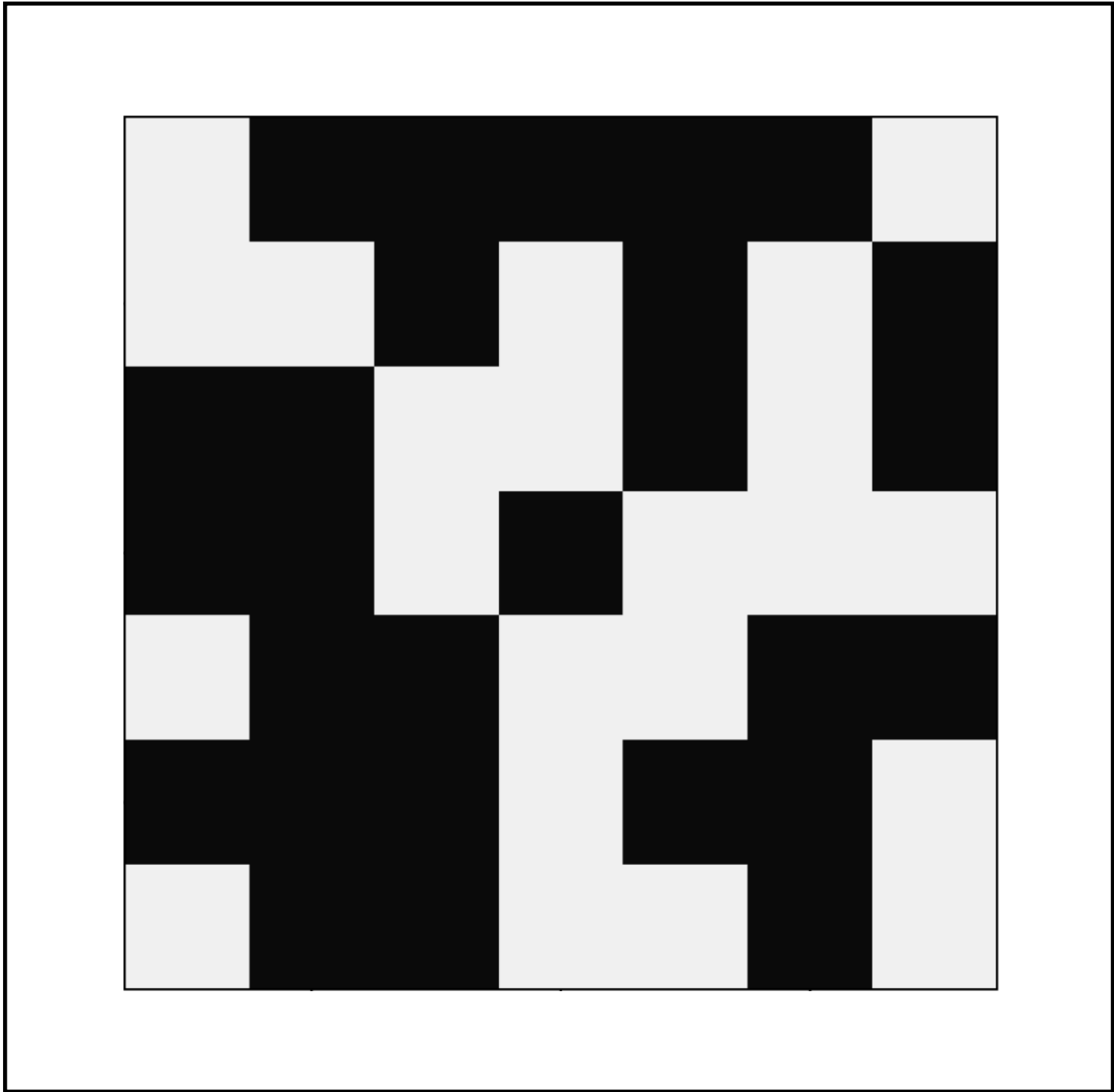
How can we quantify the structure of the black and white image?

One way: **Minkowski Functionals**



# Example

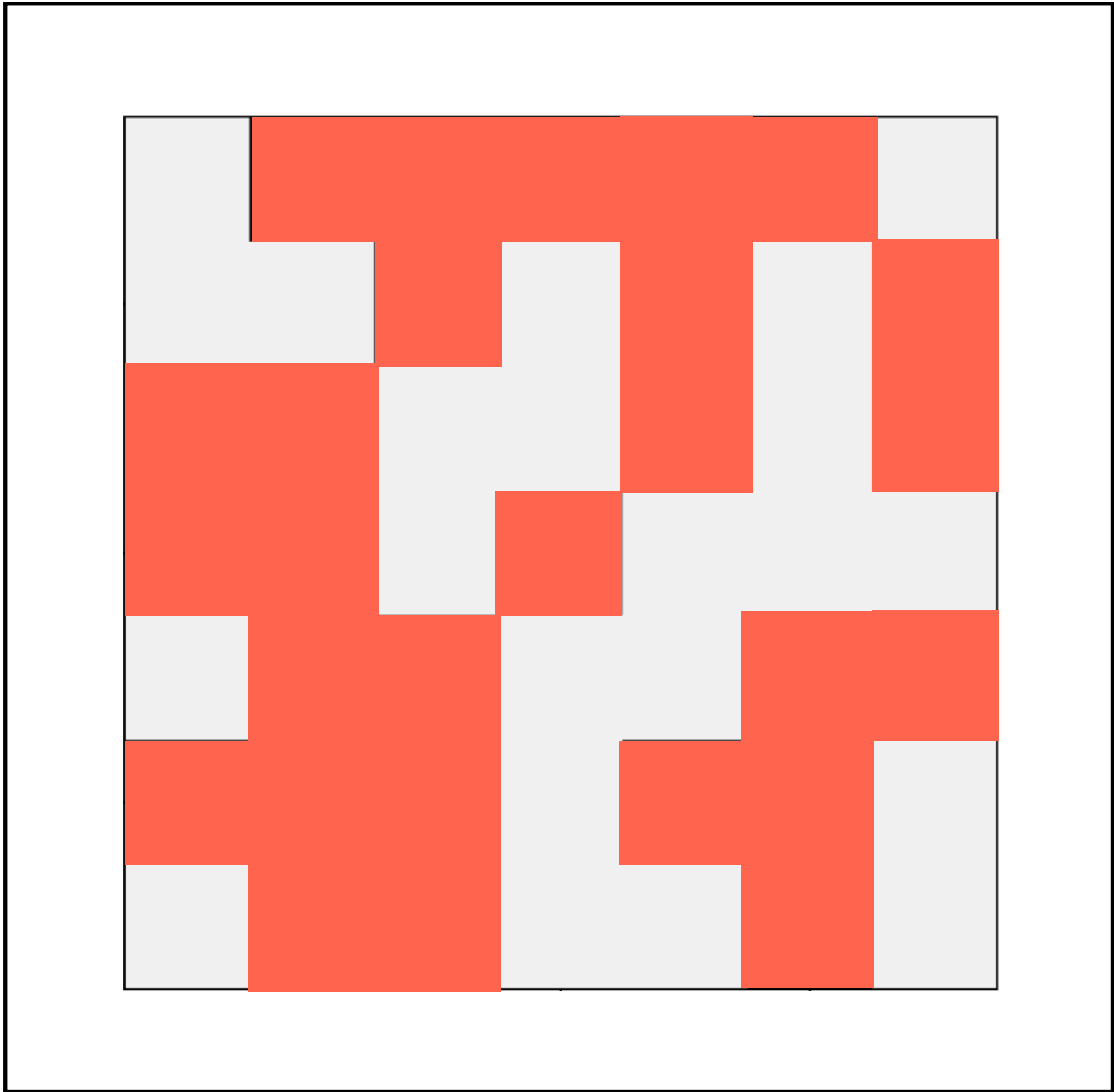
In two dimensions there are 3 such functionals. (fullfilling some desirable properties)



# Example

In two dimensions there are 3 such functionals. (fullfilling some desireable properties)

Area: 27



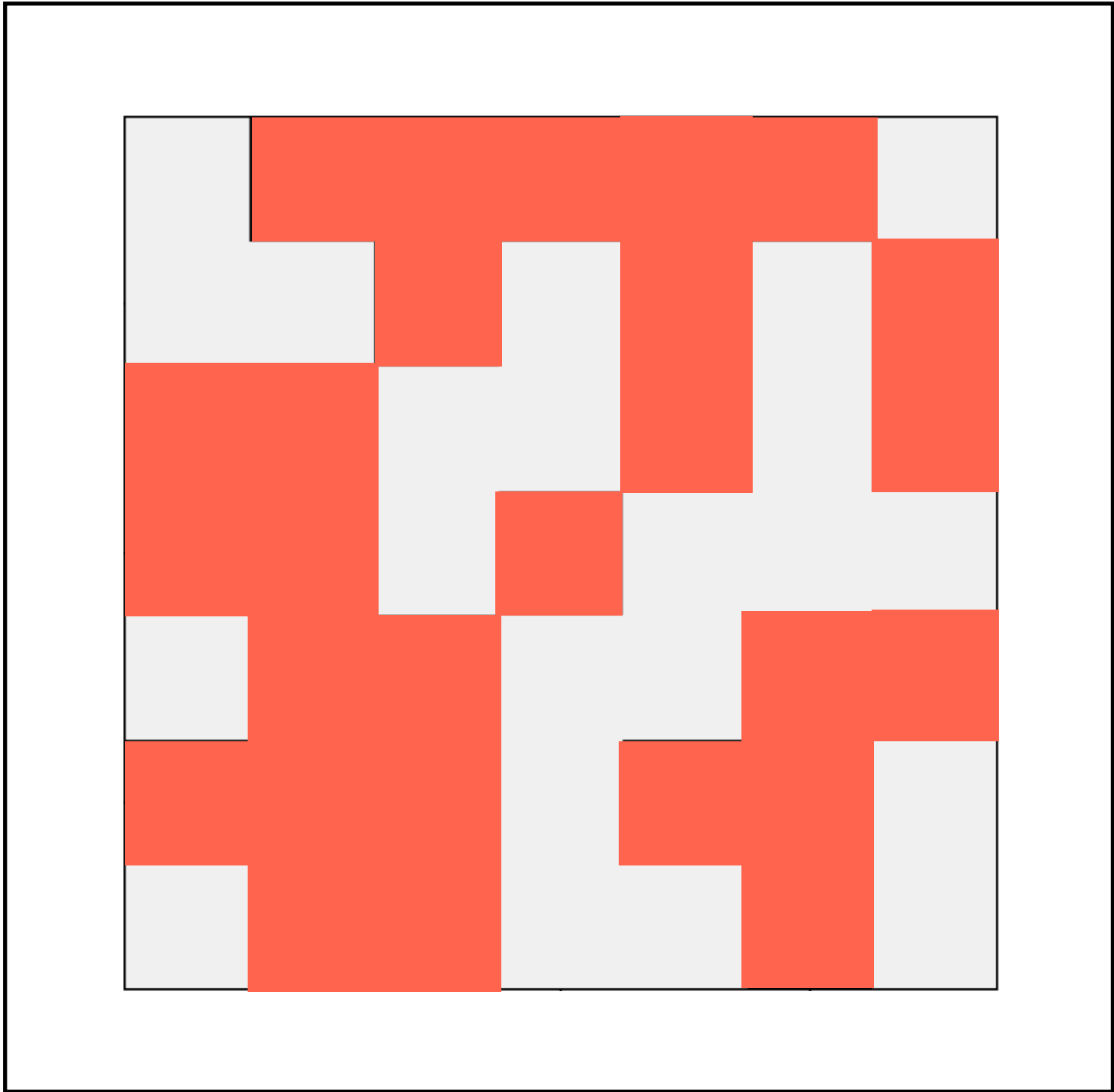


# Example

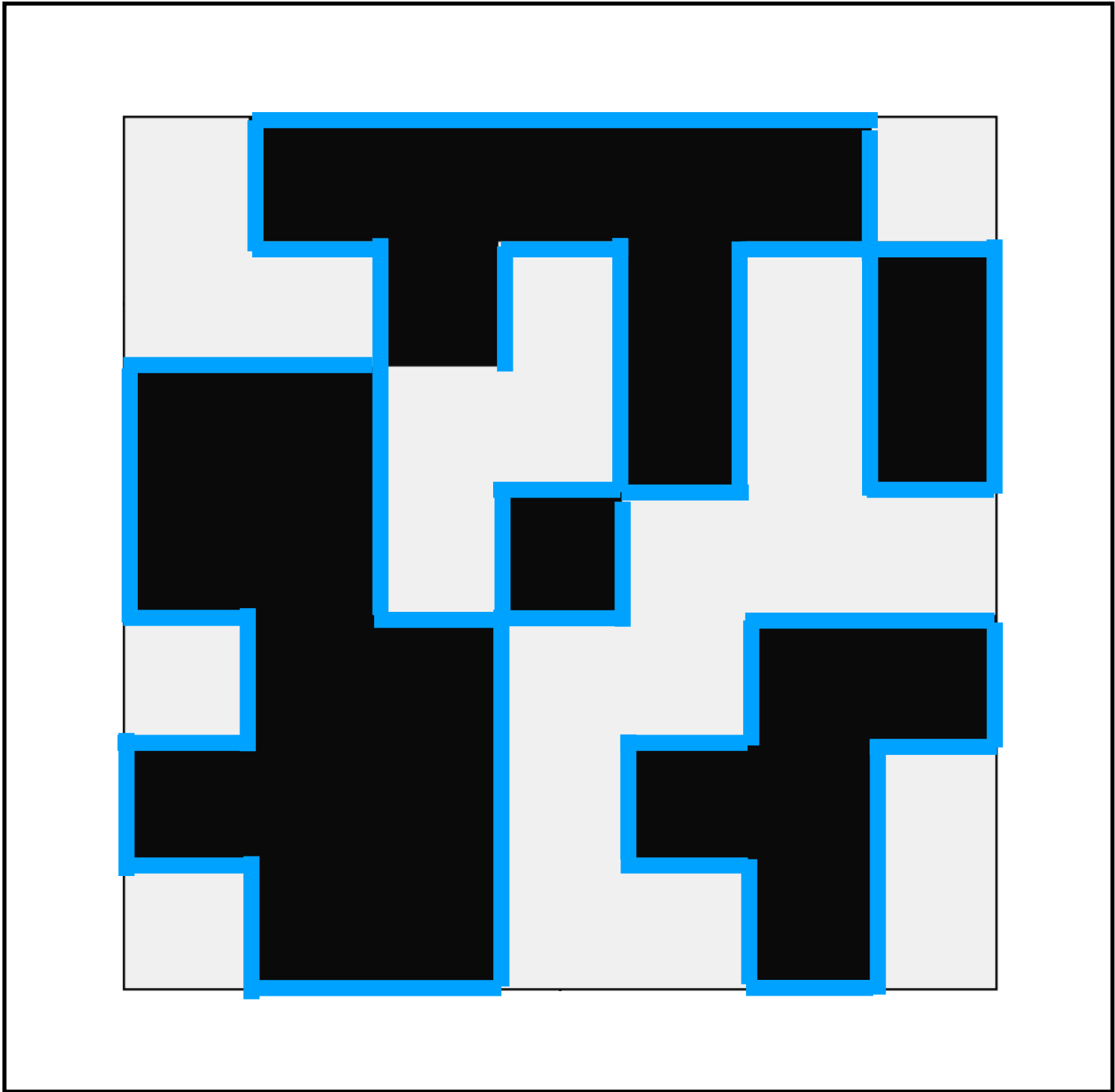
In two dimensions there are 3 such functionals. (fullfilling some desireable properties)



Area: 27



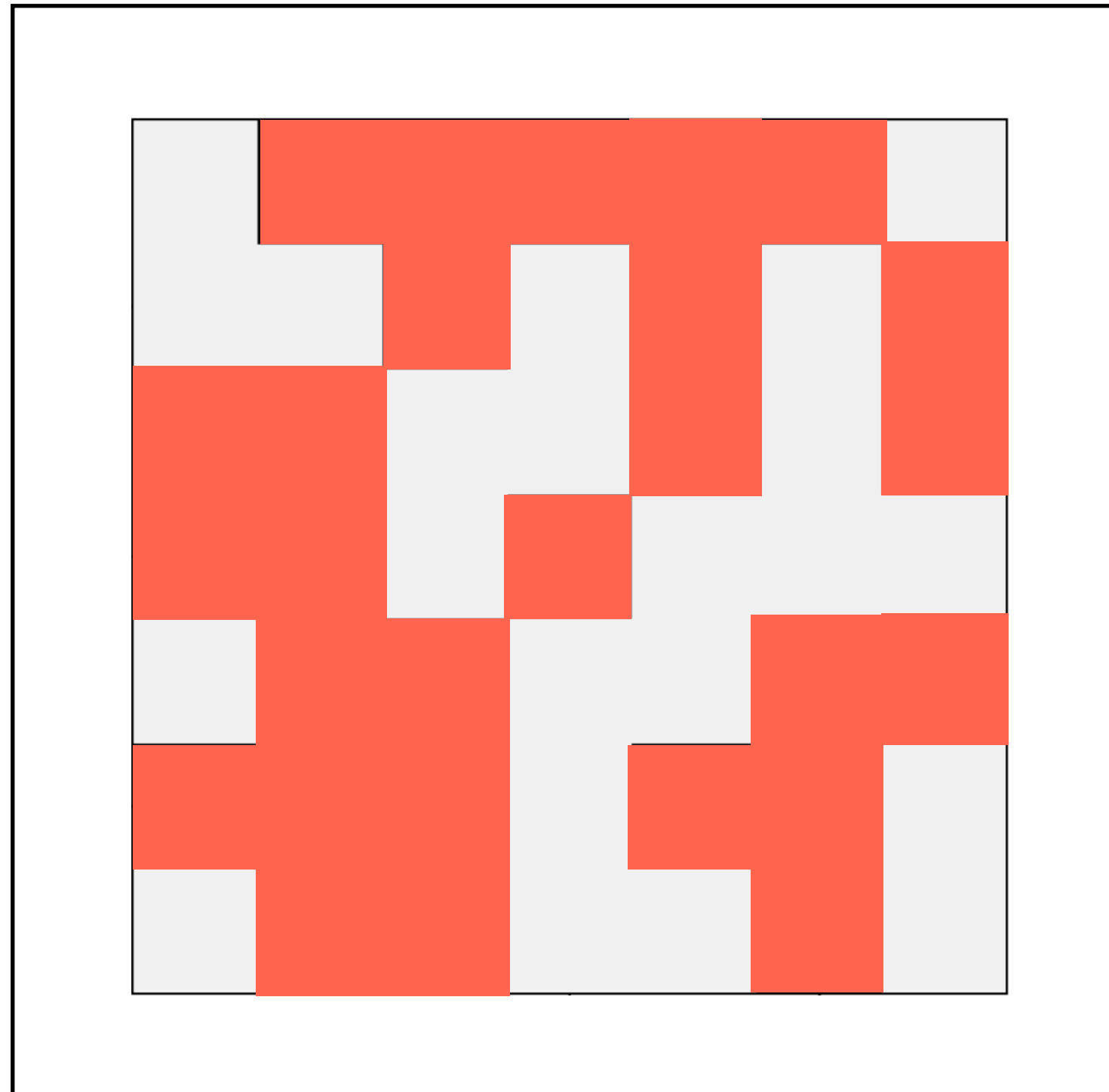
Perimeter: 58



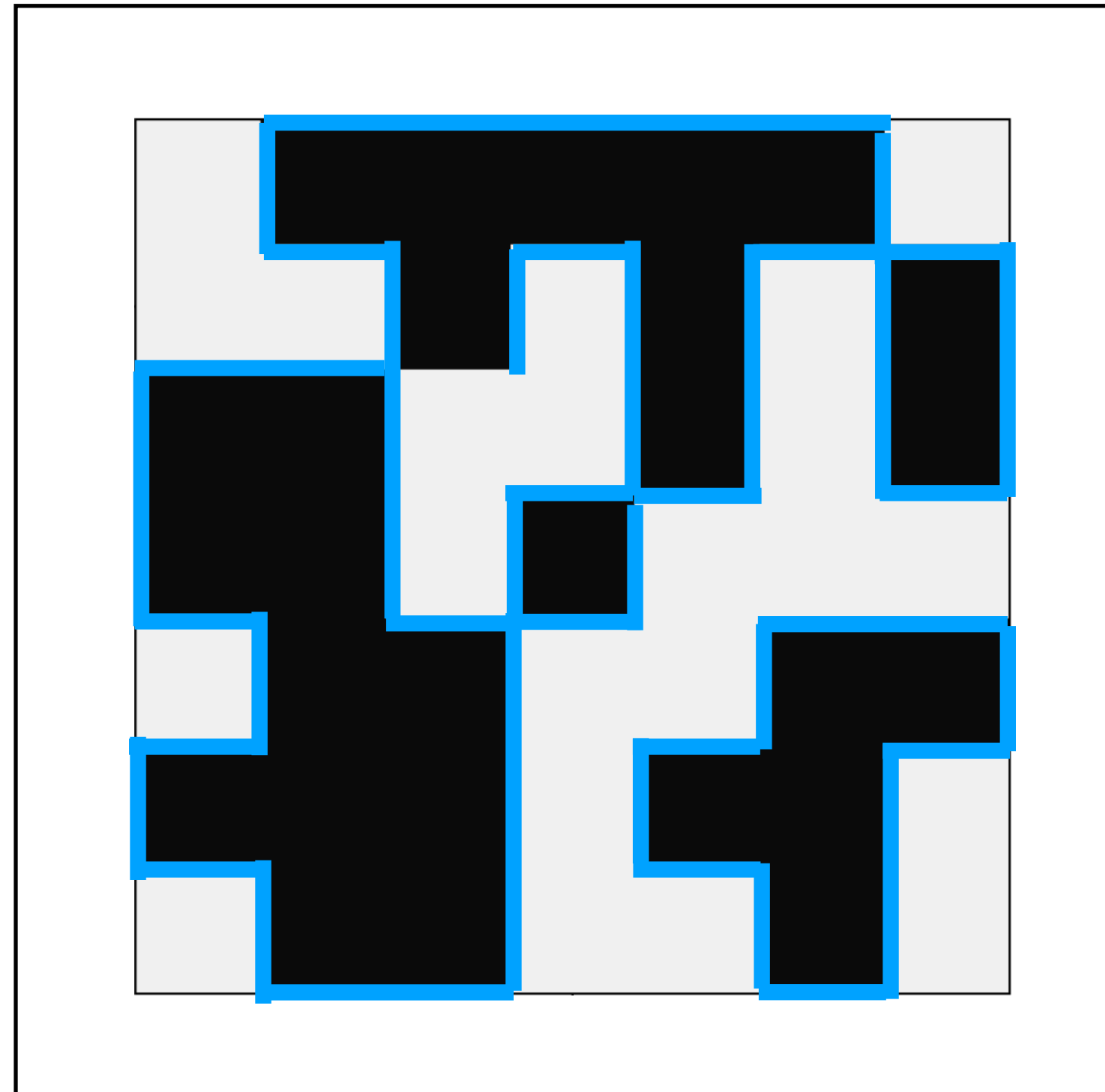
# Example

In two dimensions there are 3 such functionals. (fulfilling some desirable properties)

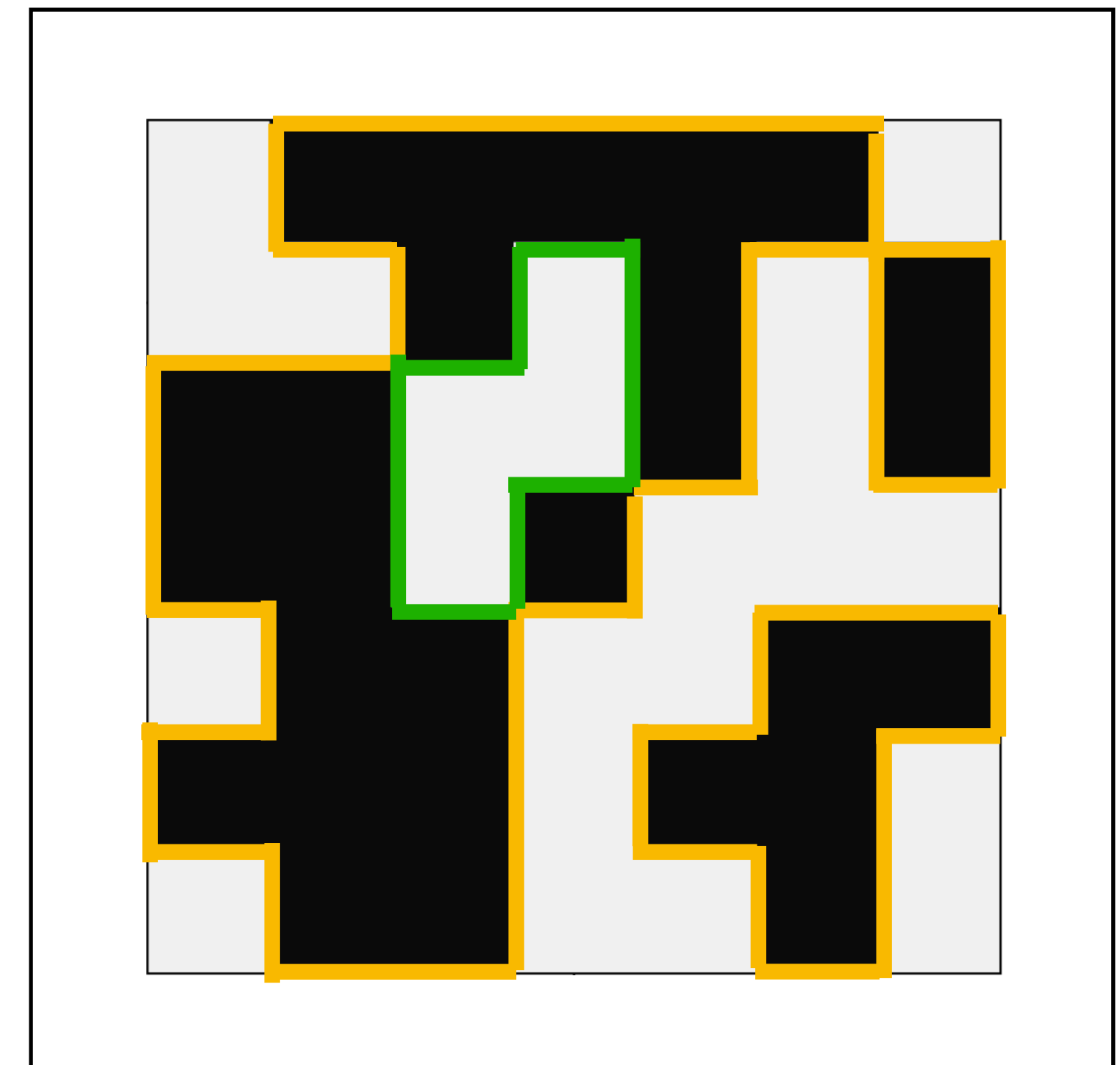
Area: 27



Perimeter: 58



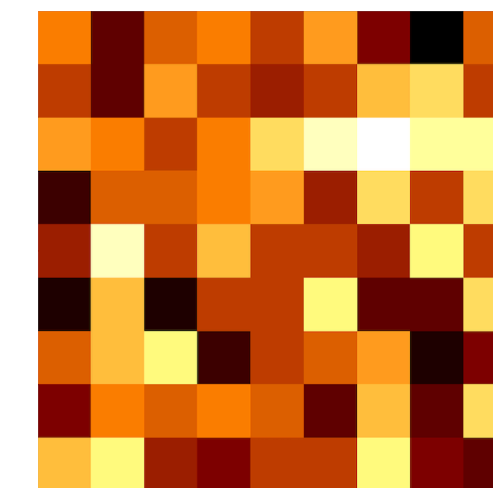
Euler Characteristic: 1





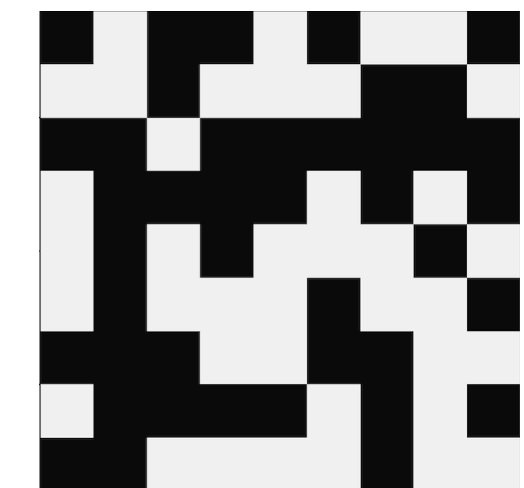
# Minkowski functionals and gamma-ray astronomy

- Cosmic ray events will be isotropically distributed in all directions → homogeneous Poisson background
- Generate black and white images by thresholding



$k \geq 10$

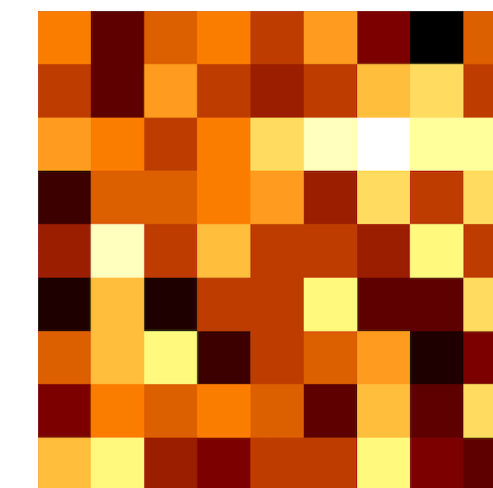
→



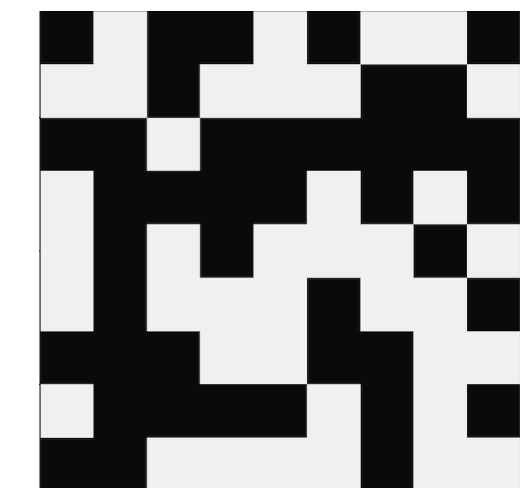
A: 42, P: 96,  $\chi$ : 1

# Minkowski functionals and gamma-ray astronomy

- Cosmic ray events will be isotropically distributed in all directions  $\rightarrow$  homogeneous Poisson background
- Generate black and white images by thresholding
- We can calculate all possible black and white images for a certain system size  $N \rightarrow$  distributions for Minkowski functionals

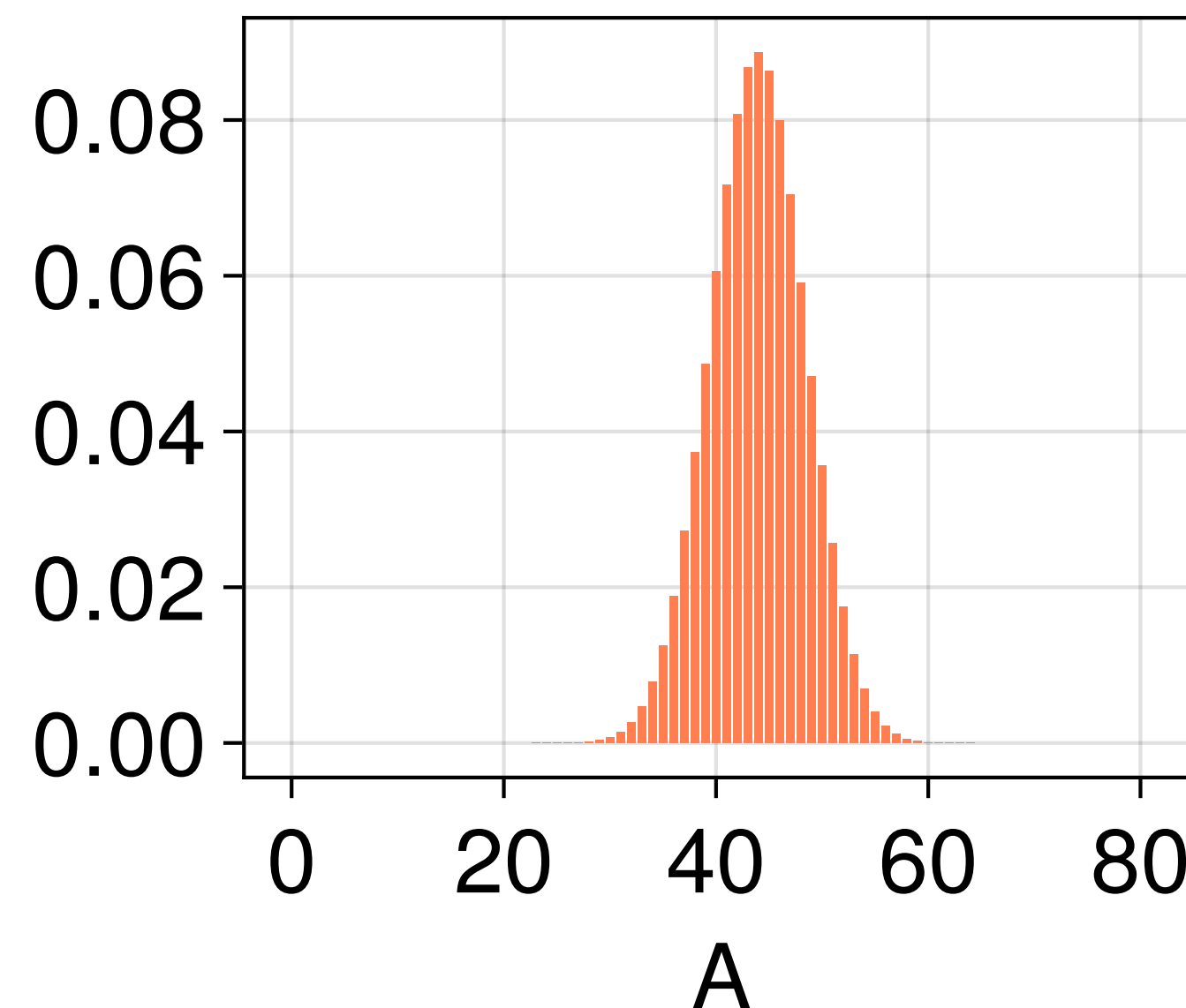


$k \geq 10$

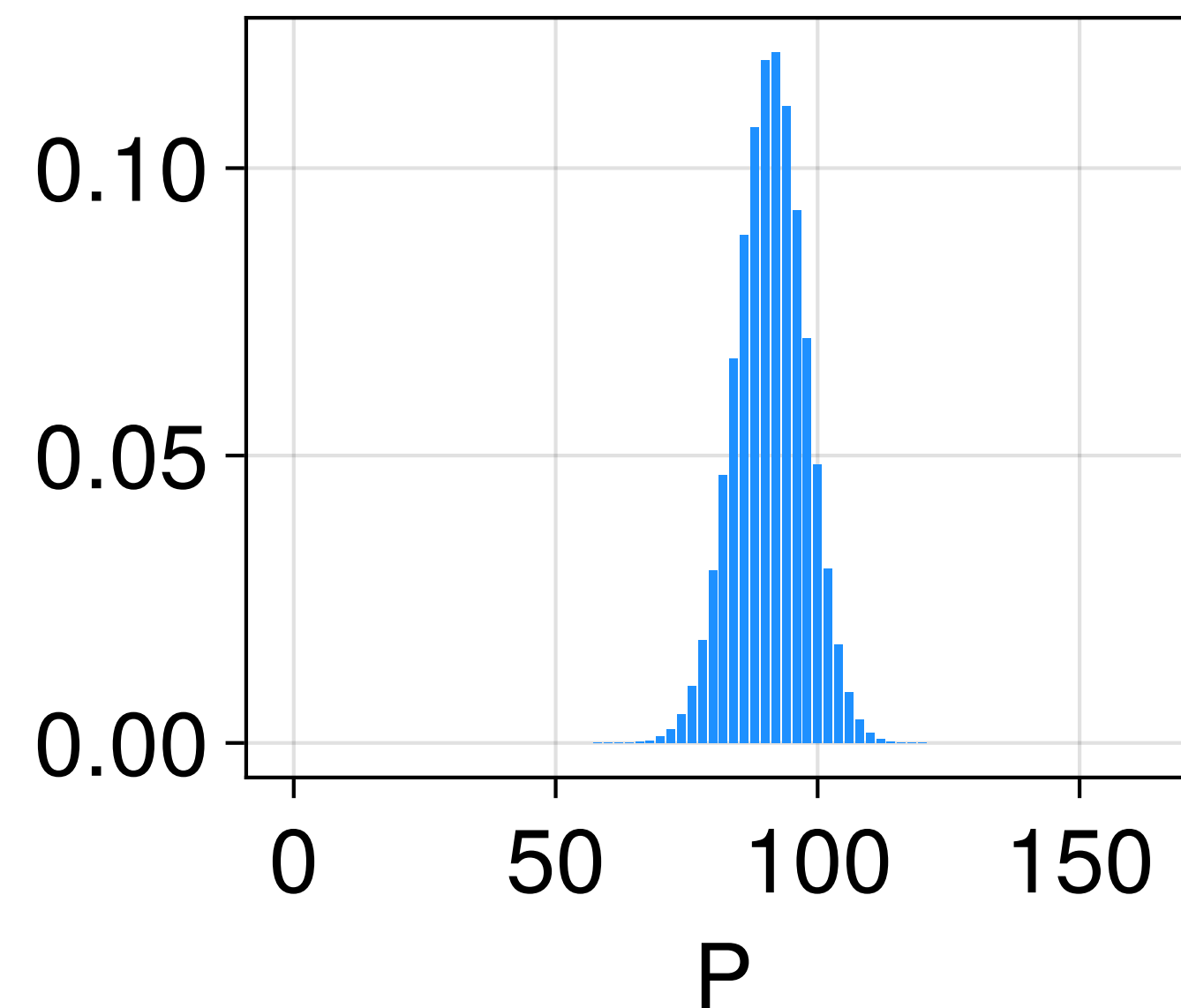


A: 42, P: 96,  $\chi$ : 1

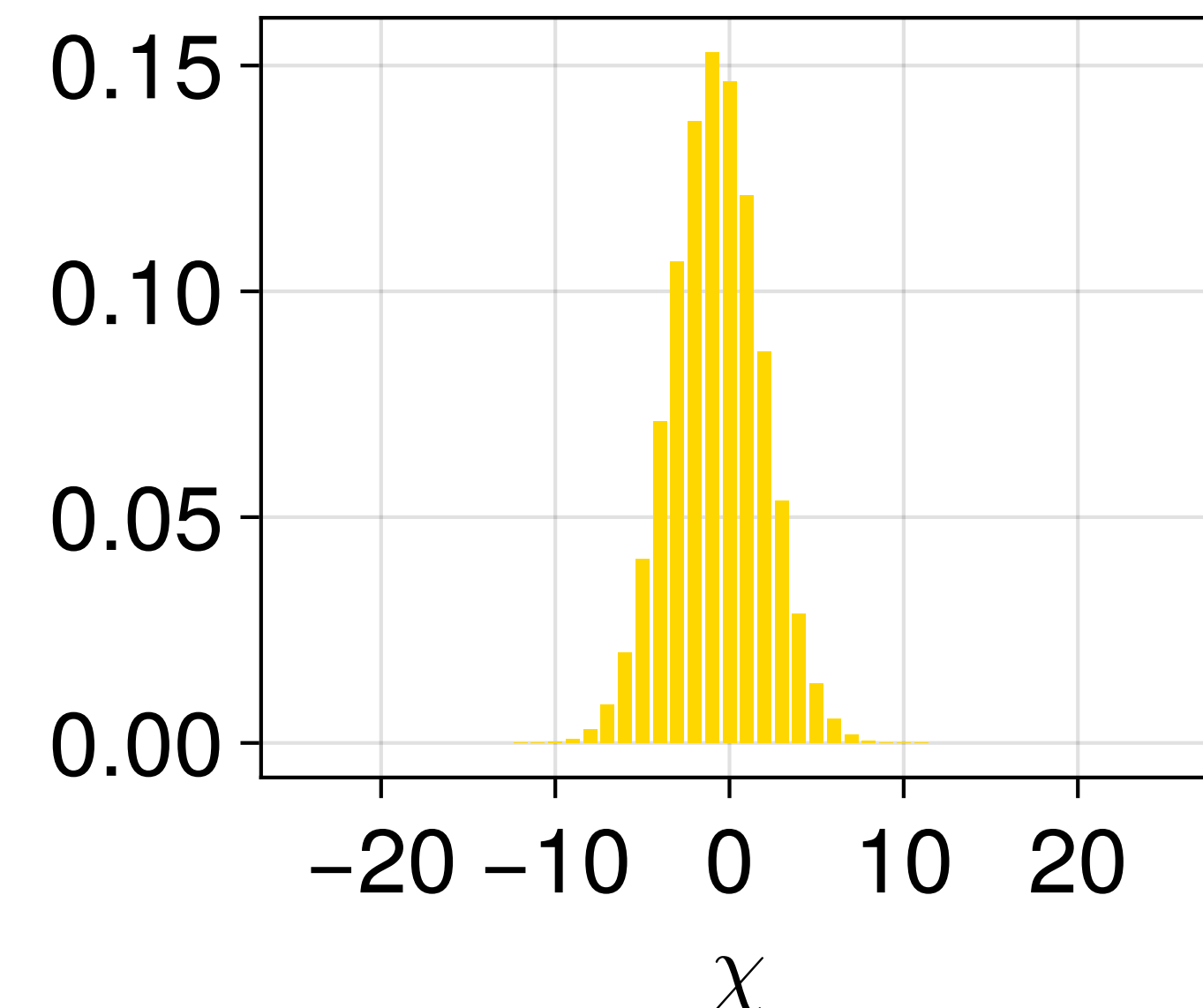
**Area**



**Perimeter**



**Euler Characteristic**



# Minkowski functionals and gamma-ray astronomy



Idea: Generate, from a counts map, a series of black and white images by thresholding. (Set pixel black if  $>$  threshold otherwise white)

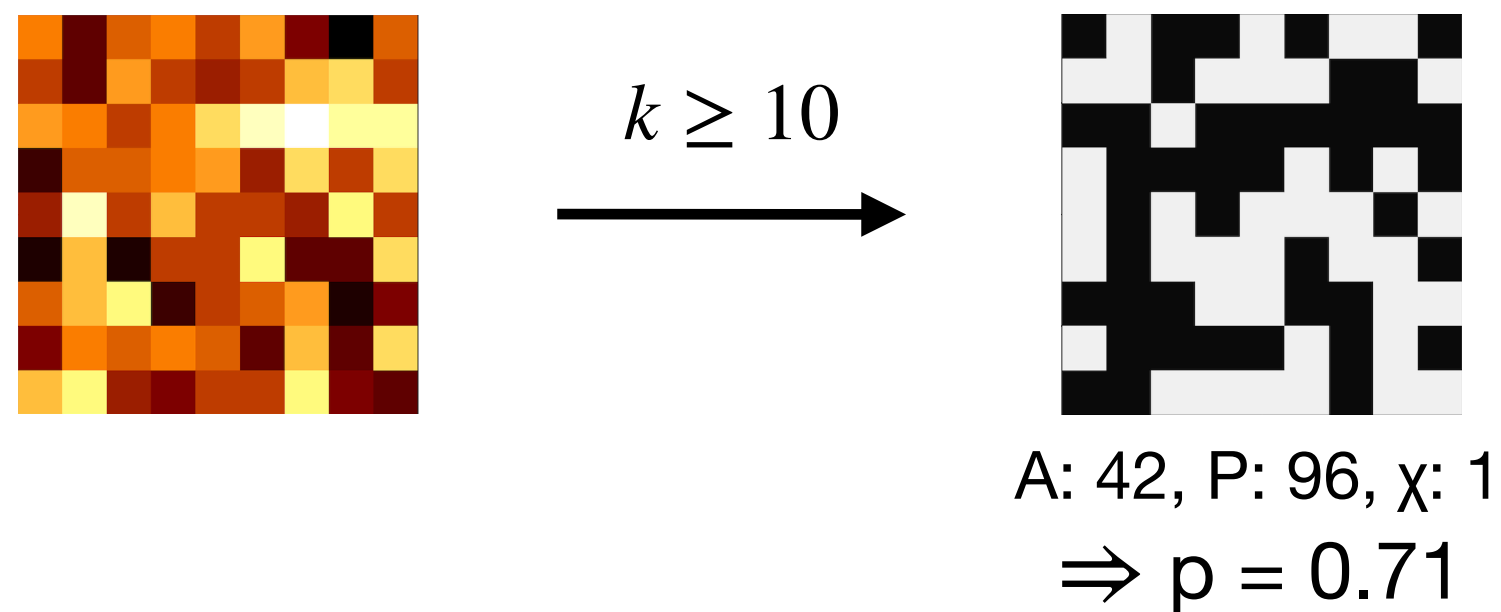
Hypothesis test for each black and white image  $\rightarrow$  **Null hypothesis: Observed black and white image is from a Poisson background with mean  $\lambda$ , at threshold X.**



# Minkowski functionals and gamma-ray astronomy

Idea: Generate, from a counts map, a series of black and white images by thresholding. (Set pixel black if  $>$  threshold otherwise white)

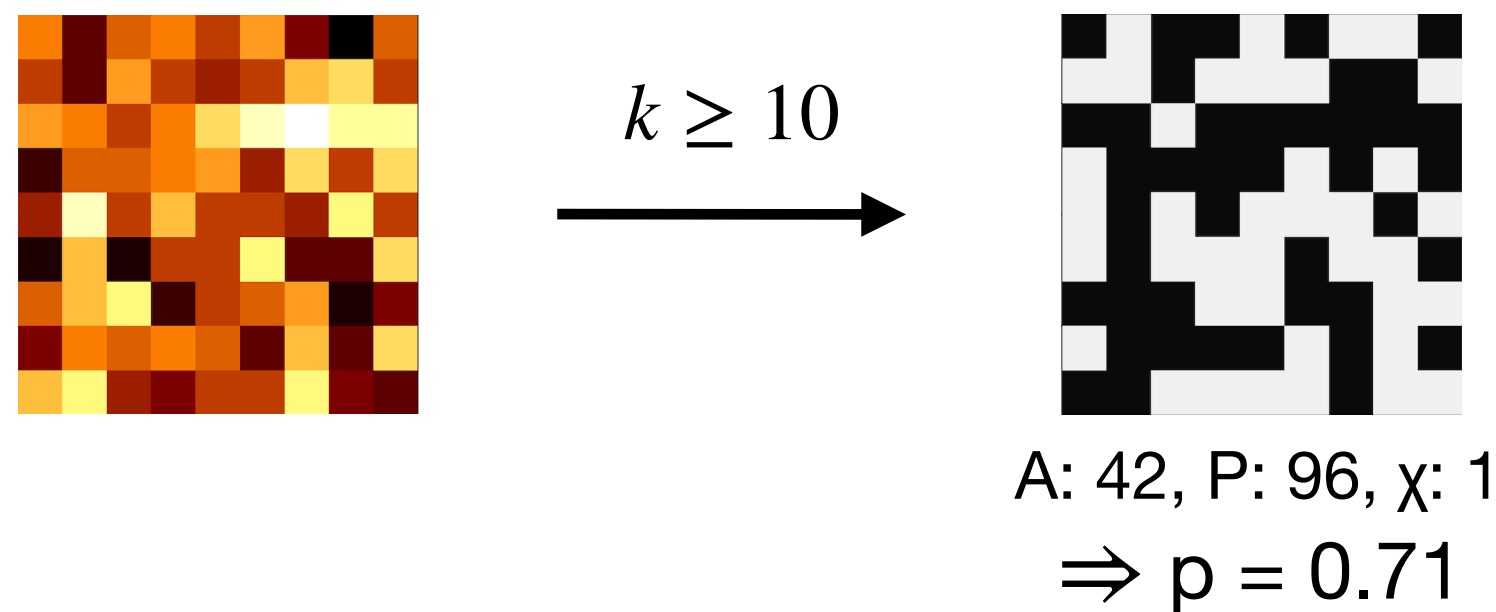
Hypothesis test for each black and white image  $\rightarrow$  **Null hypothesis: Observed black and white image is from a Poisson background with mean  $\lambda$ , at threshold X.**



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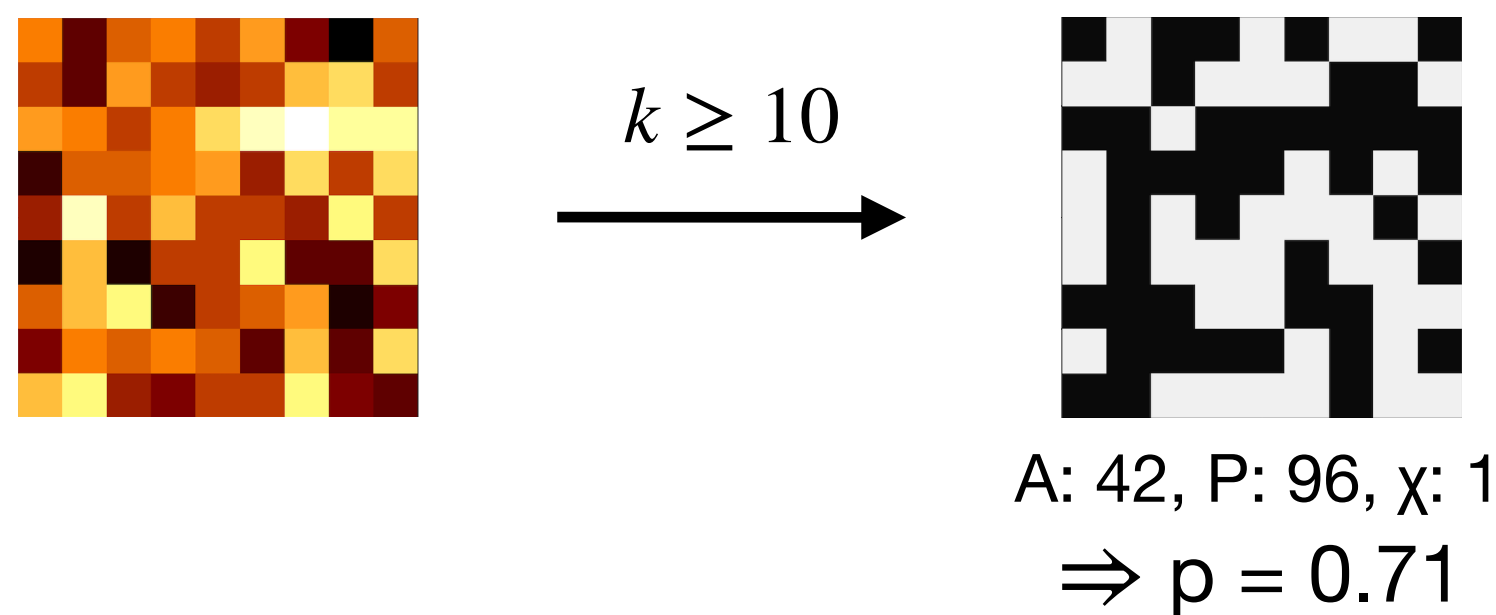
$P = 0.71$  in words: Given the ***hypothesis is true*** there is a chance of 71% that this black and white image (at threshold 10) or an even more unlikely one appears.

# Minkowski functionals and gamma-ray astronomy

Idea: Generate, from a counts map, a series of black and white images by thresholding. (Set pixel black if  $>$  threshold otherwise white)

Hypothesis test for each black and white image  $\rightarrow$  **Null hypothesis: Observed black and white image is from a Poisson background with mean  $\lambda$ , at threshold X.**

For all possible thresholds:



$P = 0.71$  in words: Given the ***hypothesis is true*** there is a chance of 71% that this black and white image (at threshold 10) or an even more unlikely one appears.

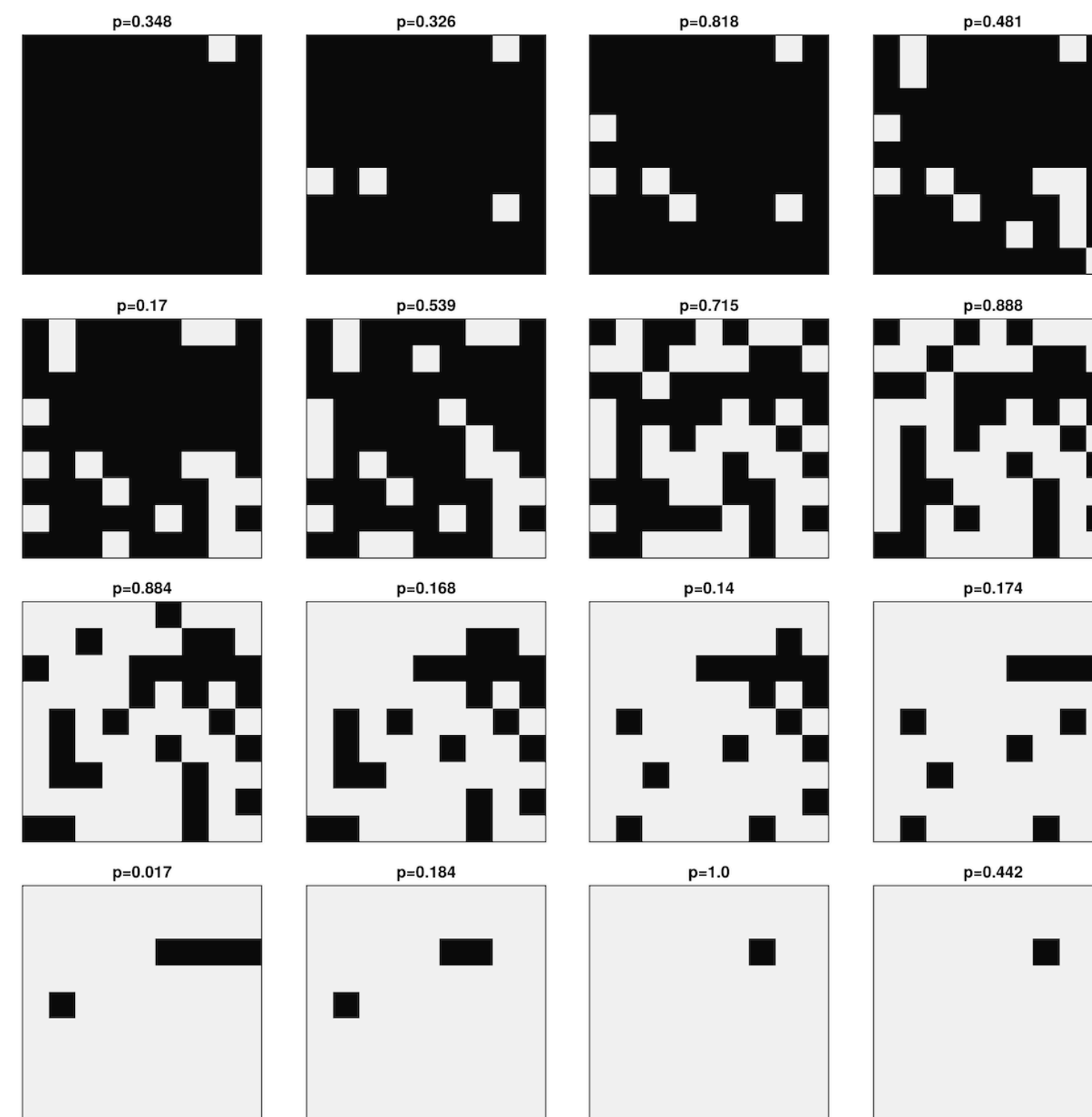
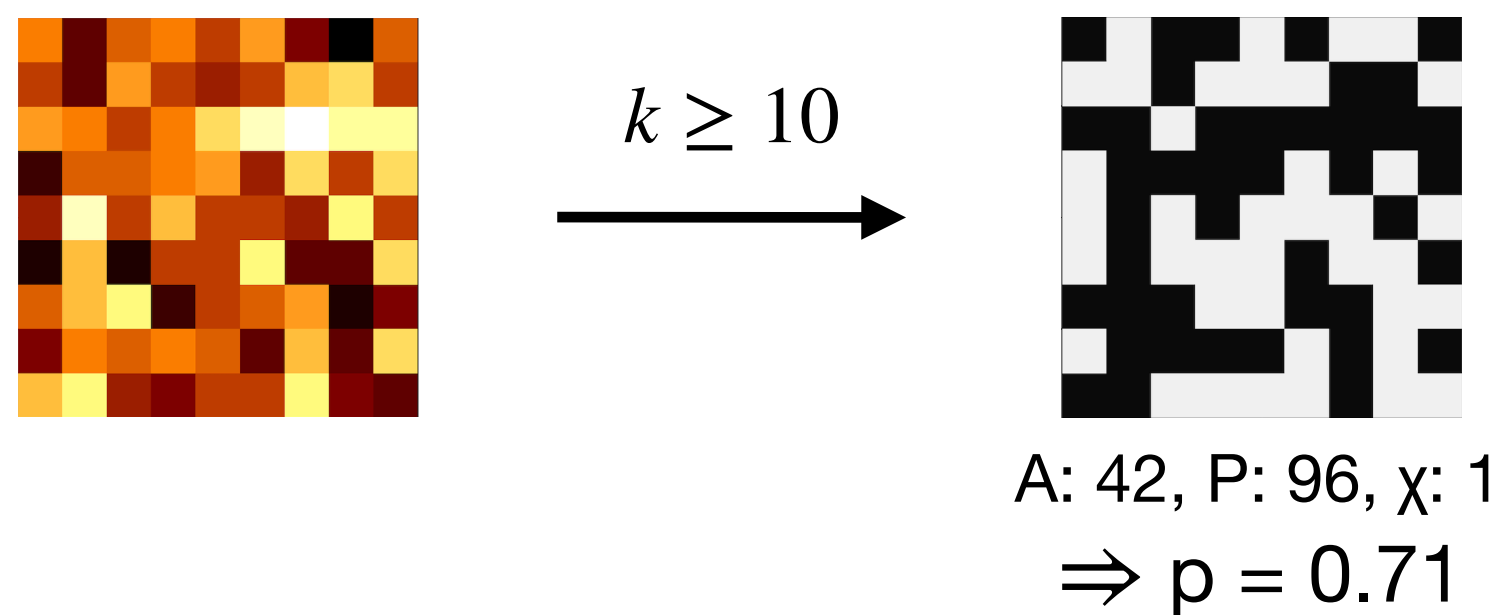


# Minkowski functionals and gamma-ray astronomy

Idea: Generate, from a counts map, a series of black and white images by thresholding. (Set pixel black if  $>$  threshold otherwise white)

Hypothesis test for each black and white image  $\rightarrow$  **Null hypothesis: Observed black and white image is from a Poisson background with mean  $\lambda$ , at threshold X.**

For all possible thresholds:



Take **smallest p-value** and **correct for trials**  $\Rightarrow p = 0.24$

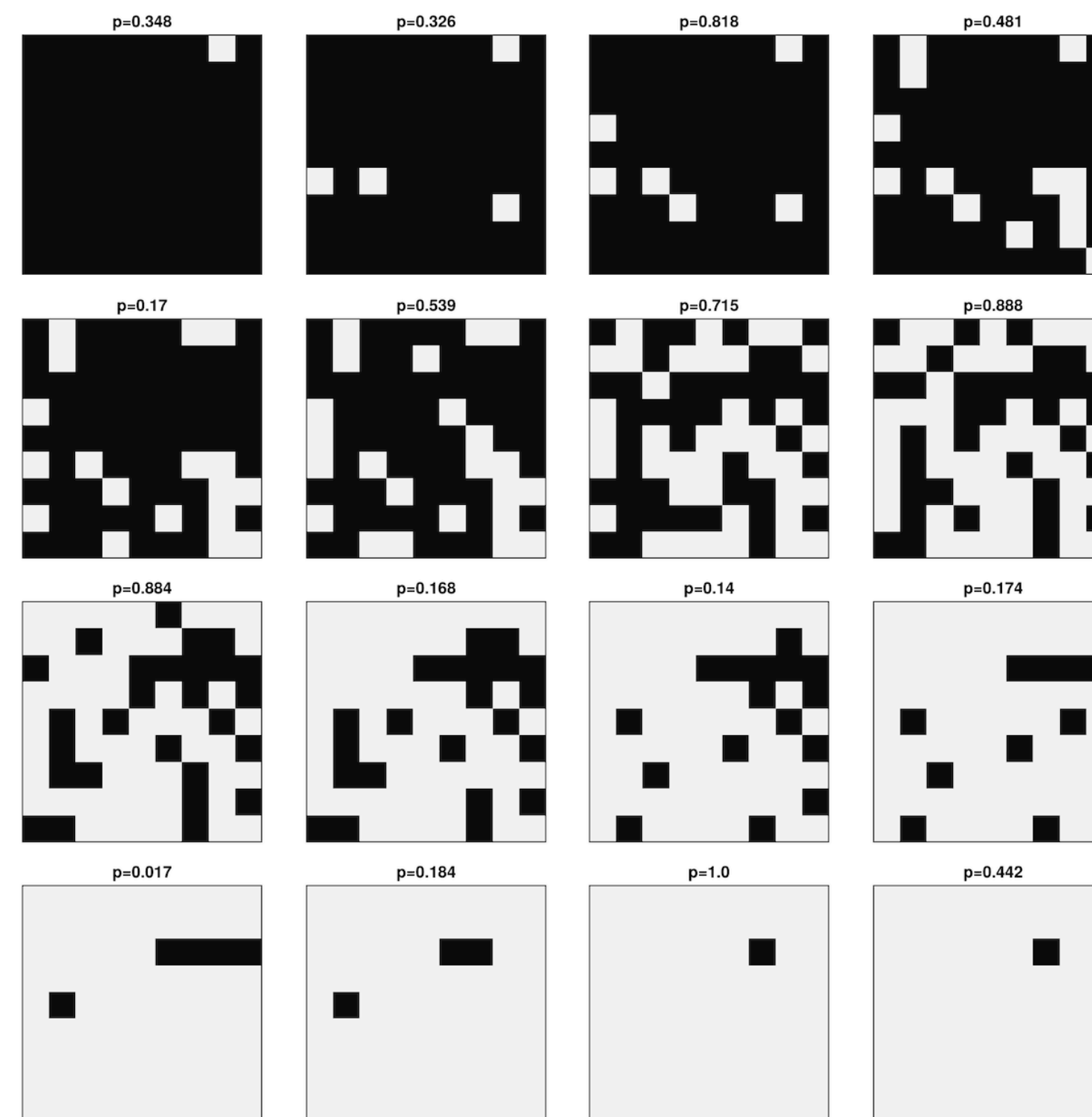
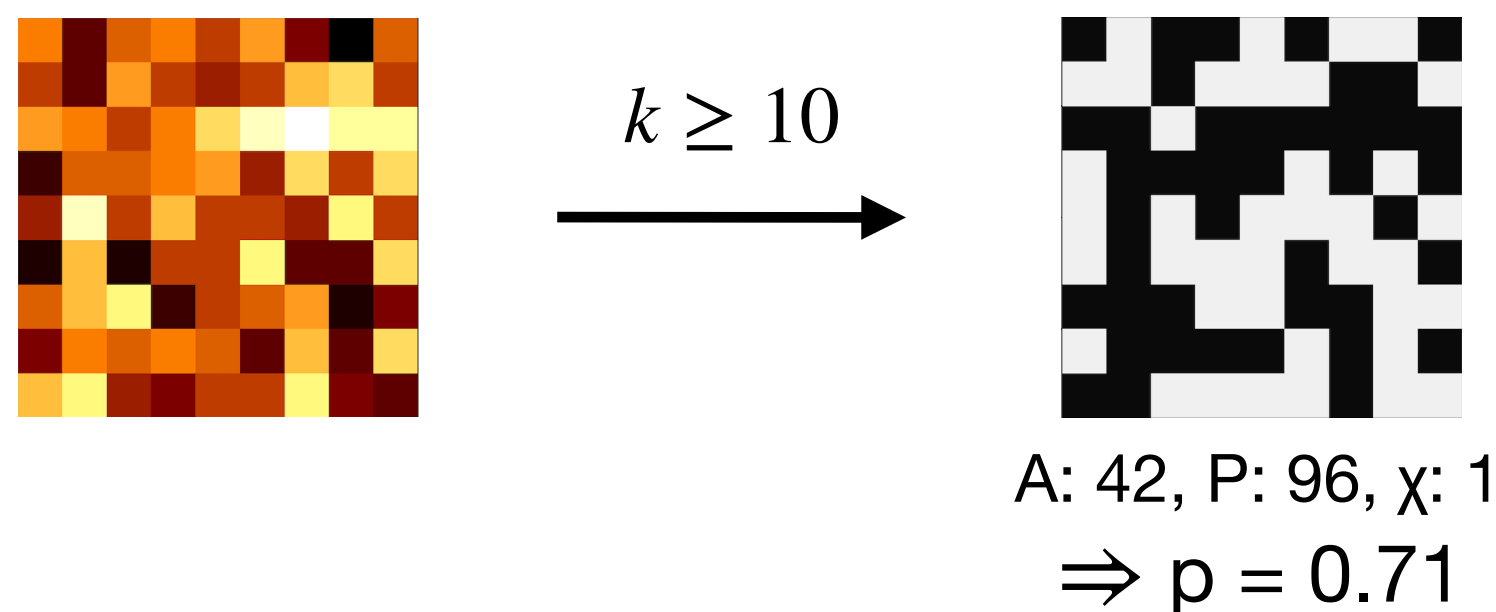
*$P = 0.71$  in words: Given the **hypothesis is true** there is a chance of 71% that this black and white image (at threshold 10) or an even more unlikely one appears.*

# Minkowski functionals and gamma-ray astronomy

Idea: Generate, from a counts map, a series of black and white images by thresholding. (Set pixel black if  $>$  threshold otherwise white)

Hypothesis test for each black and white image  $\rightarrow$  **Null hypothesis: Observed black and white image is from a Poisson background with mean  $\lambda$ , at threshold X.**

For all possible thresholds:



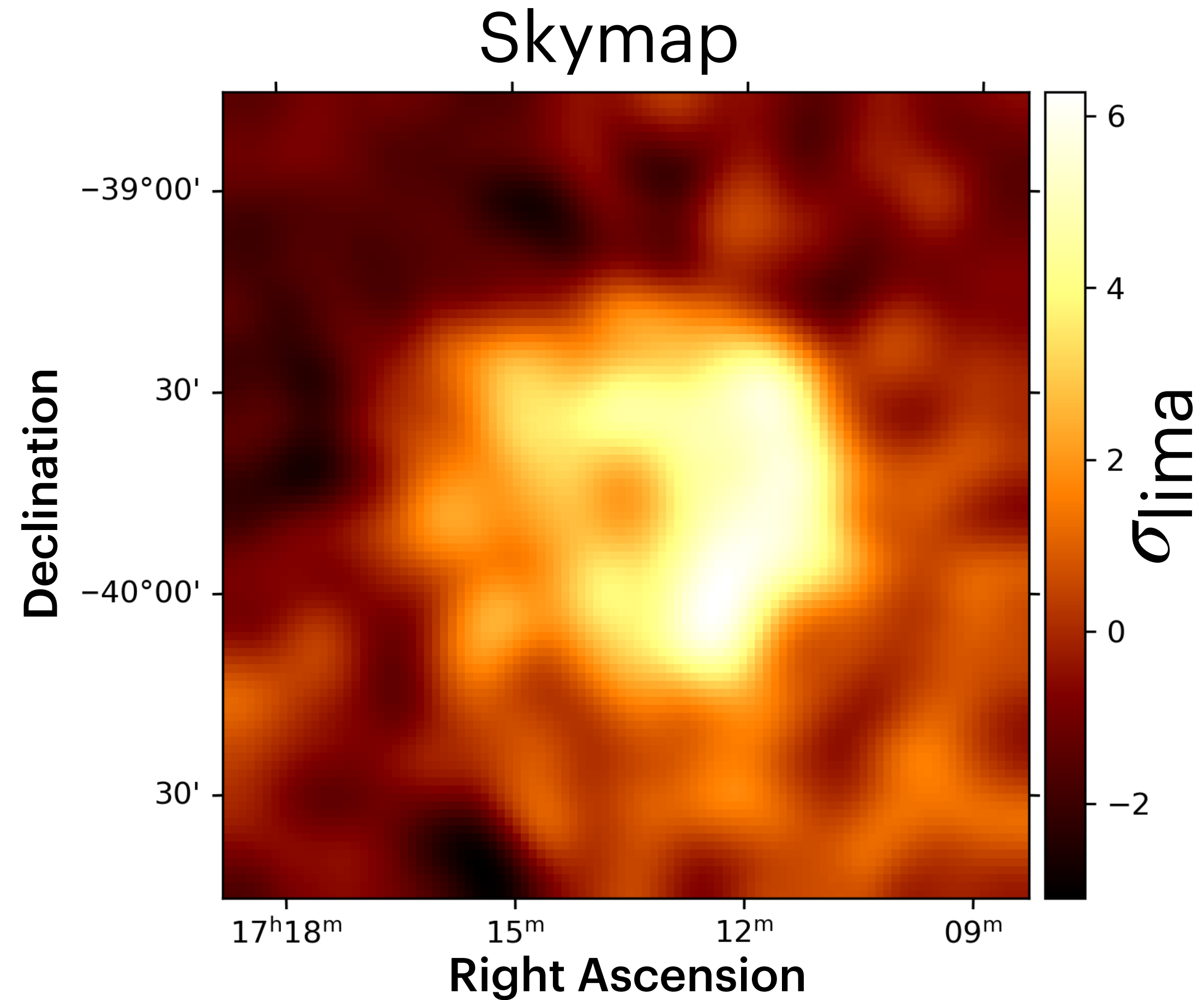
$P = 0.71$  in words: Given the ***hypothesis is true*** there is a chance of 71% that this black and white image (at threshold 10) or an even more unlikely one appears.

Take **smallest p-value** and **correct for trials**  $\Rightarrow p = 0.24$

*This we can convert to  $\sigma = 1.17$ ,*

# Minkowski sky maps: RX J1713.7 3946

- 15 observations from 2004 released in HESS open data 2018



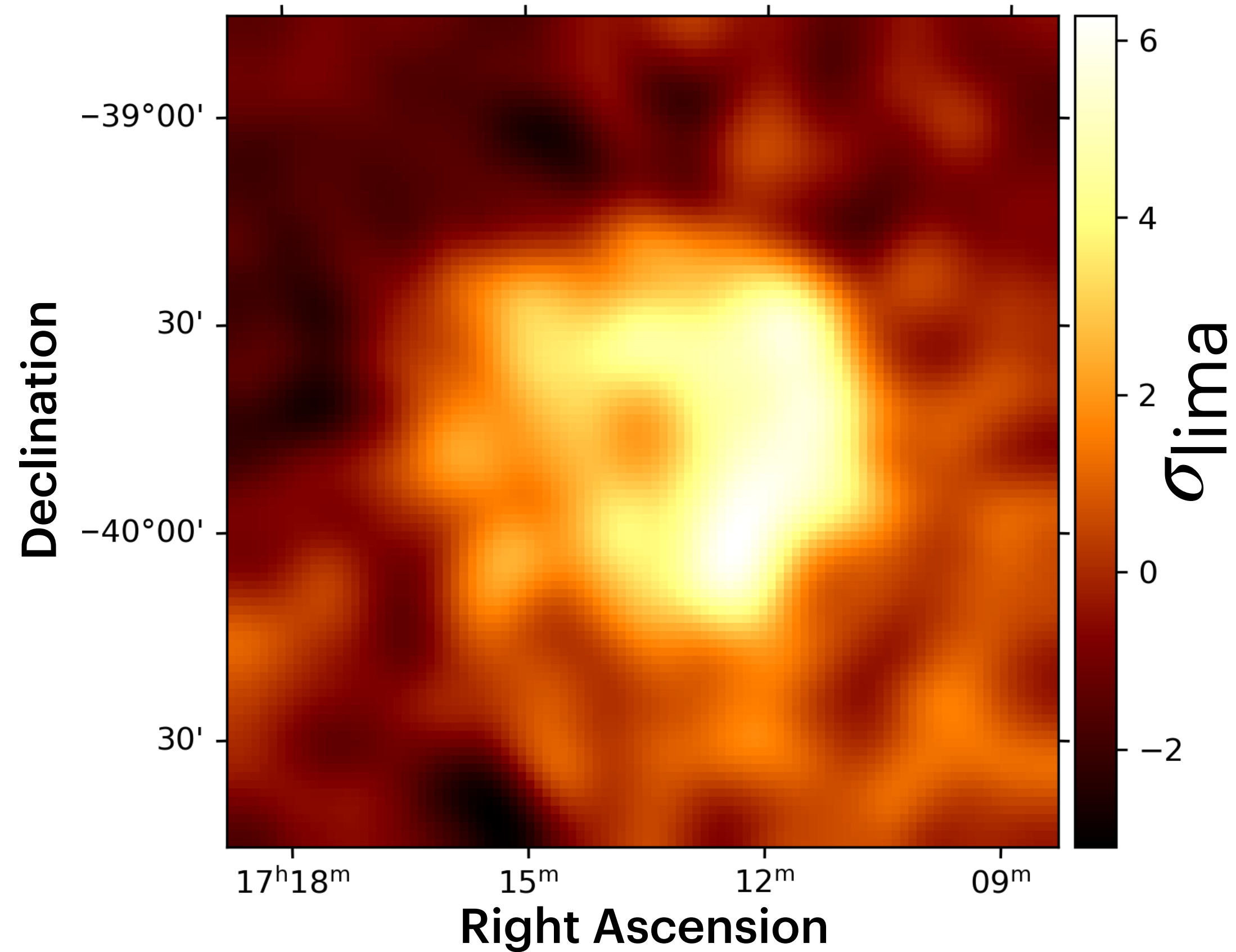


# Minkowski sky maps: RX J1713.7 3946

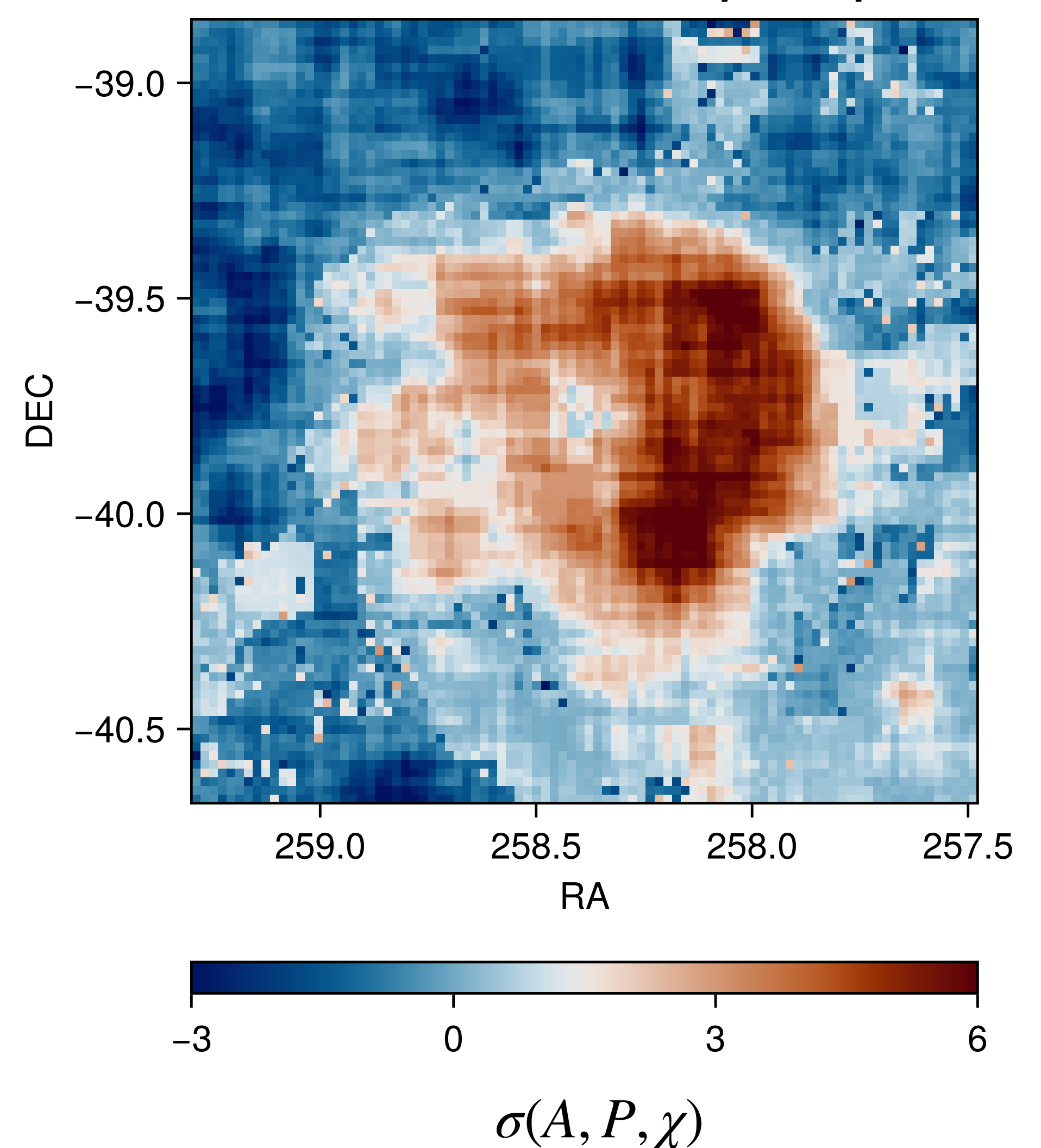
- 15 observations from 2004 released in HESS open data 2018



Skymap



Minkowski Skymap



## Advantages

- Standard LiMa method only sensitive to excess counts over background
- Minkowski functionals are sensitive to structural information
- Minkowski method can detect structures at different scales using the same kernel size

## Disadvantages

- Computationally expensive
- Less sensitive if no complex structure is in the observation

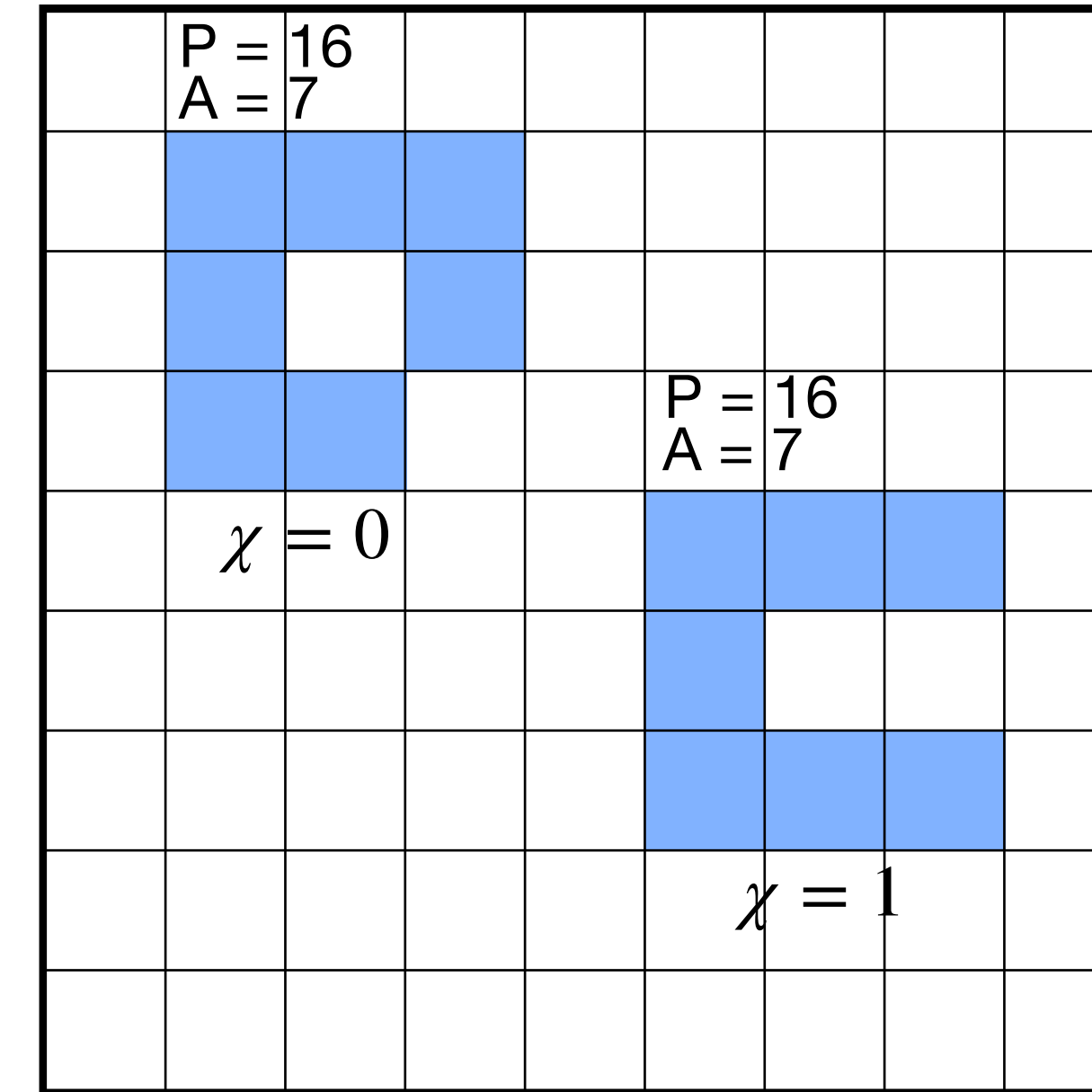
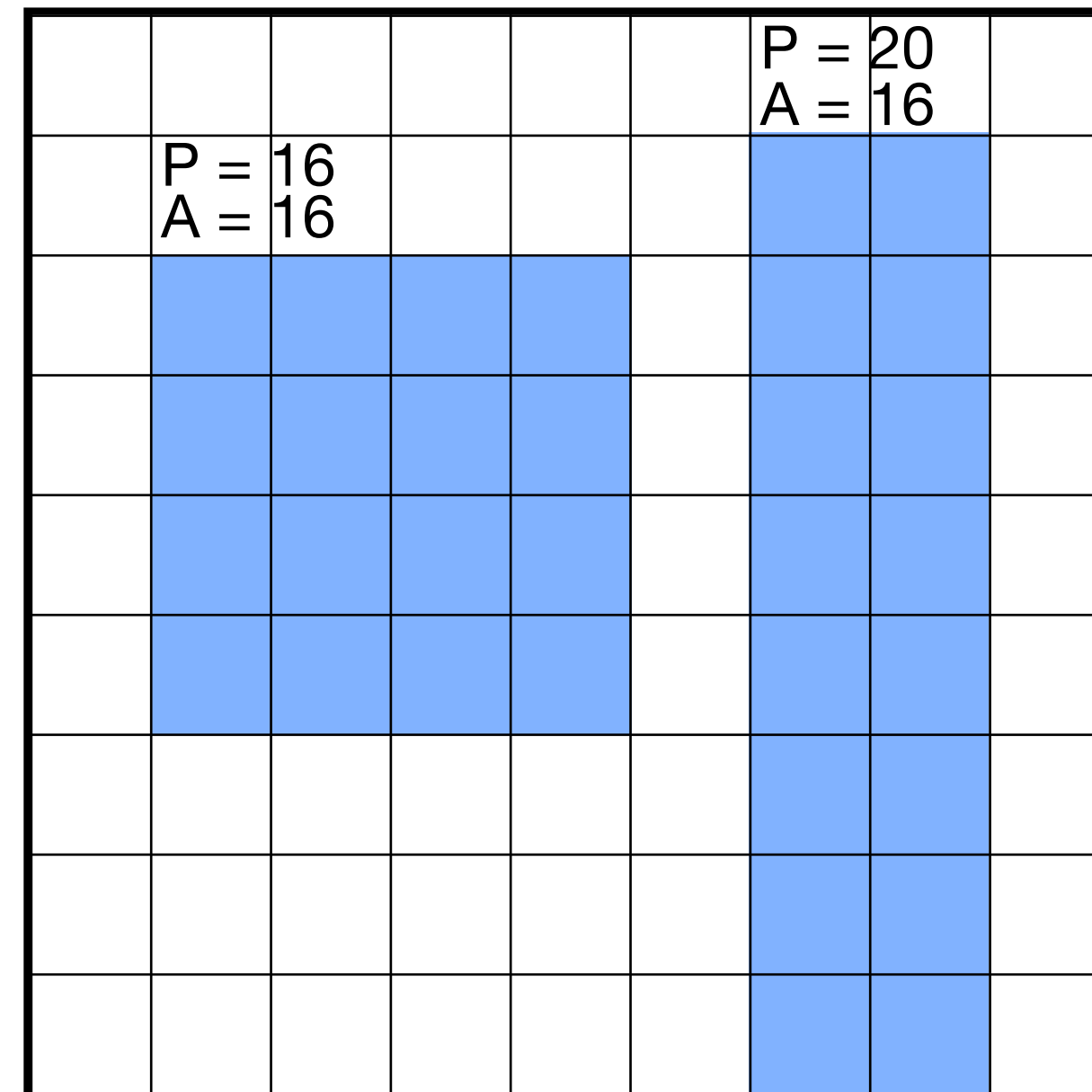
**Thank you for your attention!**



# Backup slides

Q: “How to evaluate the “geometry” of a subset of  $\mathbb{R}^n$  in a continuous and motion invariant way”

Area (A)  
Perimeter (P)  
Euler Characteristic ( $\chi$ )  
→ Counts “holes”



*Hardwiger’s Theorem in mathematics states these are all functionals which exist.  
With that we can classify the structure of black and white images.*