

The AGN - Neutrino Connection

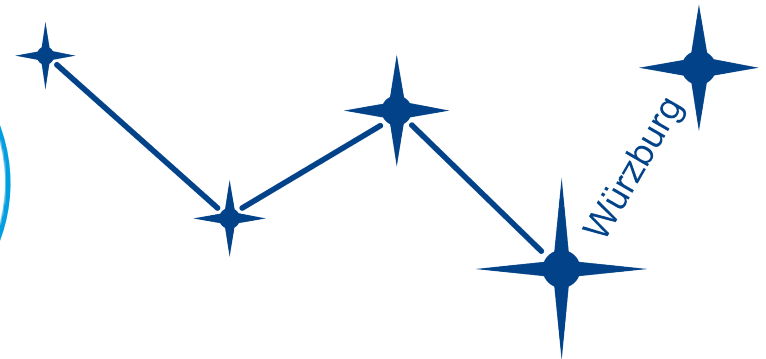
Annette Bremer

On behalf of the MessMapp group and Collaborators

FRANCI Meeting 2025, Bamberg

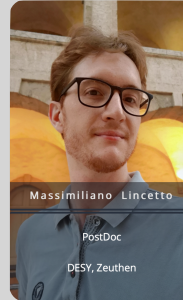
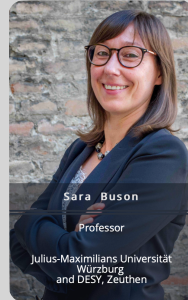


Based at:

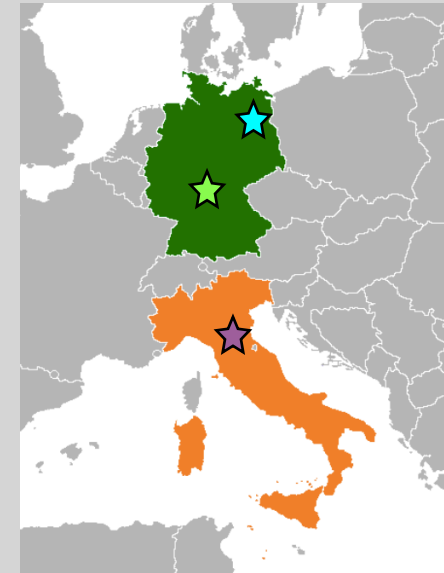
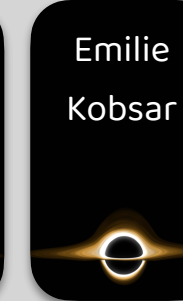
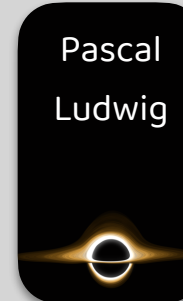
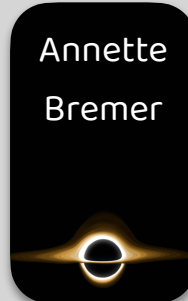
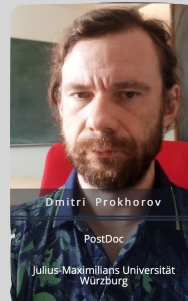


The MessMapp Research Team

★ DESY, Zeuthen:



★ JMU, Würzburg:



Collaborators:
M. Ajello, A. Coleiro, S. Marchesi,
M. Santander, A. Tramacere,
F. Vazza

Our Focus

AGN

Multiwavelength and
multimessenger properties

Observational data +
Theory

Here: focus on **neutrinos**

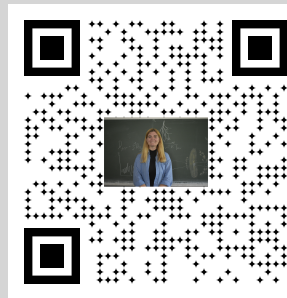
The physical properties of candidate neutrino-emitter blazars

Results

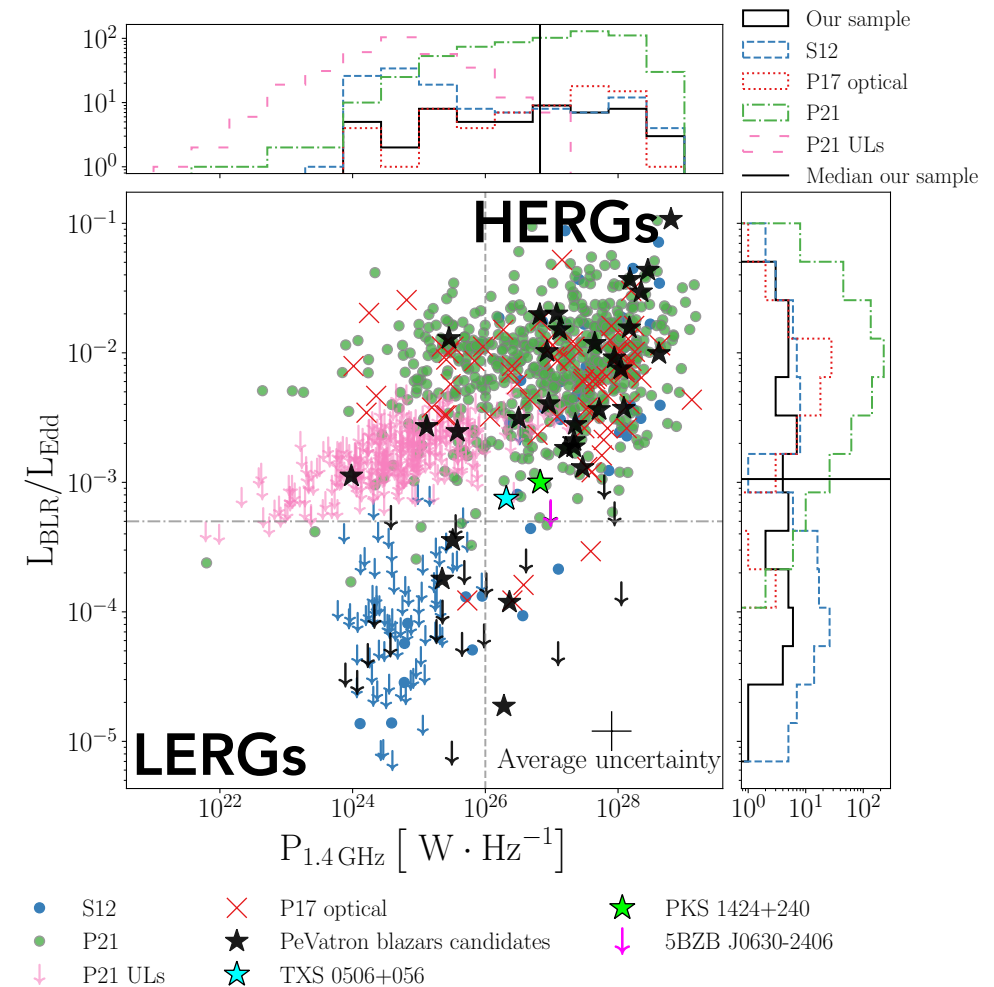
Mild tendency towards intense radiation fields and radiatively efficient accretion

Mild tendency towards powerful jets and high radio power

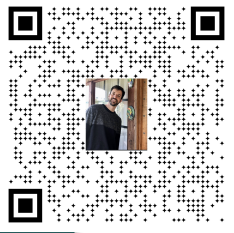
~60% HERG-like



Azzollini et al. 2025 (accepted)



X-Ray Spectral Variability as a Probe of Multimessenger Emission in Blazar 5BZB J0630-2406



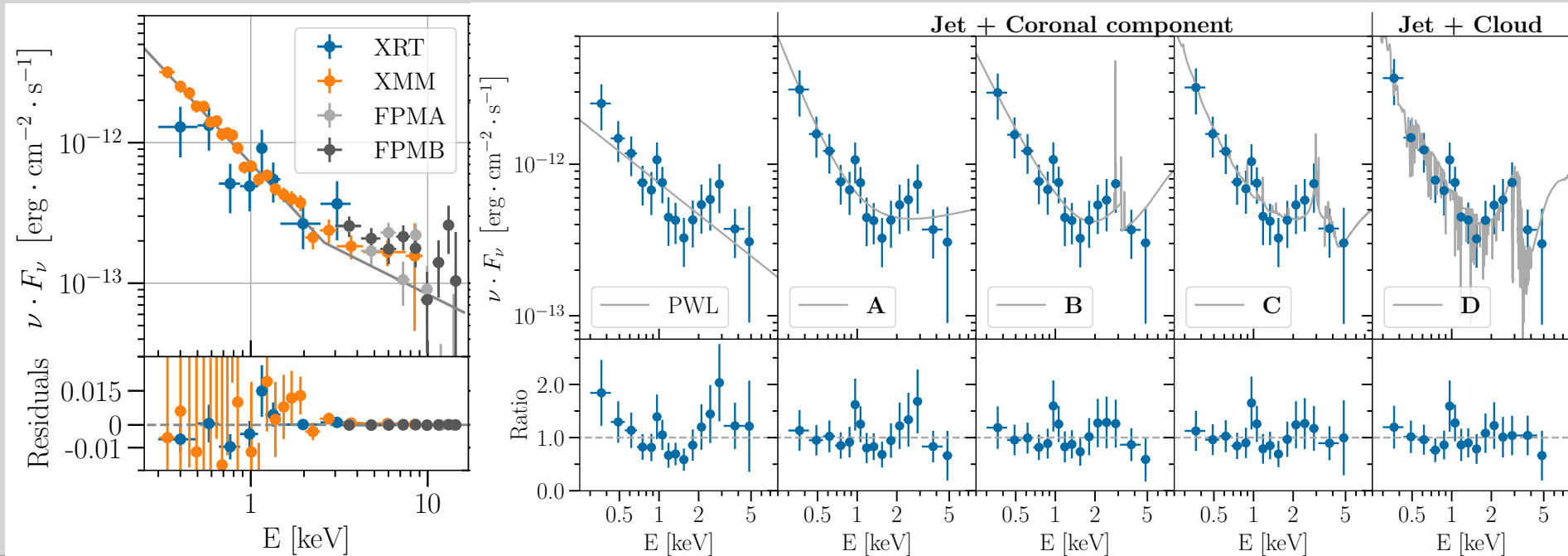
Additional component in low state spectrum

Prompts similarities to NGC 1068

Compatible with coronal emission

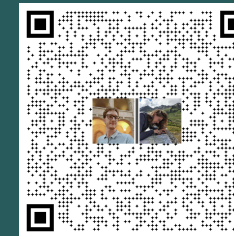
Hint for two neutrino processes

Sanchez Zaballa et al. 2025



IceCubePy

An open access software for neutrino data analysis

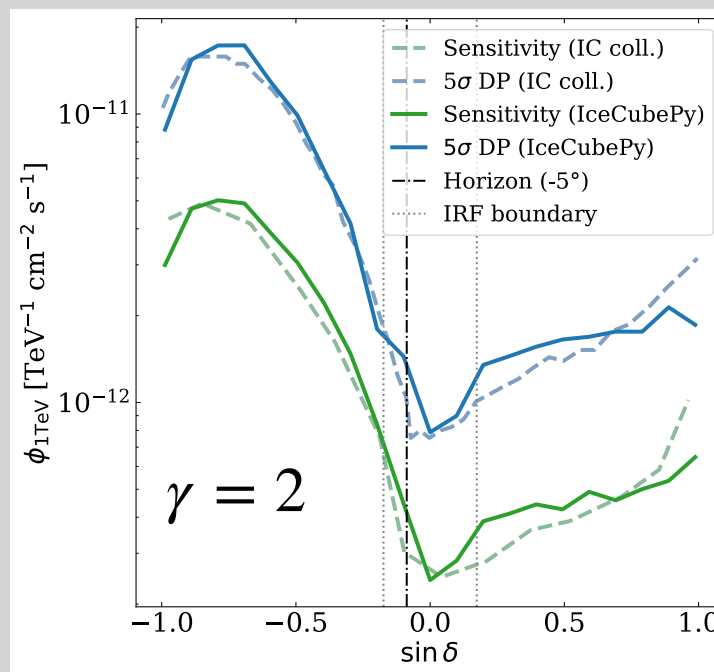


Python-based analysis tool using
unbinned likelihood

10-year IceCube muon track dataset

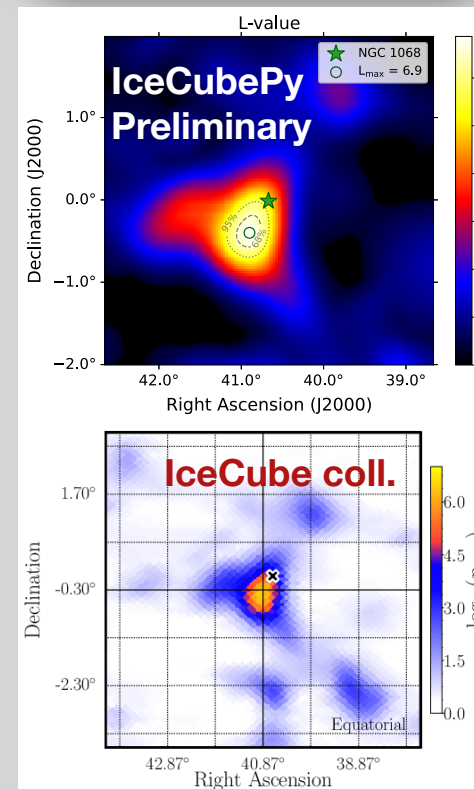
Reproduces IceCube collaboration results

Deviations can be attributed to the limited
accuracy of the released instrument response

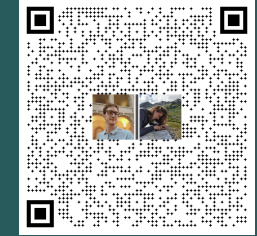


Lincetto, AB et al. in prep.

NCG 1068 Hotspot



IceCubePy — TXS 0506+056 Hotspot

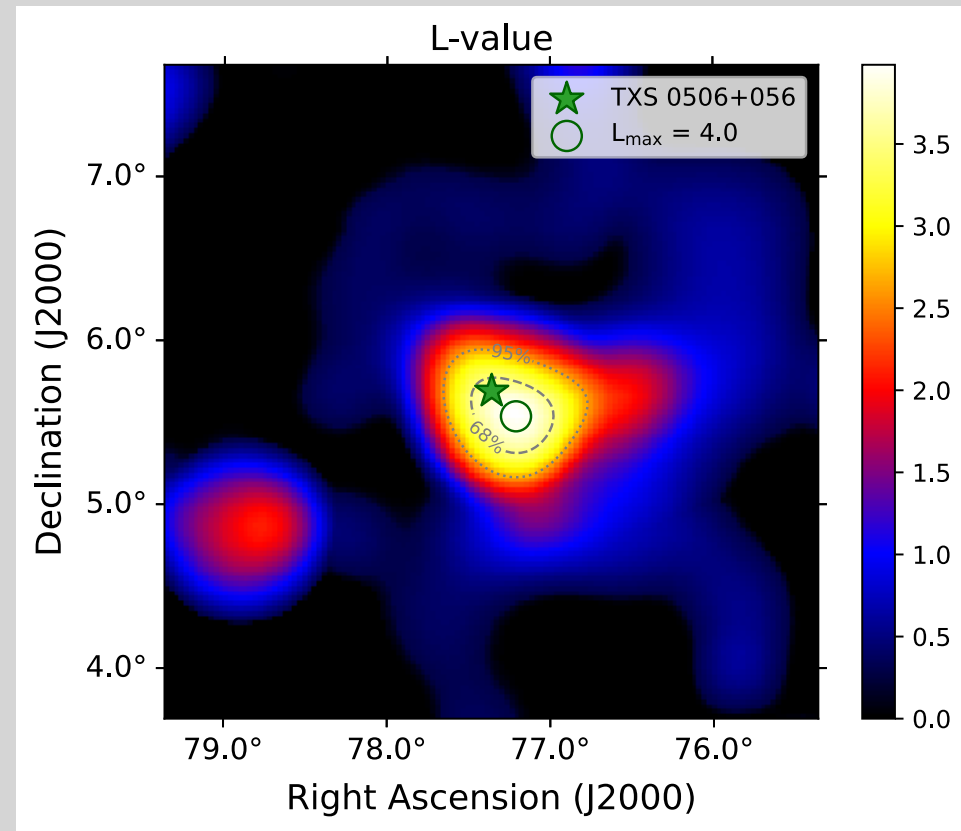


Map of local significance

L-value = $-\log_{10}(\text{p-value})$

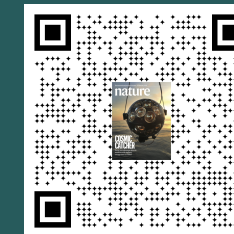
At TXS 0506+056 coords (77.36, +5.69):

- TS = 14.1
- $n_s = 16.1$
- $\gamma = 2.3$
- L-value = 3.5
 - IceCube coll. L = 3.72



Lincetto, AB et al. in prep.

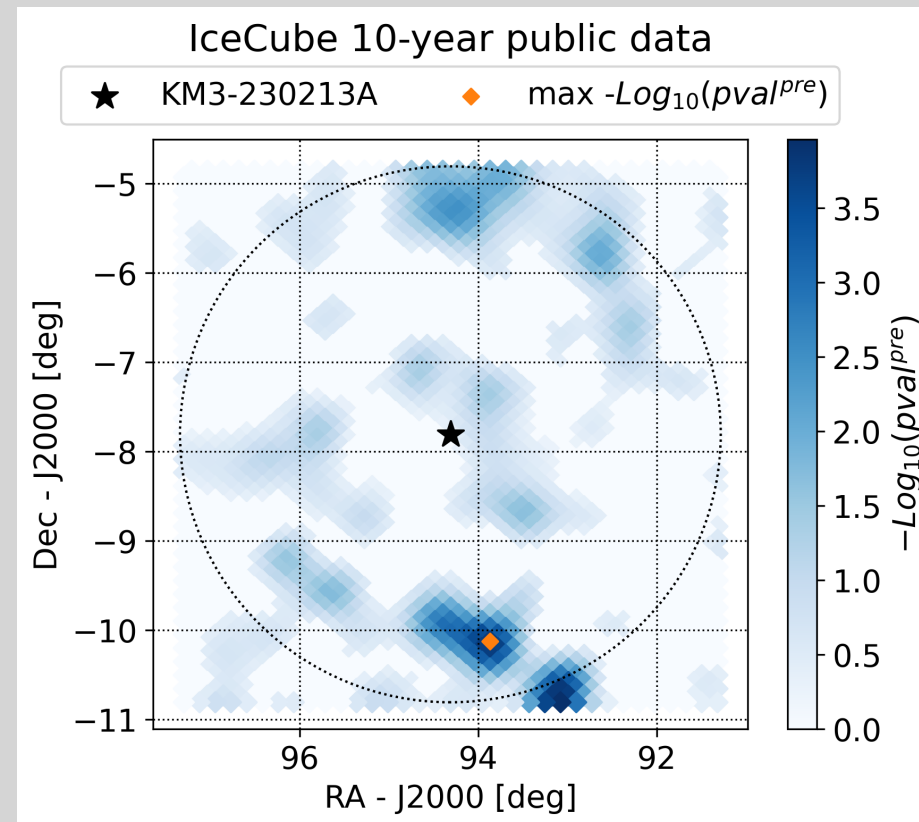
UHE Event KM3-230213A



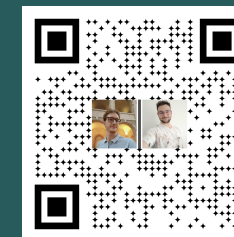
Confidence region of KM3-230213A

IceCubePy produced most stringent upper limits on steady neutrino flux from UHE event direction

Dataset	n_{sig}	$-\log_{10}(p\text{-value})$	Flux Upper Limit [$\text{GeV}^{-1}\text{cm}^{-2}\text{s}^{-1}$]	(RA, Dec) [deg, deg]	Distance [deg]	$P\text{-value}$
ARCA6-21	0.4	0.044	1.8×10^{-8}	(94.3, -7.8)	0.0	-
	1.3	1.308	1.9×10^{-8}	(96.7, -6.8)	2.6	0.44
ORCA6-18	0	-	2.1×10^{-7}	-	-	-
ORCA18-23	0	-	2.3×10^{-6}	-	-	-
ORCA-combined	-	-	2.0×10^{-7}	-	-	-
ANTARES	0	-	1.1×10^{-8}	(94.3, -7.8)	0.0	-
	1.9	1.936	1.7×10^{-8}	(94.4, -5.3)	2.5	0.53
IceCube	1.4	0.327	1.2×10^{-9}	(94.3, -7.8)	0.0	-
	15.1	3.782	6.3×10^{-9}	(93.9, -10.1)	2.4	0.07

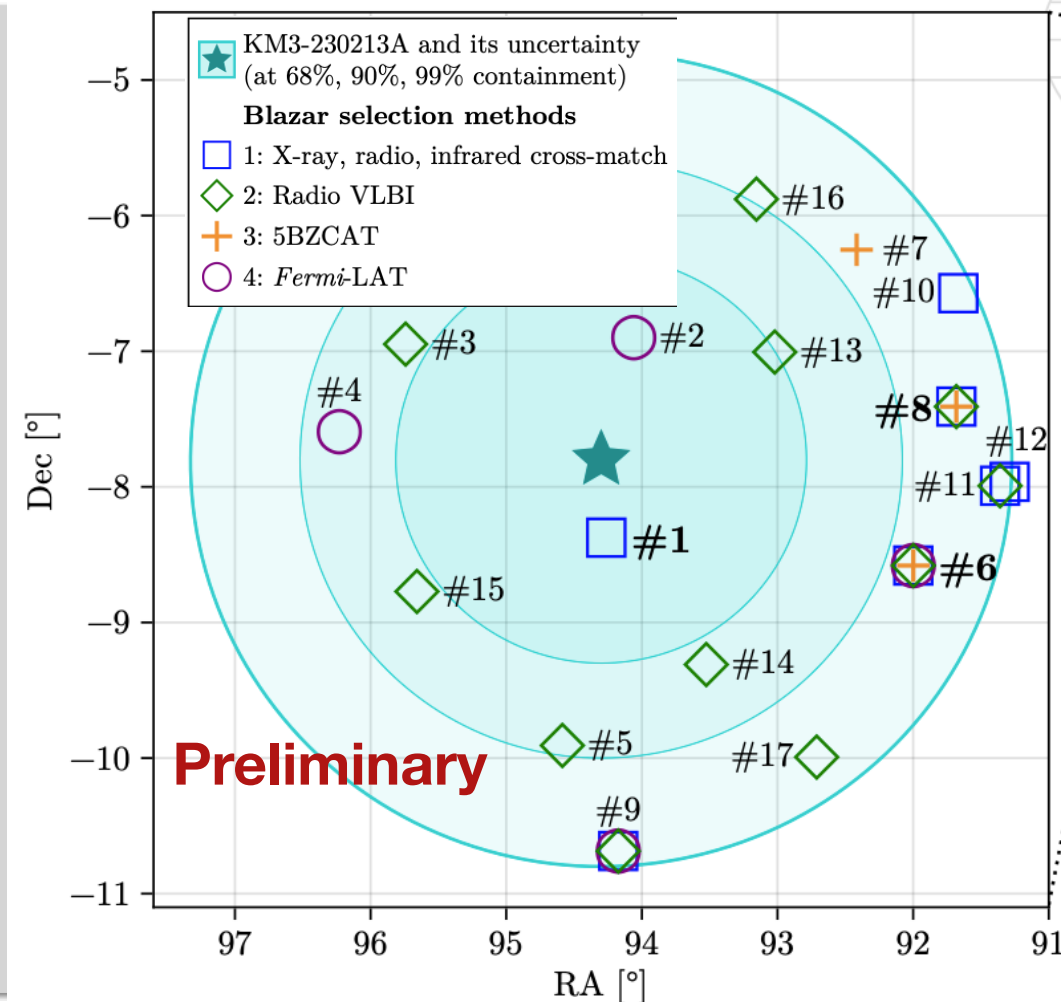
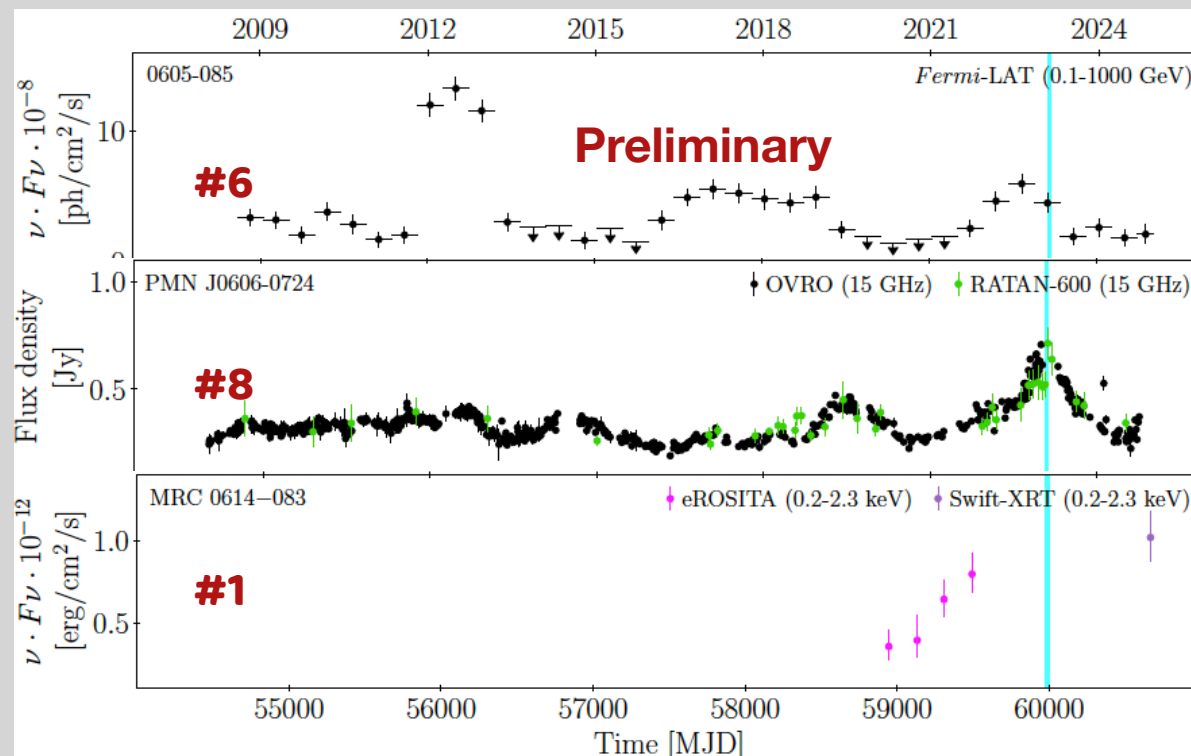


UHE Event KM3-230213A

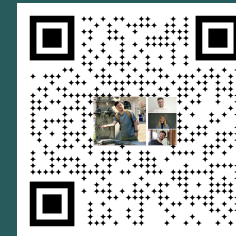


17 potential blazar-like counterparts

No conclusive evidence for individual source



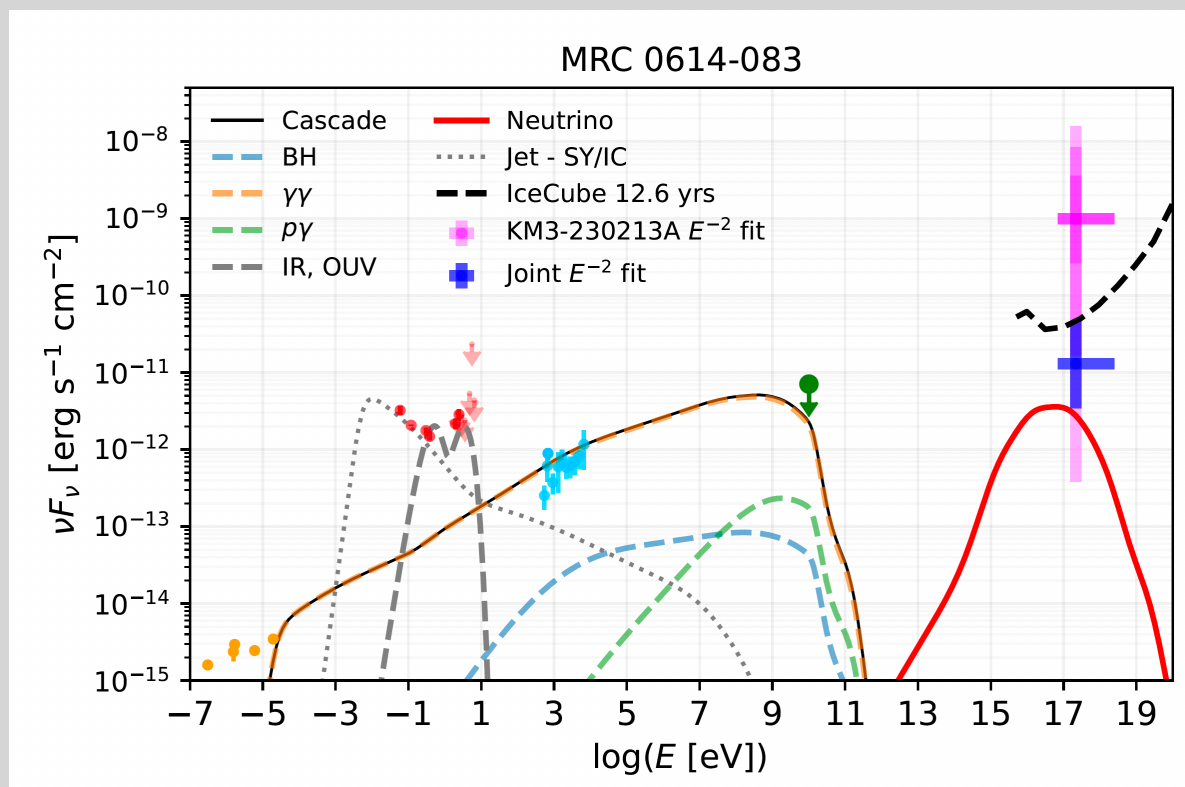
KM3-230213A origin: Accretion Flare Hypothesis



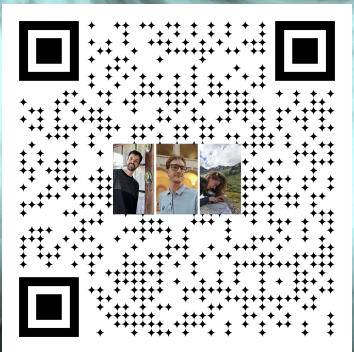
Closest Blazar to KM3NeT-230213A (#1)

SED and light curves described with
accretion flare

Predicts X-ray flare and neutrino production



[Yuan et al. (submitted)]



Summary

We embrace a **Multimessenger** approach combining **Neutrino** observational approach (IceCube, KM3NeT), **Theoretical** angle (SED modeling: AM3), **Multiwavelength** (photon) observations

- Analysis and results reproducible
- Datasets available via e.g. Zenodo
- Software hosted in GitHub: First public release: AM3, upcoming: IceCubePy
[GitHub.com/messmapp/IceCubePy-public](https://github.com/messmapp/IceCubePy-public)