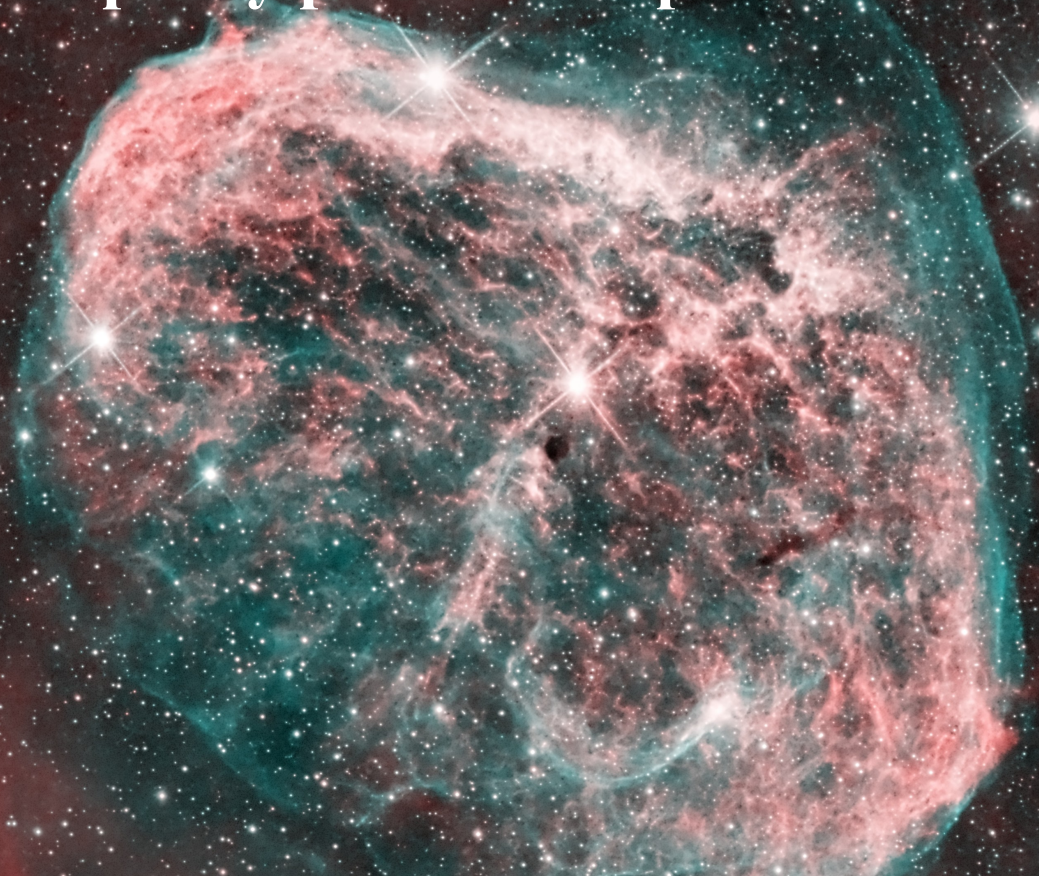


Astrophotography 101

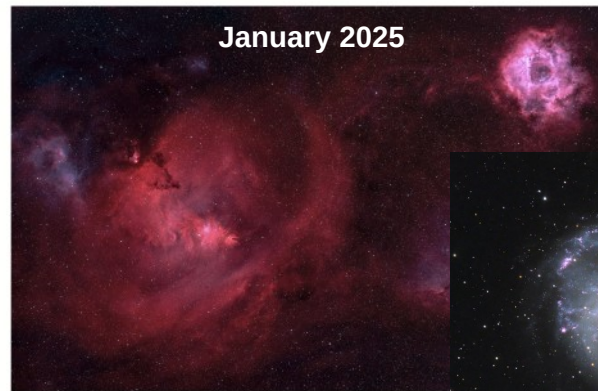
So, you want to take pretty pictures of space?

Astroparticle School 2025
Dr. Martin Rongen



Speaker introduction

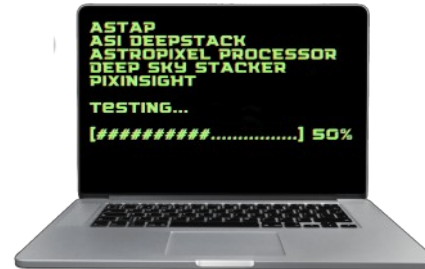
- Astroparticlephysicist @ ECAP working on the IceCube Neutrino Observatory
 - 2019: PHD from RWTH Aachen
 - 2020-2023: Postdoc in Mainz
 - 2023: joined ECAP as habilitation candidate
- Started in astrophotography in August 2024 (so still rather new to all this...)



Astrophotography as your miniature science experiment

- Chose your science case
- Design, fund & build your experiment
- Gather the data
- Process the data

SNRs
milky Way
Target
Nebulae
Planet
Galaxy



Types of astrophotography

DSLR & Tripod
F500 Rule

Short Scope
& Star Tracker

Long Scope
& Heavy Mount

Landscape

Widefield: Milky Way nebulae

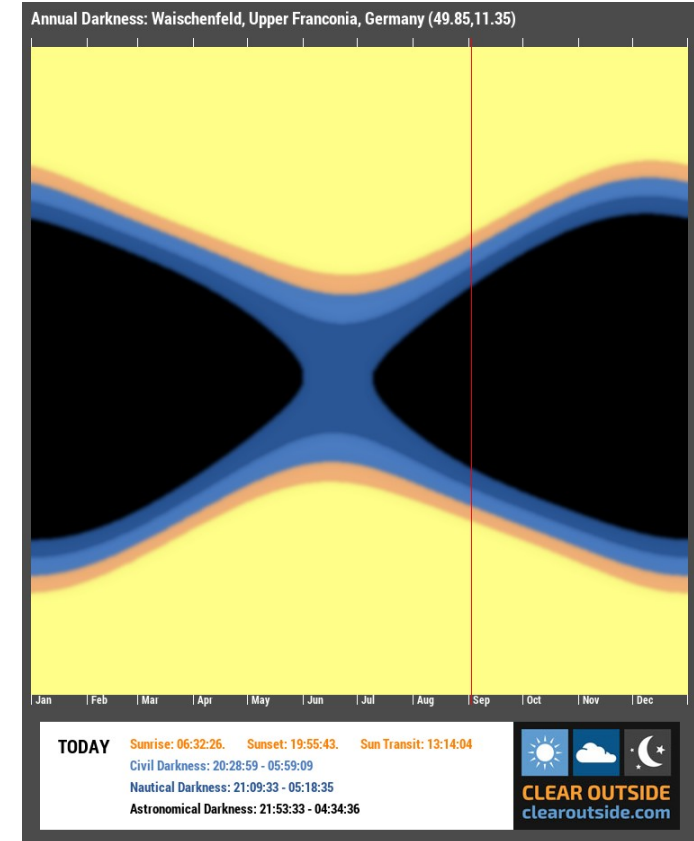
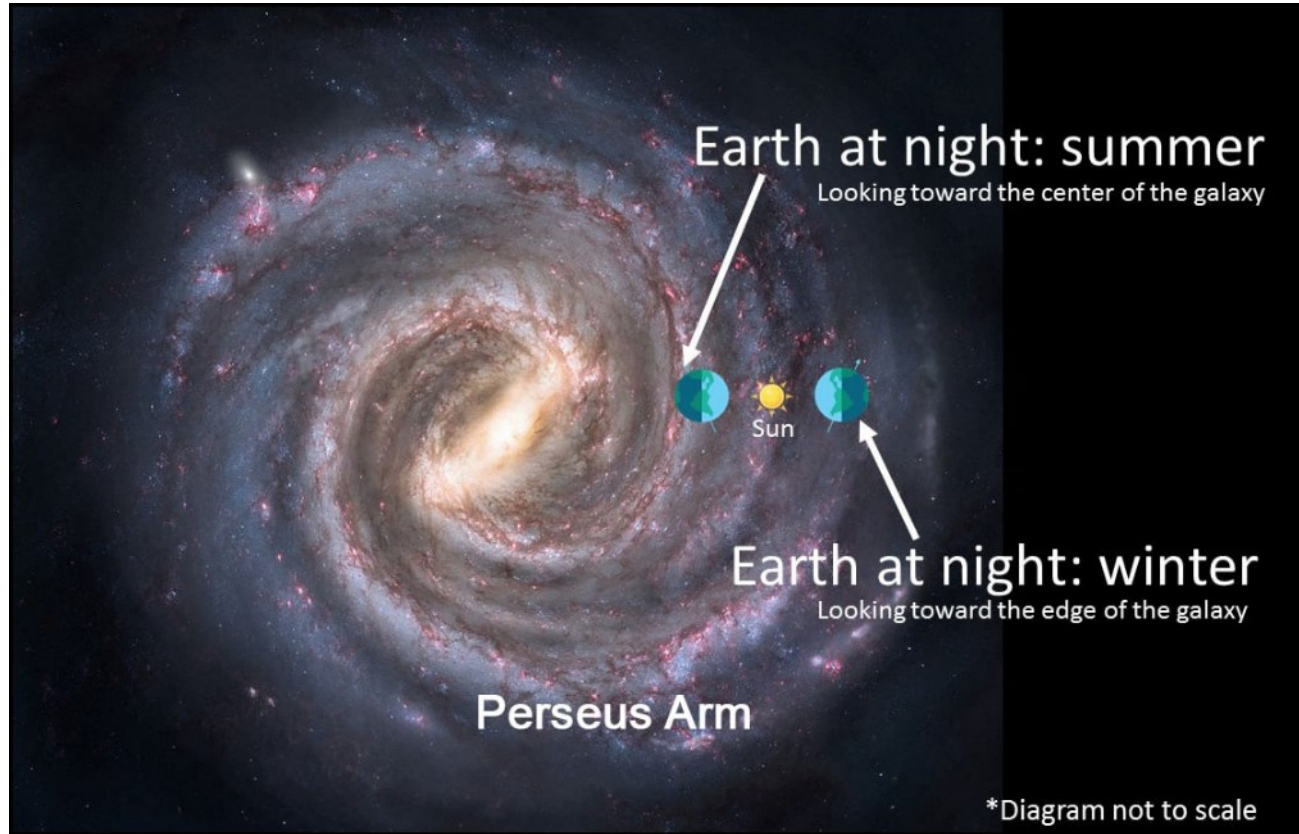
Field of view (FOV)

Cost & complexity

© Astrobiscuit

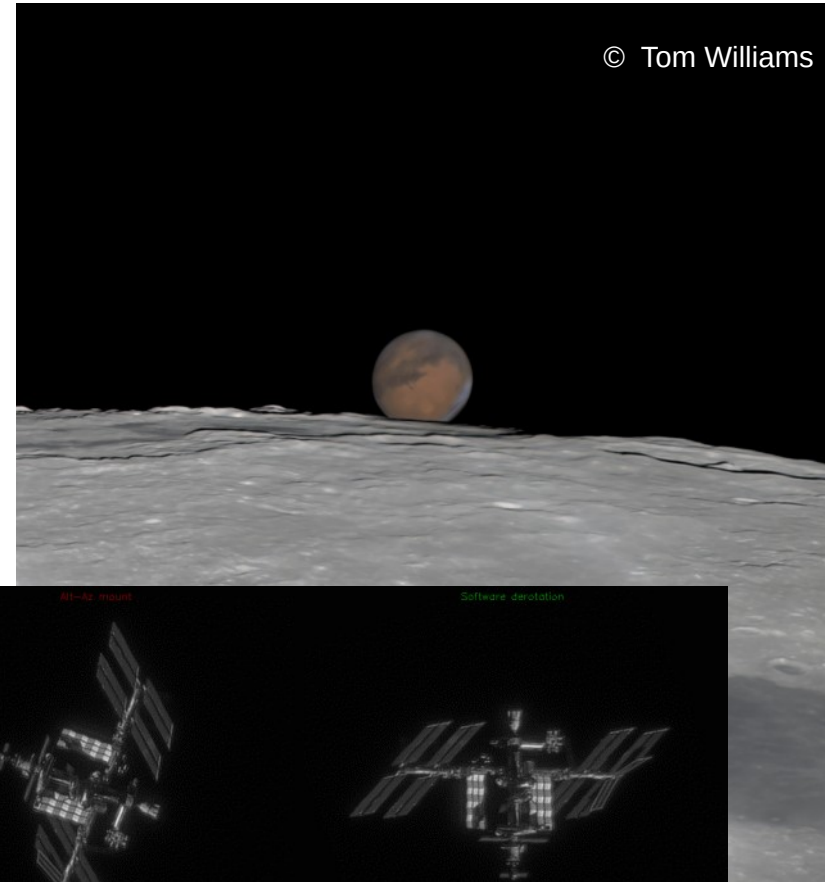
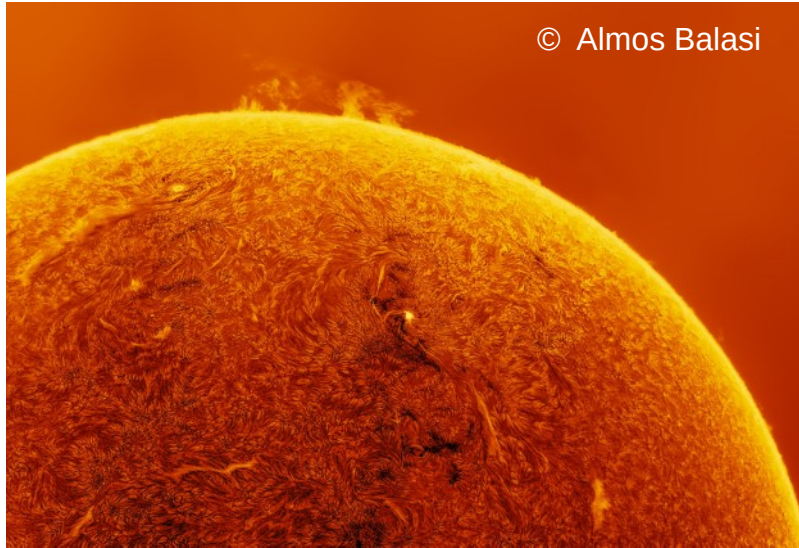
© DSC

The seasons of astrophotography



Landscape / milky way nebulae astrophotography in summer & autumn.

Solar, lunar, planetary & satellites



Plenty of other specialized fields
requiring unique techniques
(e.g. lucky imaging) and equipment.



Equipment overview



Every setup consists of 3 major components:

- A sturdy mount (tripod – EQ mount)
- The optics (lens – telescope)
- A camera (DSLR – astro camera)

And lots and lots of little things:

- Power supplies
- Data storage
- Focusing masks

Untracked setup

- Camera, Lens, Tripod, **Intervallometer**
- Landscape astrophotography sadly only makes sense with very little light pollution
- Long exposures yield star trails
- F500 rule for sharp stars:
 - Exposure: $500s/\text{focal length}[\text{mm}]$
 - Modulo crop factor, ...
e.g. PhotoPills app for details
- Sensible focal length range:
 - 50mm: $\sim 20^\circ$ FOV
 - 135mm: $\sim 7^\circ$ FOV
(moon is about 0.5° diameter)



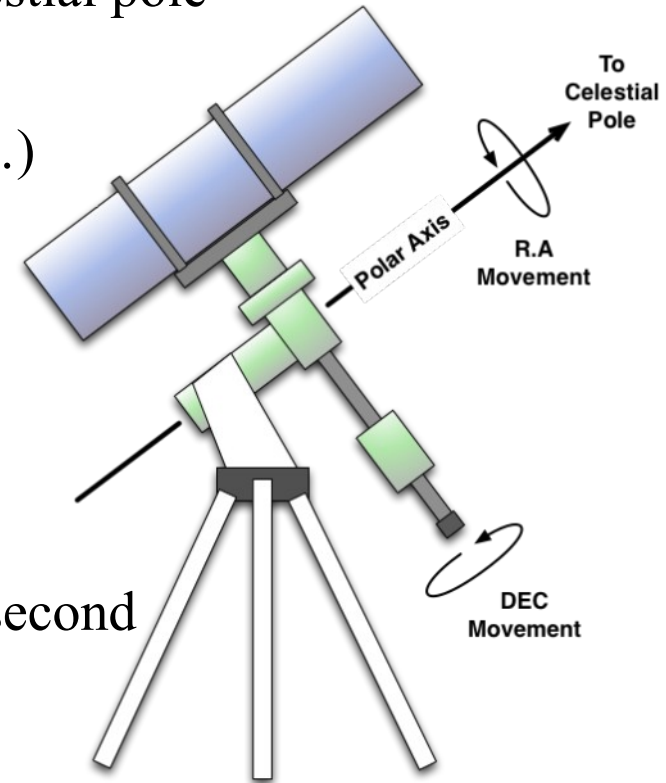
Star trackers

- Same setup as before, but now adding one-axis constant-speed motor to compensate Earth rotation
- Exposure times limited by accuracy of tracker, generally around 30-45 seconds
- Focal length limited by maximum weight tracker can support, 135mm max still sane

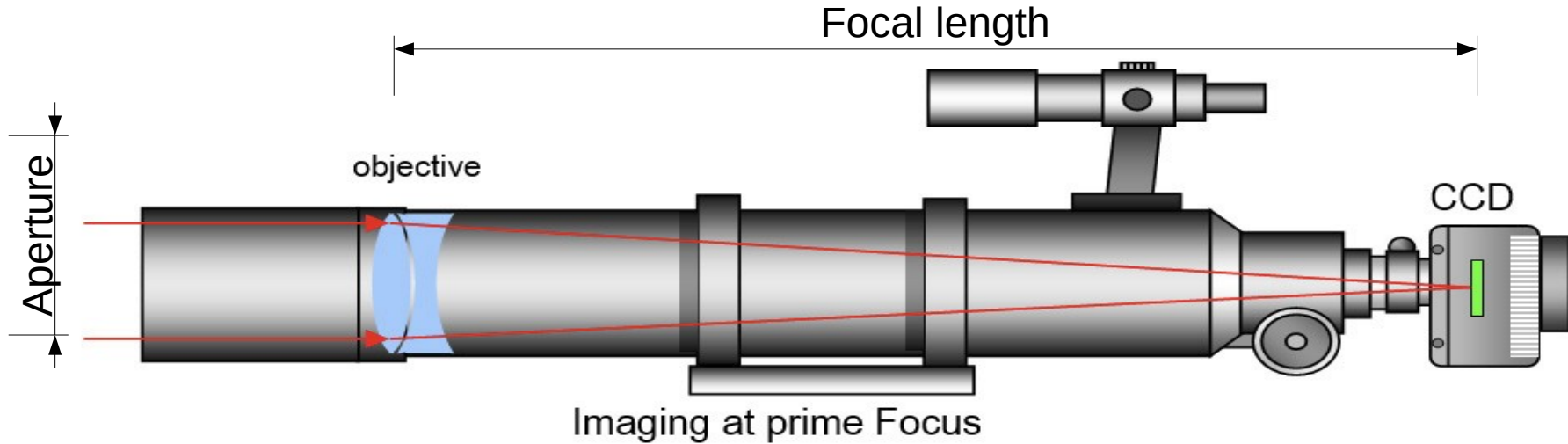


Equatorial mounts

- Telescope mounted rigidly and two axes to be able to slew to any direction
- Right ascension axis aligned with Earth rotation / celestial pole
Declination axis at 90°
- Alt-Azimuth mounts have their specific uses (visual...) BUT are generally not preferred as they introduce field rotation (can keep target centered, but image rotates)
- Also allows for guiding:
 - While primary telescope is taking long exposures
 - Secondary guide telescope takes an image every second and sends correction signals to mount

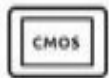


Optics basics: Focal length & aperture

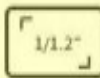


- Telescopes project an angular separation onto a focal imaging plane
- Longer focal length \rightarrow larger angular magnification \rightarrow less intensity per unit area
- Larger aperture \rightarrow more photons in total
- Image brightness given by **f-ratio: focal length/aperture**
- Slow scopes: Large f-ratio; Fast scopes: Small f-ratio

Camera basics: Pixel characteristics



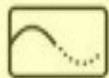
Sensor
IMX585



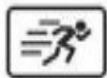
1/1.2"
11.2×6.3mm



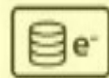
Resolution
3840×2160



ADC
12bit



FPS
47



Full well
MC 40Ke/MM 38.7Ke



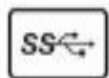
Read noise
0.7e



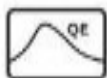
Cooling Tempe
35°C



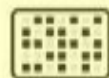
DDR3 Buffer
512MB



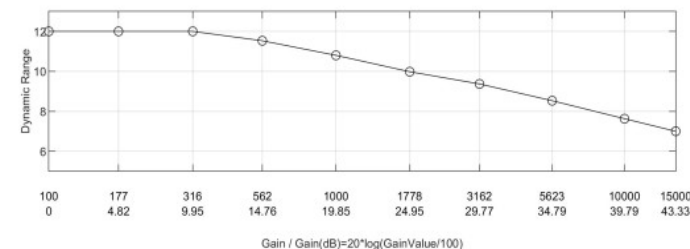
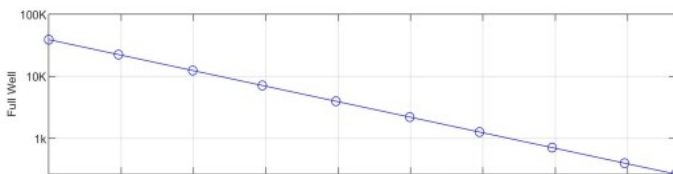
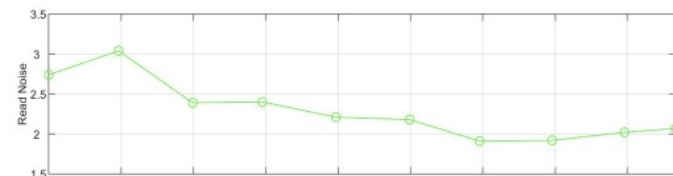
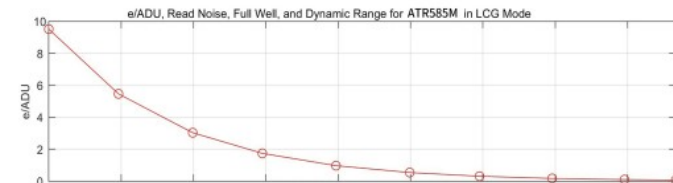
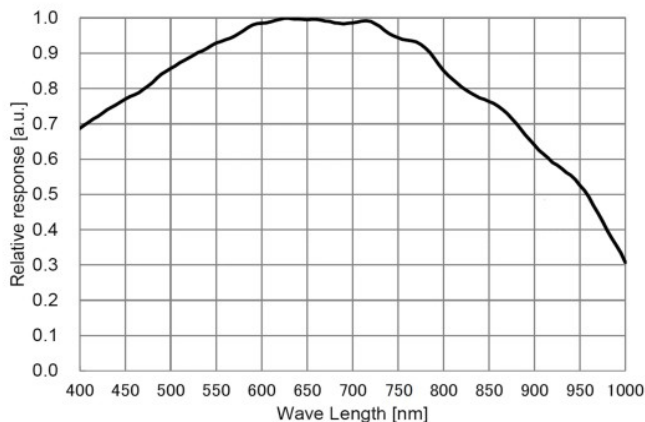
USB
3.0



QE
MC 88%/MM 91%

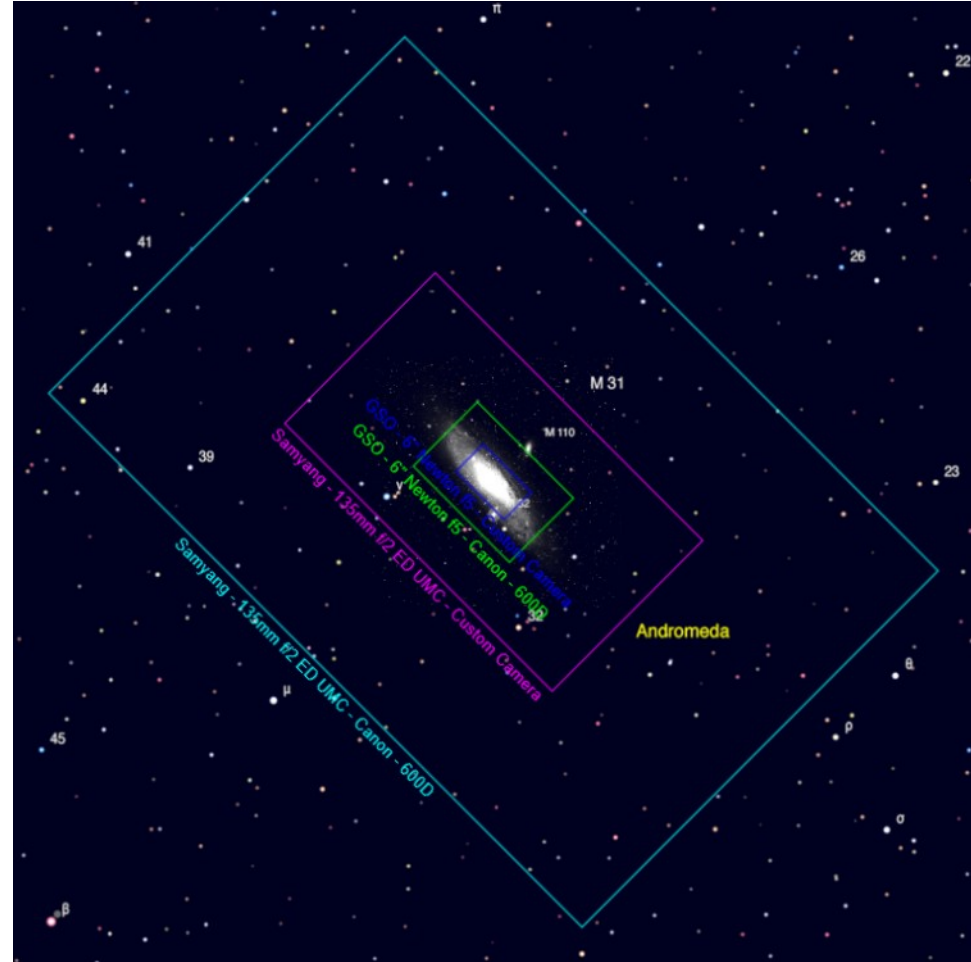


Pixel Size
2.9μm



Angular resolution and FOV

- The telescope focal length and camera pixel size determines the per-pixel angular resolution
- Multiplied with horizontal and vertical number of pixels yields the FOV
→ needs to be matched to target
(or dedicate to a mosaic)
- Actual resolution mostly limited by atmospheric seeing (typically 1")
 - Oversampling does not improve image
 - Undersampling can increase SNR but degrade resolution
- Additional diffraction Rayleigh limit for small apertures
- Calculator @ https://astronomy.tools/calculators/ccd_suitability



Refractors & lenses



- Most camera lenses are BAD
(Your stars are purple and large?? You have chromatic aberration....)
- Prime lenses can be good (well known Rokinon/Samyang 135mm f2)
but often need to be stepped down (closing the aperture for larger f-ratio)
- Dedicated astro-lenses / refractors often use expensive fluorite optics



Canon kit lens



Canon zoom lens

MR 2024

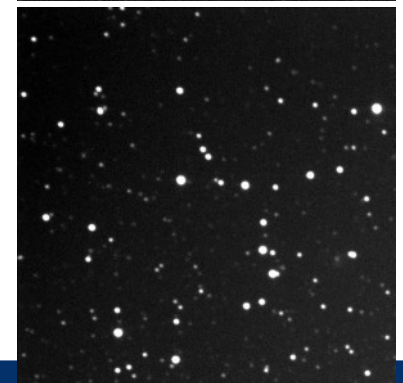
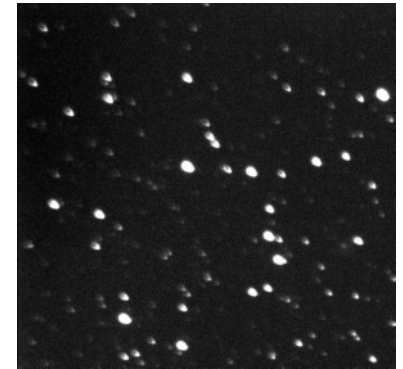
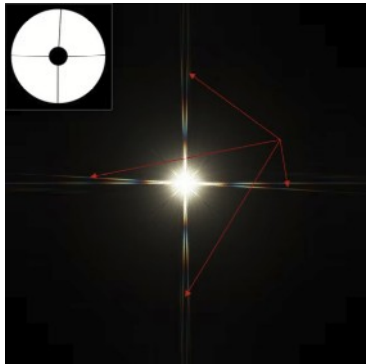


Samyang 135mm

MR 2024

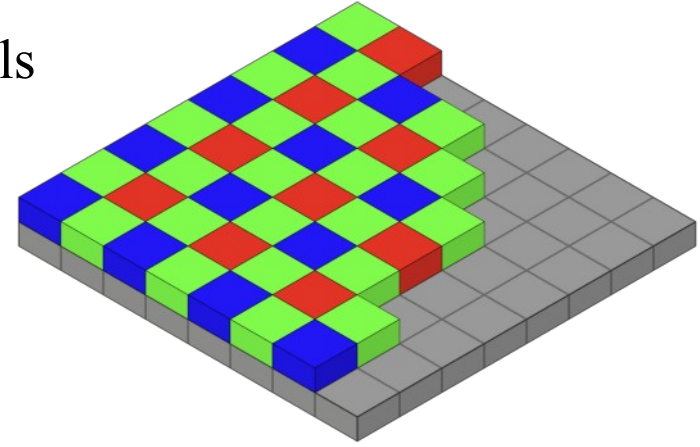
Reflectors (here only discussing Newtonians)

- Cheapest way to get a decent f-ratio at larger focal lengths
- But reflectors suffer from comatic aberration requiring another expensive piece of equipment (coma corrector)
- Newtonians also require alignment of the optical train (called collimation) (can even drift as telescope slews)
- And lots of fine-tuning for good optical quality (focussing, secondary spider alignment for symmetric diffraction spikes)



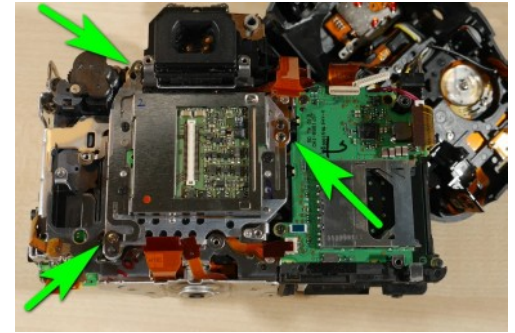
Color vs. Mono cameras

- Color cameras use Bayer color mask in front of pixels (2 green, 1 blue, 1 red to match human color perception)
- „Real resolution“ only 50% of what is quoted
- Few things in sky actually green...
- Mono cameras do not feature Bayer mask
 - 3x photon count without filters (luminance)
 - Stacking images with color filters gives color image
 - Allows for efficient narrowband imaging
- Filter wheels allow for automatic cycling of filters during imaging session



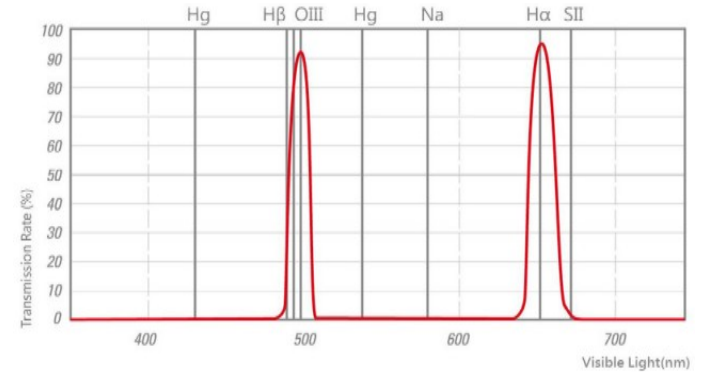
DSLRs vs. Dedicated astro cameras

- DSLR:
 - Multi-purpose, often already on hand
 - Cheap when used (< 200€)
 - Low sensitivity, but UV-IR cutoff filters can be removed
 - uncooled → high noise (especially in Summer)
- Astrocams:
 - Require dedicated PC to readout
 - Can be mono!!
 - Cooled and low noise



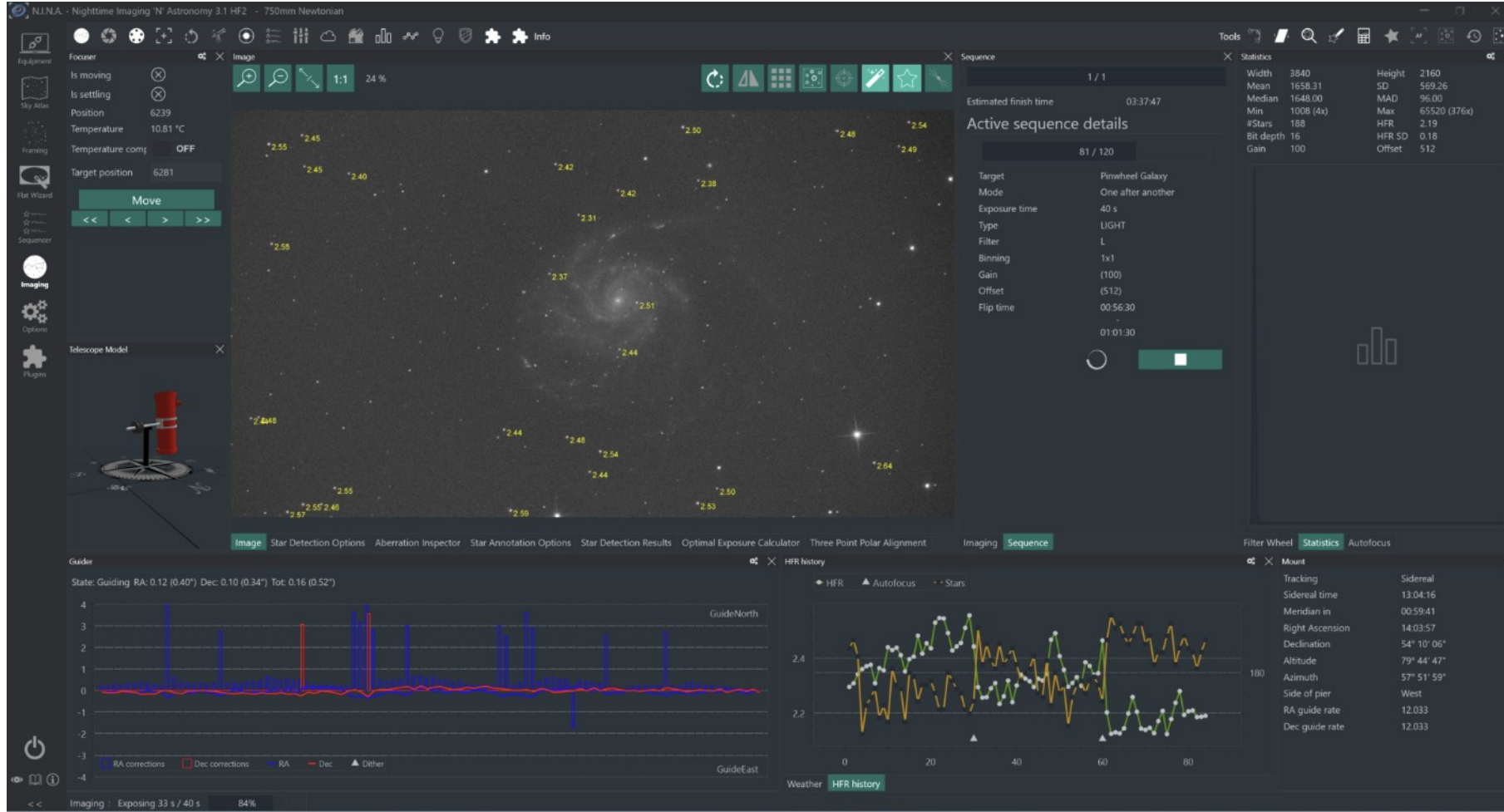
(Narrowband) filters and SNR

- Selecting discrete emission lines (non-thermal component)
 - Interesting to look at specific elements
 - Suppresses most light pollution → higher SNR
- Dual-narrowband (Ha, OIII) amazing for DSLRs
- Broad-band light pollution filters mostly not worth it



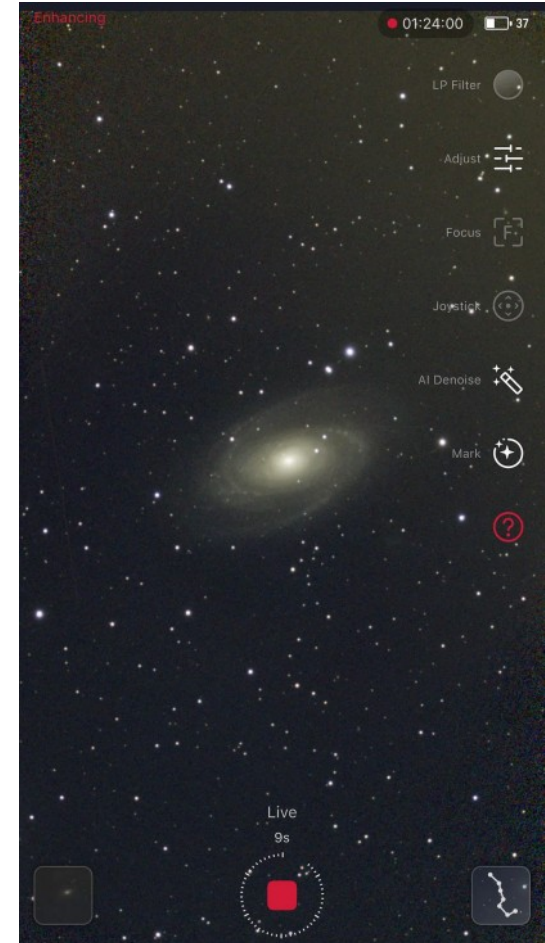
Aquisitions software (ASlair, NINA)

Central control hub for all systems



Smart telescopes

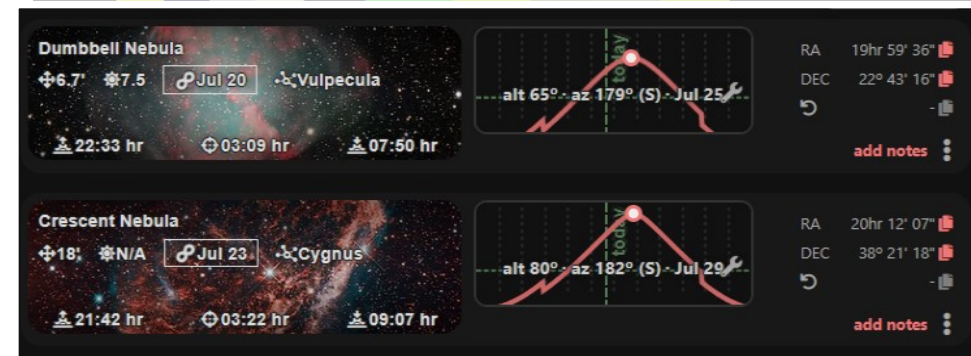
- Extremely good value for money
- Easy entry point into descent astrophotography, with imaging being fully controlled in phone ap
- To me not really interesting as I enjoy building and understanding my setup as much as using it....



Planning an astro-night

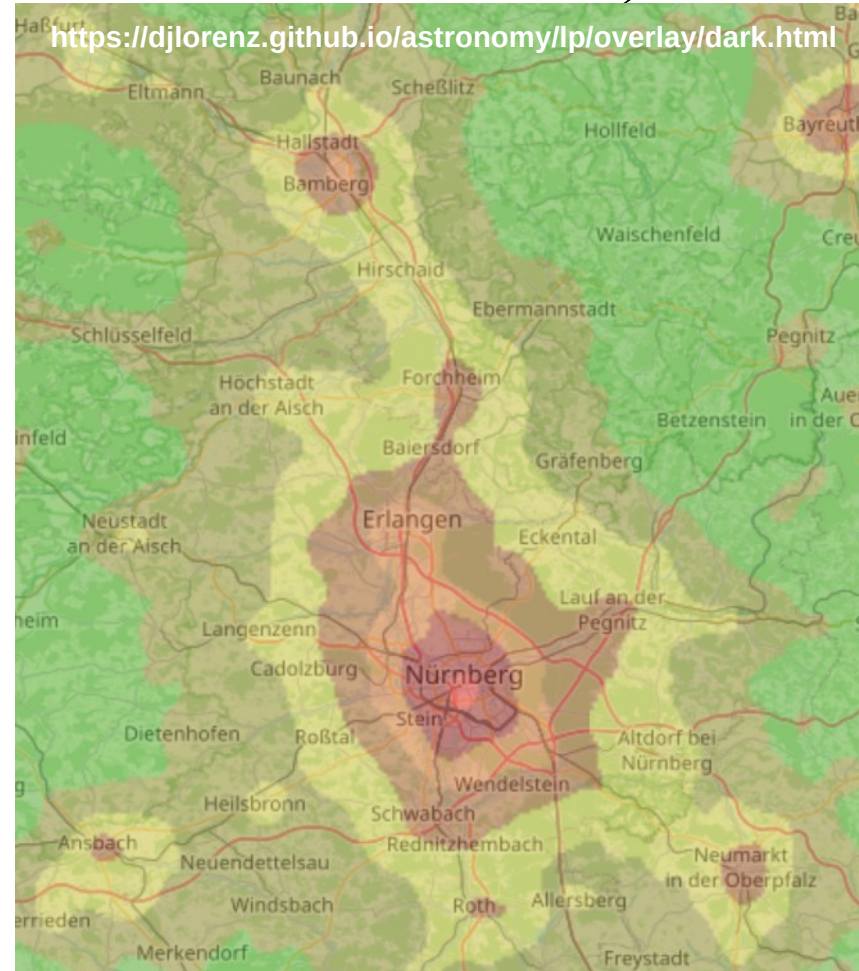
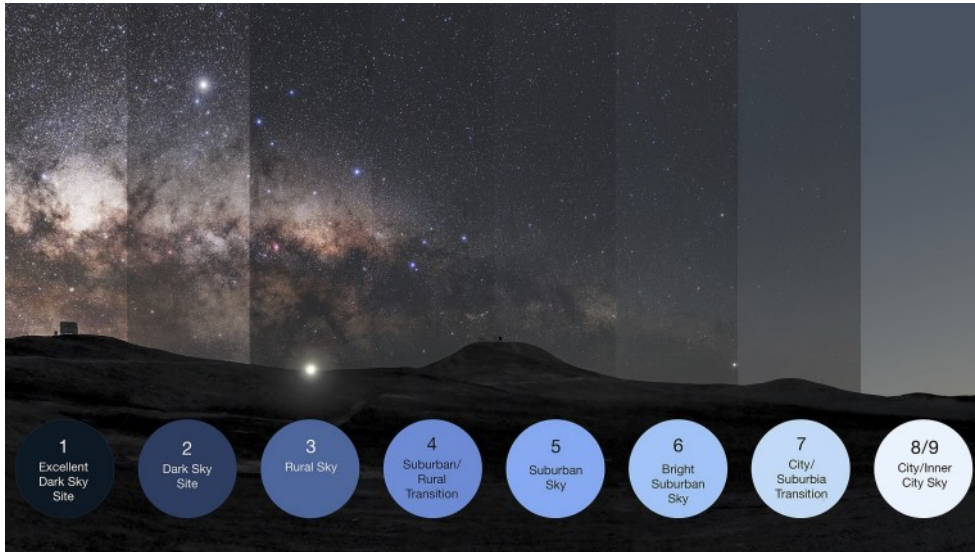
- Is the weather fine?
- What is the seeing like?
- What moon phase are we in?
- Which targets are high in the sky?
(close to opposition)
- Want to drive to a dark site?
→ target selection
(narrow vs broadband target)
- Pick and prepare
equipment to match target

	Clouds			Arc Sec.	Index		Jet Stream	Bad Layers			Ground		
	Low	Mid	High		1	2		Bot (km)	Top (km)	K/100m	Temp	Rel. Hum.	Celestial Bodies
23	0	0	0	1.35	5	5	15 m/s	01.6	02.6	0.6 K	18 °C	47%	---M----
Sat 2025-06-21 sunrise: 05:11 sunset: 21:19 moonrise: 01:55 moonset: 16:52 moonphase: 21%													
0	0	0	0	1.32	5	5	15 m/s	01.6	02.6	0.6 K	17 °C	49%	---M----P
1	0	0	0	1.29	5	5	15 m/s	01.6	02.6	0.7 K	16 °C	52%	-----P
2	0	0	0	1.24	5	4	16 m/s	01.6	02.6	0.7 K	16 °C	55%	-----S-NP
3	0	0	29	1.19	5	4	16 m/s	01.6	02.1	0.8 K	15 °C	58%	L----S-NP
4	0	0	70	1.16	5	4	14 m/s	01.6	02.1	0.9 K	14 °C	61%	L-V--SUNP
5	0	0	100	1.14	5	4	13 m/s	04.4	05.1	0.5 K	14 °C	63%	L-V--SUNP
6	0	1	100	1.13	5	4	11 m/s	01.6	02.1	0.9 K	15 °C	60%	L-V--SUNP
7	0	4	100	1.13	5	4	10 m/s	01.6	02.1	0.8 K	16 °C	56%	L-V--SUNP
8	0	6	99	1.12	5	4	9 m/s	01.6	02.1	0.8 K	18 °C	50%	LMV--SUN-
9	0	4	67	1.11	5	5	9 m/s	01.6	02.1	0.8 K	19 °C	46%	LMV--SUN-
10	0	1	24	1.09	5	5	9 m/s	01.6	02.1	0.8 K	21 °C	42%	LMV--SUN-
11	0	6	2	1.07	5	5	9 m/s	01.6	02.1	0.8 K	23 °C	38%	LMVMJSUN-
12	0	27	24	1.06	5	5	9 m/s	01.6	02.1	0.8 K	24 °C	33%	LMVMJSUN-
13	0	56	67	1.06	5	5	10 m/s	01.6	02.1	0.8 K	25 °C	31%	LMVMJ--UN-
14	0	77	98	1.05	5	5	10 m/s	01.6	02.1	0.8 K	26 °C	29%	LMVMJ--U--
15	0	83	100	1.04	5	5	10 m/s	01.6	02.1	0.7 K	26 °C	26%	LMVMJ--U--
16	0	81	99	1.02	5	5	10 m/s	01.6	02.1	0.7 K	26 °C	25%	LMVMJ--U--
17	0	71	86	1.01	5	5	10 m/s	01.6	02.1	0.7 K	26 °C	24%	-M-MJ--U--



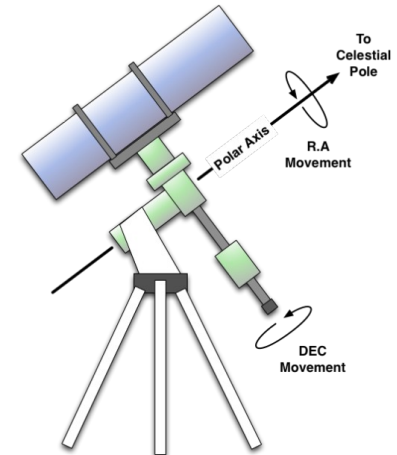
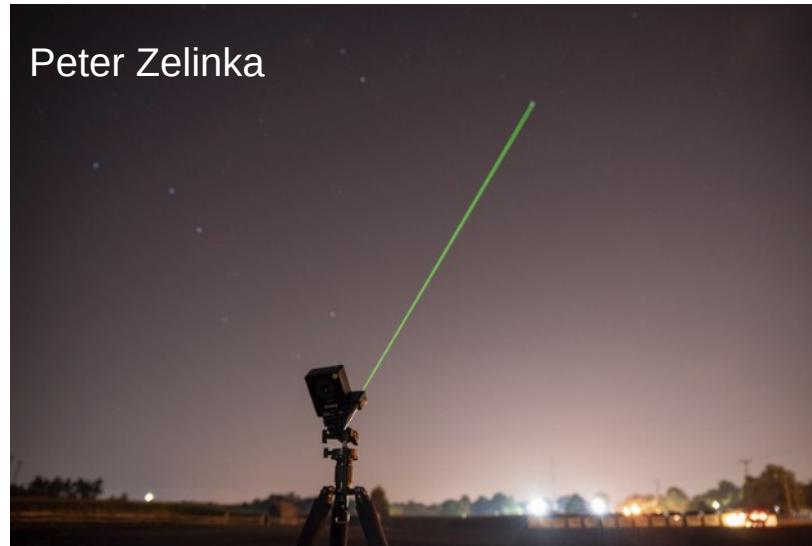
Light pollution (anthropogenic and the moon)

- Diffuse light background washes out targets
(SNR more important than #photons)
- Full moon is \sim bortel 7



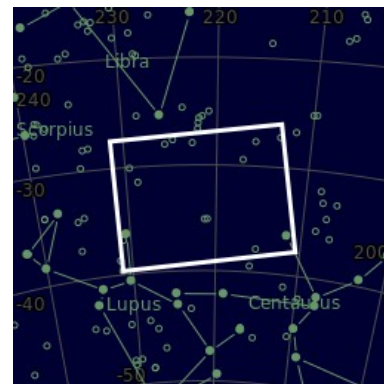
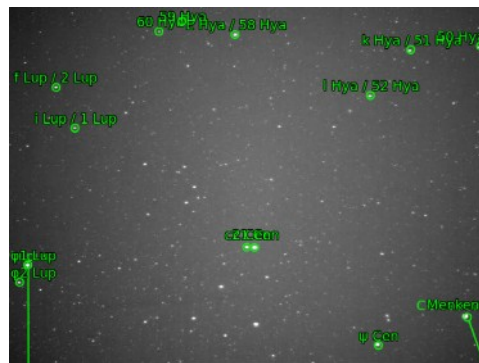
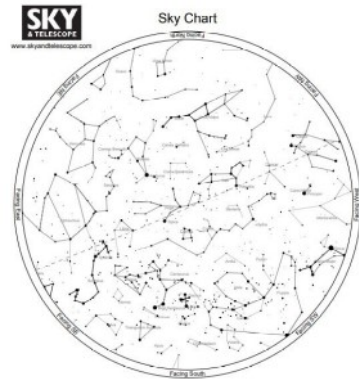
Polar alignment

- Align R.A. axis to celestial pole via small telescope inside R.A. axis
- Polaris is close but not exact ($\sim 0.45^\circ$ away)
- Star trackers sometimes feature lasers for rough alignment



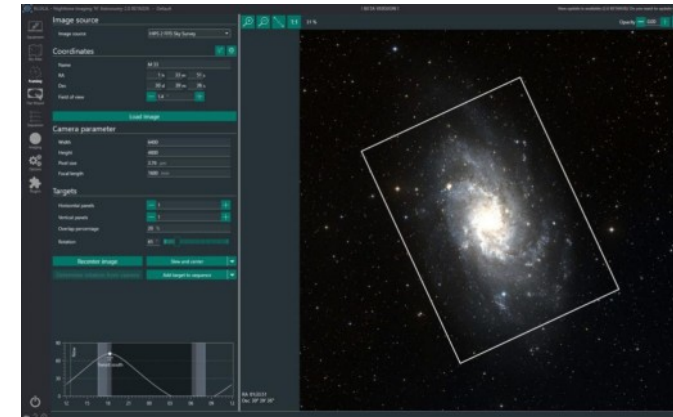
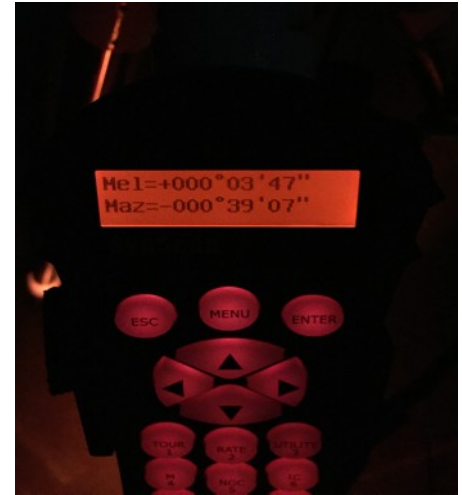
Finding the target I: Star hoping & plate solving (GPS in space)

- With a star tracker now you need to find the object by manually pointing the camera (while not knocking the tracker out of alignment)
- Star charts and phone planetarium software is helpful to go from constellation to constellation, star by star to the object
- This can be VERY frustrating and time consuming ...
- Most of the time you will not even see your DSO in the image, here plate solving helps to iterate

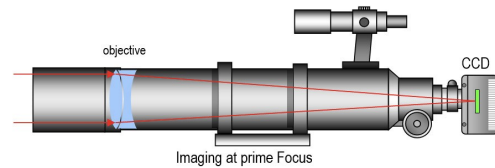


Finding the target II: Star alignment and automation

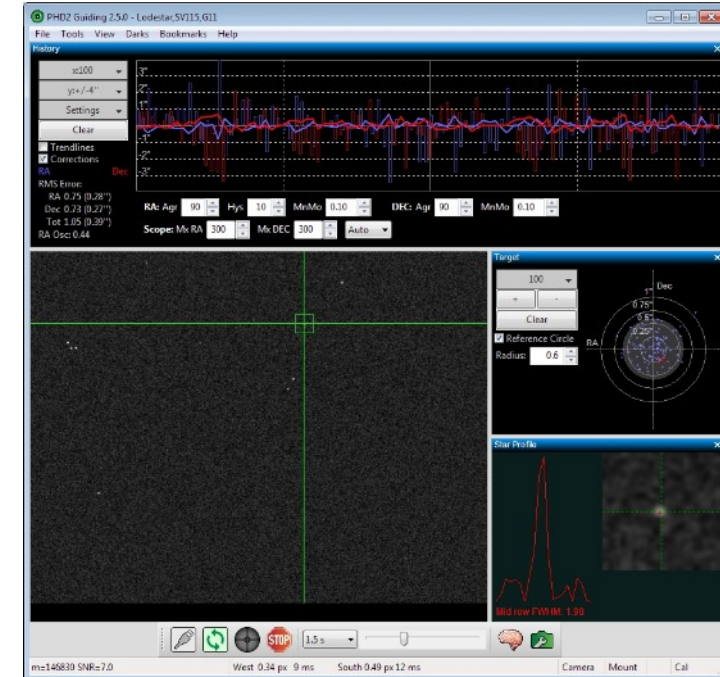
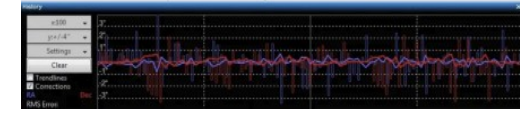
- Motorized go-to mounts feature star alignment methods
 - Mount slews to bright star
 - You manually center it in the FOV
 - Repeat 3x → mount knows orientation
- When software controlling all components (i.e. in NINA) software will plate solve every image and can automatically determine telescope orientation and slew to targets



Guiding



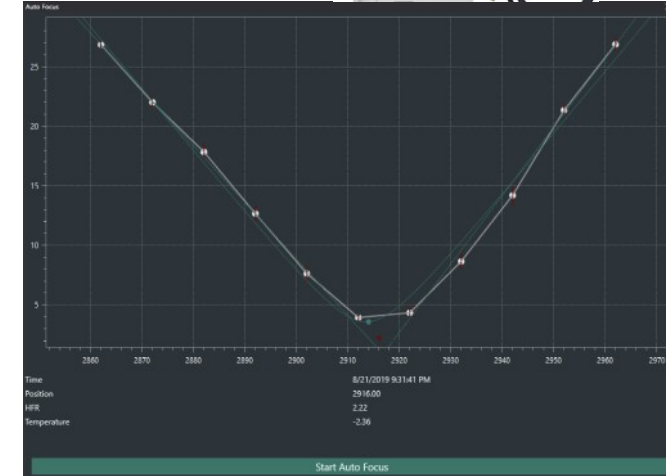
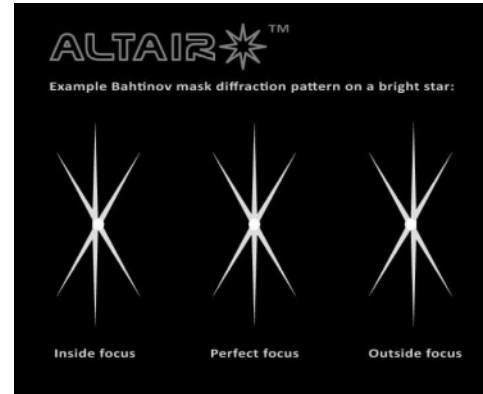
- Secondary guide scope takes an image every ~ second
- Software analyses star positions and sends correction signals to mount
 - can keep image stable to sub-arcseconds for minutes
- Also allows for dithering:
Sub-pixel resolution photography
by randomly offsetting image by fractional pixel counts



Focusing



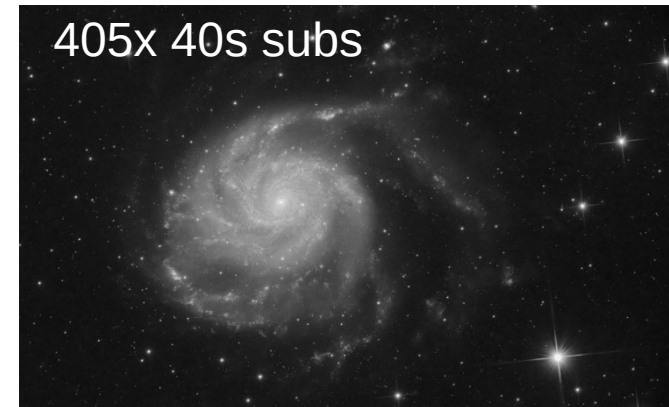
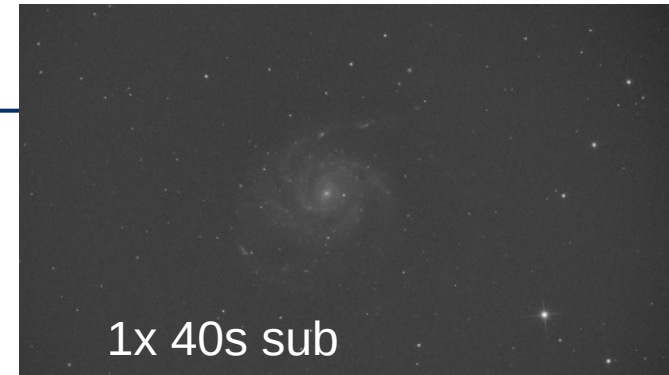
Spherical aberration call spherochromatism
typical for camera lenses



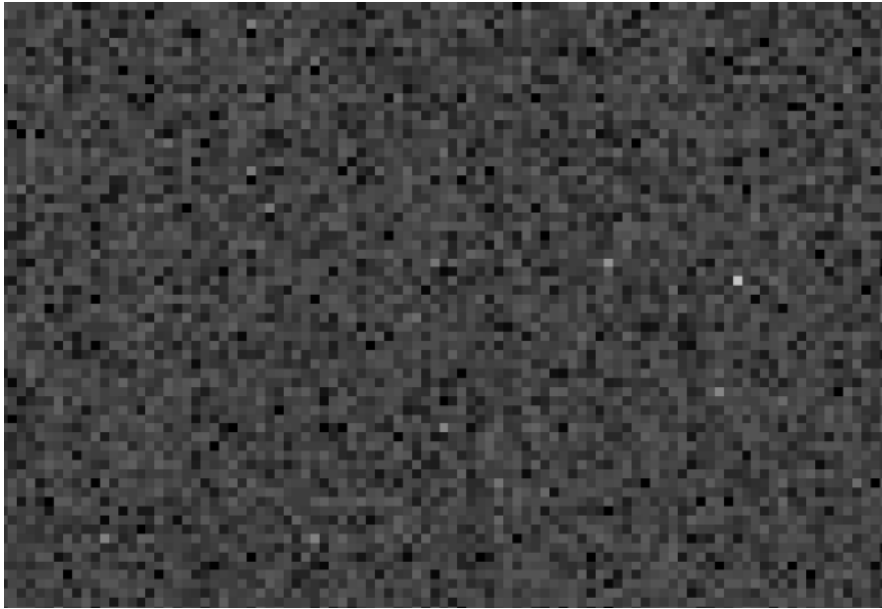
- Manual focusing aided by diffraction spikes of e.g. Bahtinov masks
- With a computer controlled setup can measure star size (& with an electronic auto-focuser set focus position with um-accuracy)

Gathering subframes (subs)

- Sub exposure time limited by:
 - Tracking / guiding stability
 - Camera dynamic range
(keep stars unsaturated for true color)
 - Satellite trails
 - Can range between seconds to several minutes
- Small project ~1h, serious projects starting at 10h, deep exposures can go to several hundred hours

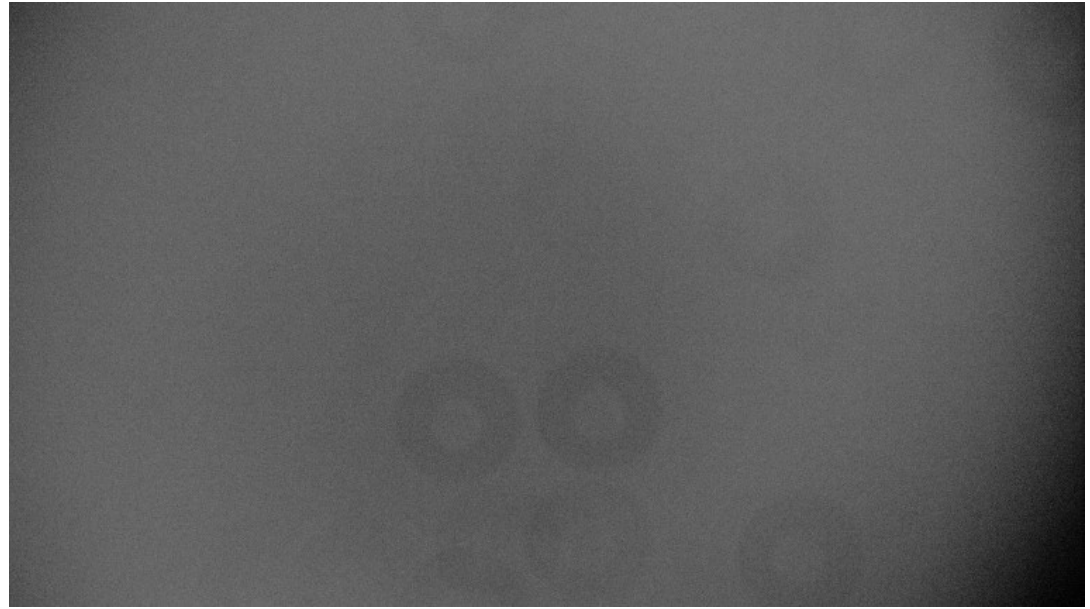


Calibration frames – darks & flats



Camera image in perfect darkness not perfectly dark:

- ADC read-noise
- Thermal dark current
- Taken at same temperature & exposure settings



Not the same sensitivity everywhere in the image:

- Sensitivity falls off at large angles
- Dust particles (motes)
- Taken using LED panels or the sky at dusk/dawn with diffuser cloth sheets

Post-processing software



Siril

Open-source, very capable for both artistic and scientific imaging

my personal favorite

PixInsight

~400€

Supposed to be THE best in software.

Not tried myself yet...



AstroPixelProcessor

~200€, supposed to be quite good



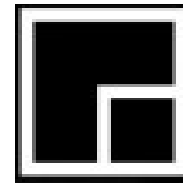
Sequator

Free, good for landscapes



SetiAstroSuite

Free, rapidly developing, strong AI integration



ASTAP

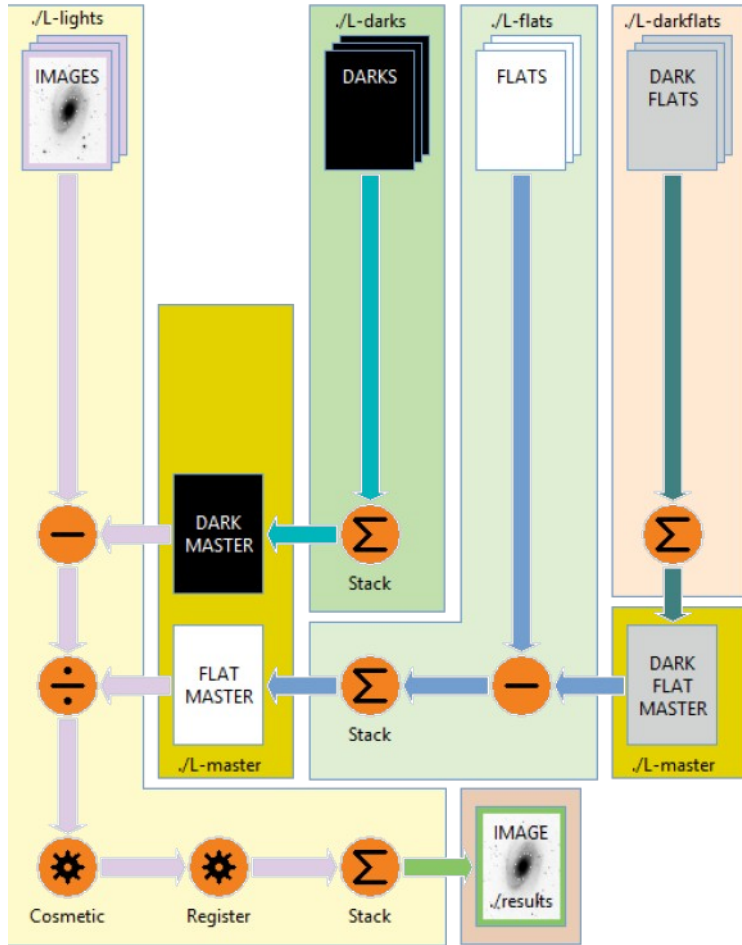
Free, great but technical, amazing plate solving



DeepSkyStacker

Free, a bit outdated

Stacking – calibrate, align & average subs

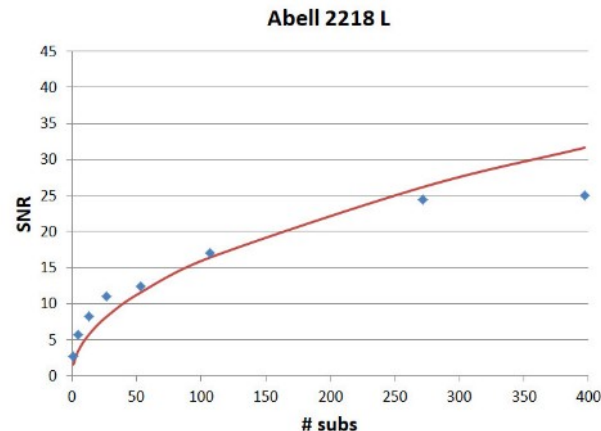


A tangent on SNR....

- Stacking is not about gathering more light, but about increasing the statistical uncertainty of the gathered light....
- Each pixel samples from a Poisson distribution of Target Flux + Sky flux
- We want to measure the Target Flux:

$$\text{SNR} = \text{Target Flux} * \text{time} / \text{sqrt}((\text{Target Flux} + \text{Sky Flux}) * \text{time})$$

-
- SNR grows as the square root of time
 - Time needed to reach a constant SNR growth linear with Sky flux



Cropping & Background extraction



- After stacking crop image to common region
- Fit and subtract low-order polynomial to remove sky flux / light pollution background
- (Or use AI models like GraXpert)

Sharpening & denoising (classic & AI)



- While there are limits to the obtainable resolution / sharpness images can be improved either by deconvolution (when knowing the PSF) or applying AI denoising/sharpening tools
- Best to go easy, as aggressive sharpening leads to artifacts



(Spectrophotometric) color calibration



Color in astrophotography is always tricky and often a matter of taste.

But initial color correction / white balancing often performed on catalog data of star colors (e.g. from the Gaia satellite mission)



Stretching



- Original image is „linear“: Twice the photon count = twice the image brightness
- We need to apply a transfer function to lift information from shadows into mid-tones and highlights
- Can be algorithmic (statistical stretches), guided (arcsinh, GHS) or manual
- Also referred to as histogram or curve transformations
- Be slow and go in steps, easy to over-do

Star processing (e.g. with StarNet++)



Foreground stars often a lot brighter compared to deep sky object.

Problem: Would want to stretch DSO strongly, but this clips stars....

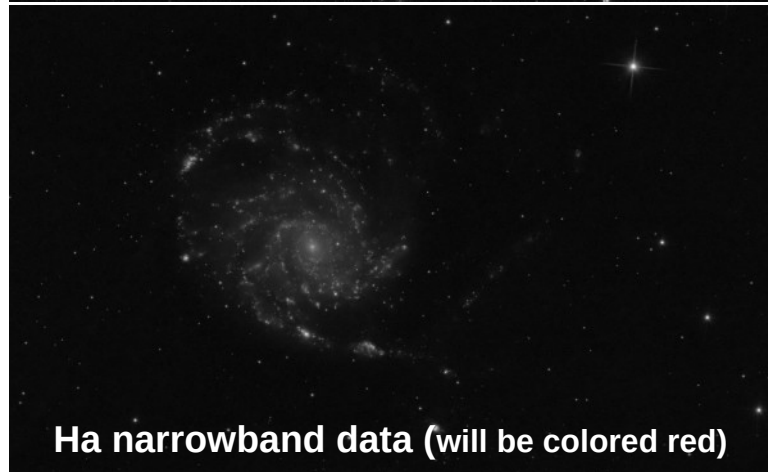
Common solution: Use (AI) tools to separate stars and DSO and process separately

Layer based image editors

- After stretching images mostly further processed in classical layer based editors (Photoshop, Gimp)
- Narrowband color assignment (color pallet picking)
- L-RGB merging (based on L mask)
- Stars and starless blending
- Artistic touchups:
 - Color contrasts
 - Wavelet sharpening
 - Curve transformations
 -



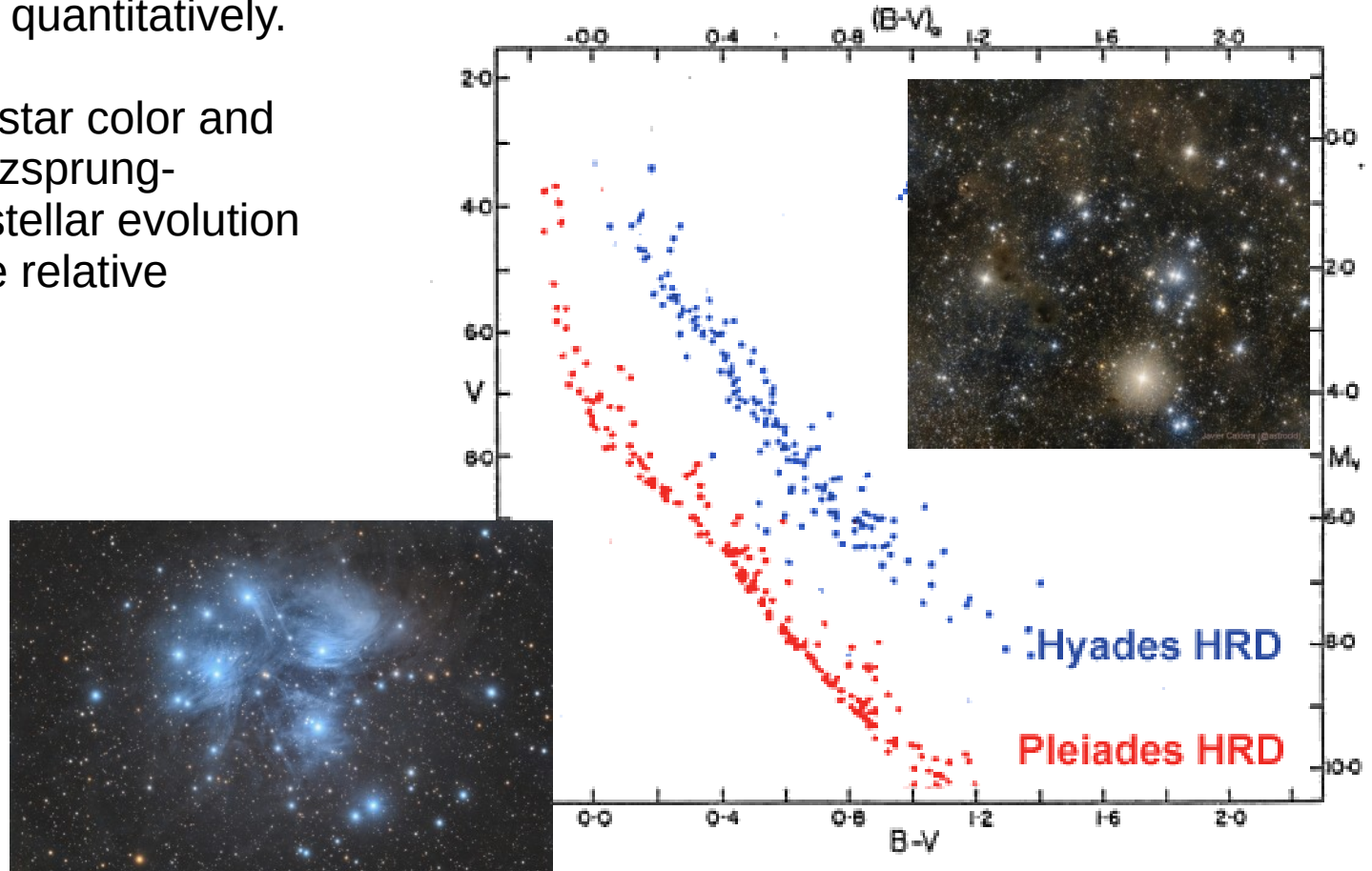
Merging narrowband and and/or L-RGB



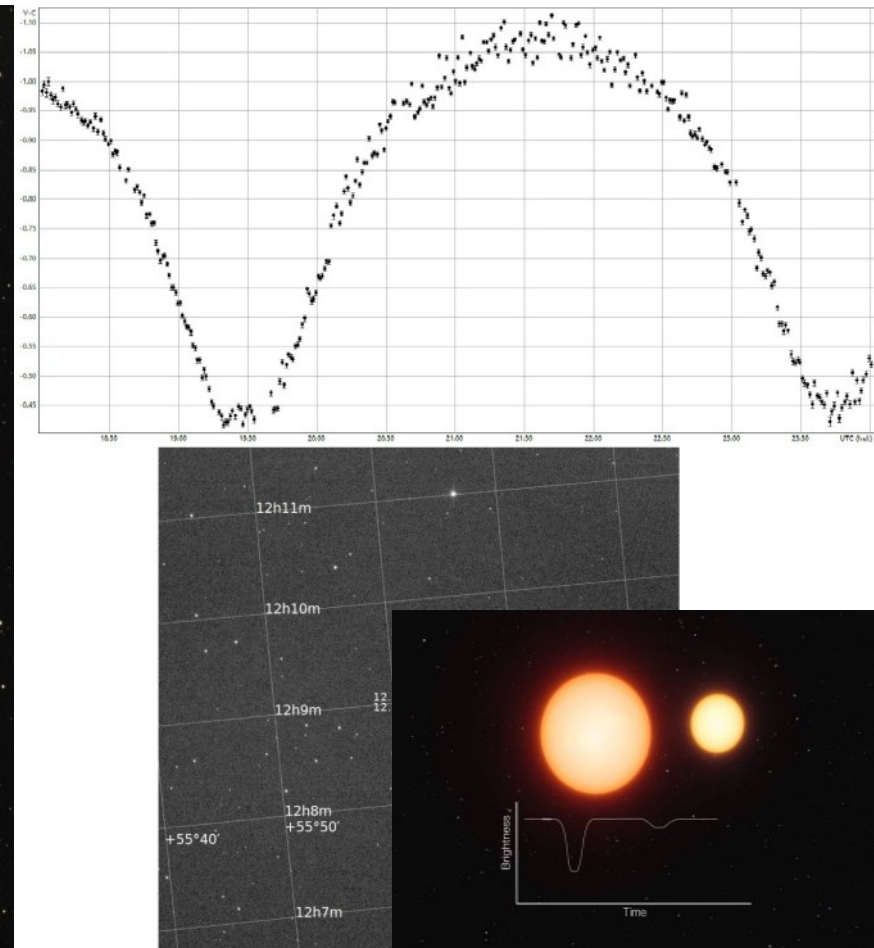
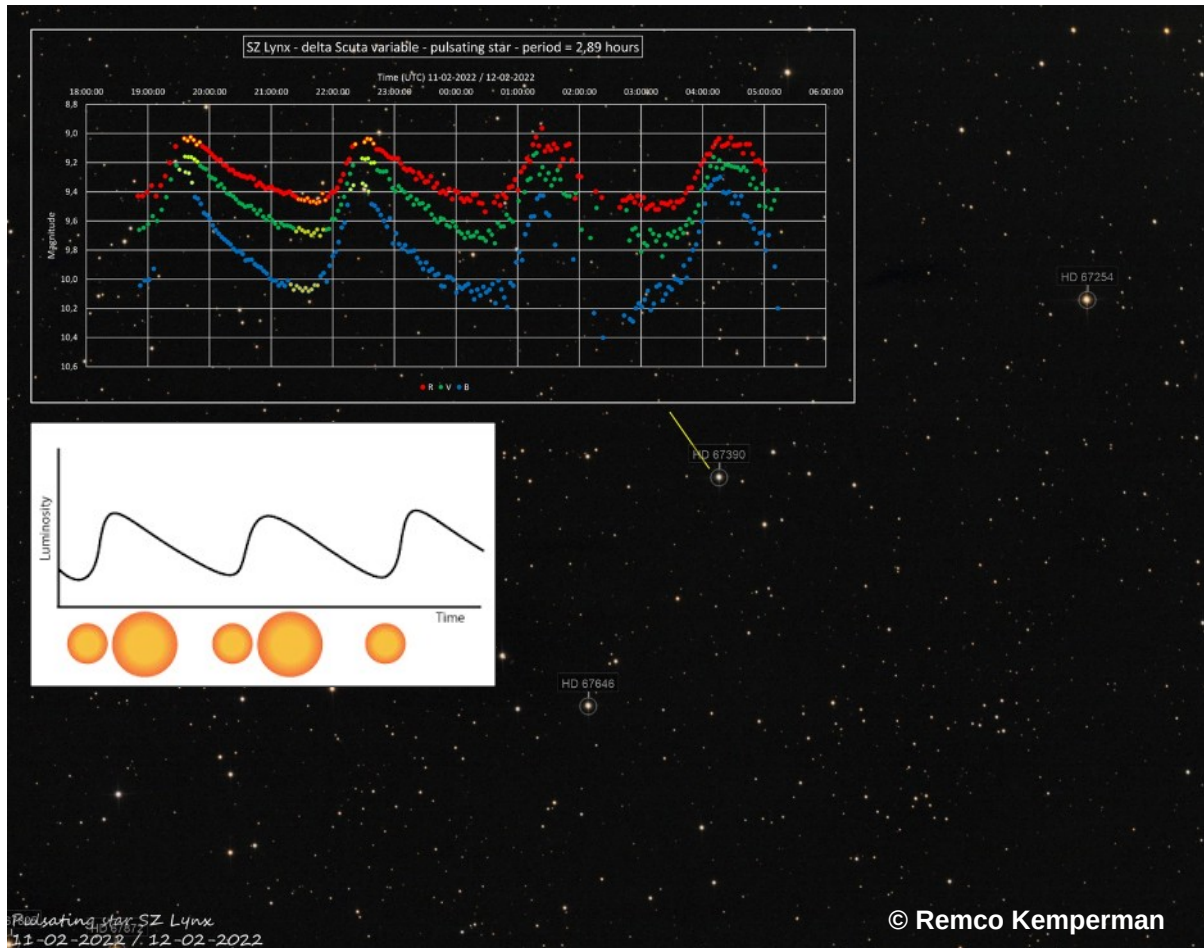
Scientific imaging: Photometry (HRD)

Images can also be used quantitatively.

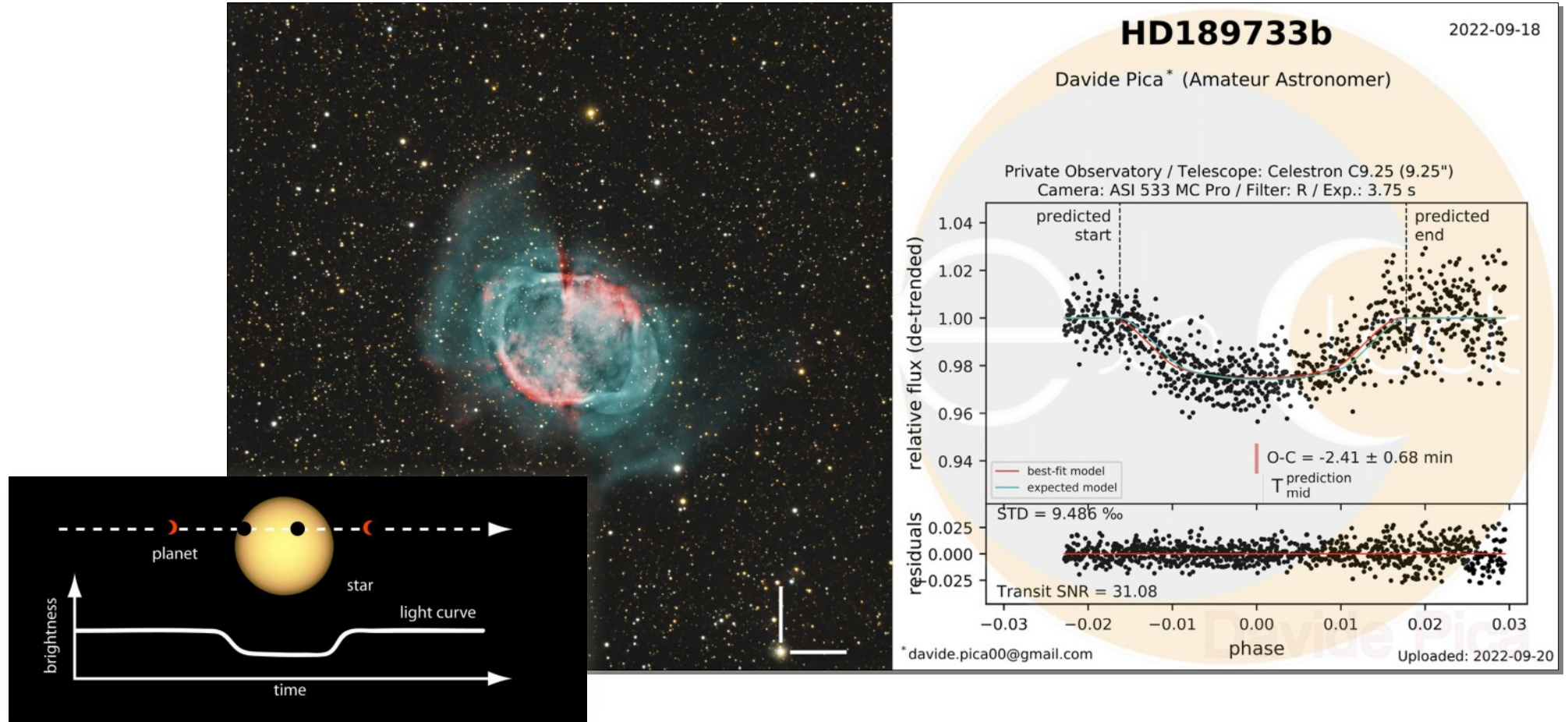
For example to measure star color and brightness, building „Hertzsprung-Russell diagram“ to see stellar evolution and for example measure relative distance of star clusters.



Scientific imaging: Variable stars



Scientific imaging: Exoplanet transients

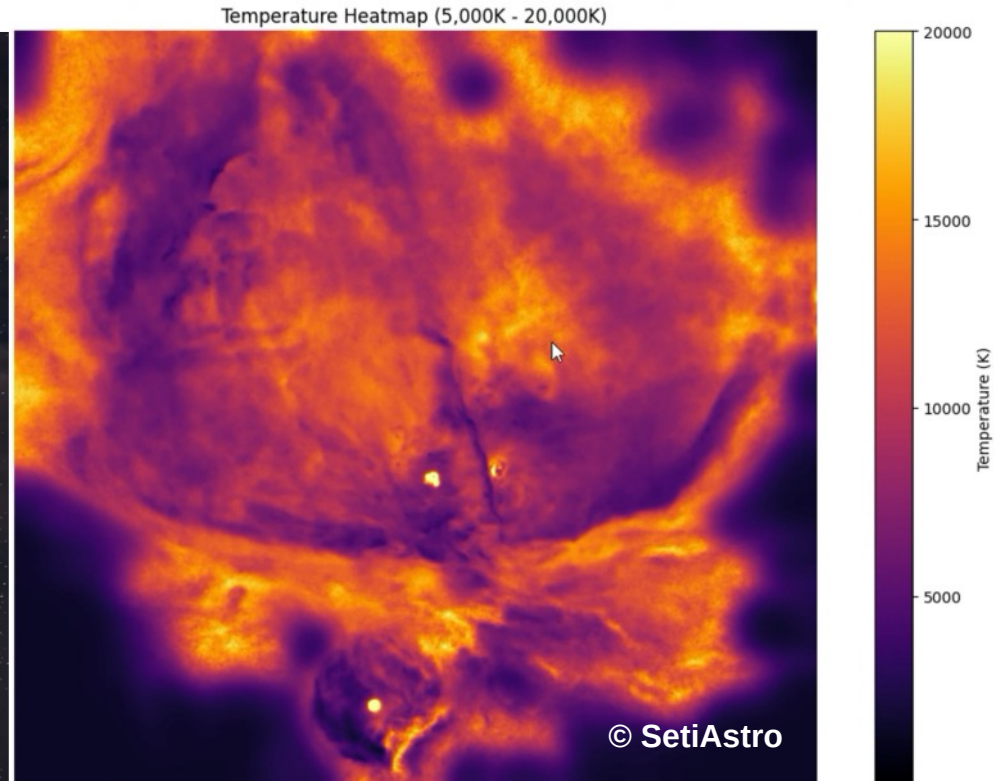


Scientific imaging: Balmer decrement imaging

Capturing the ratio of H-alpha (656nm) to H-beta (486nm) emission to deduce hydrogen gas temperature



© Itto Ogami



Temperature (K)

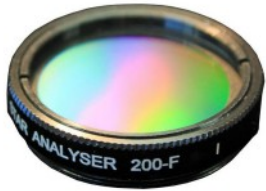
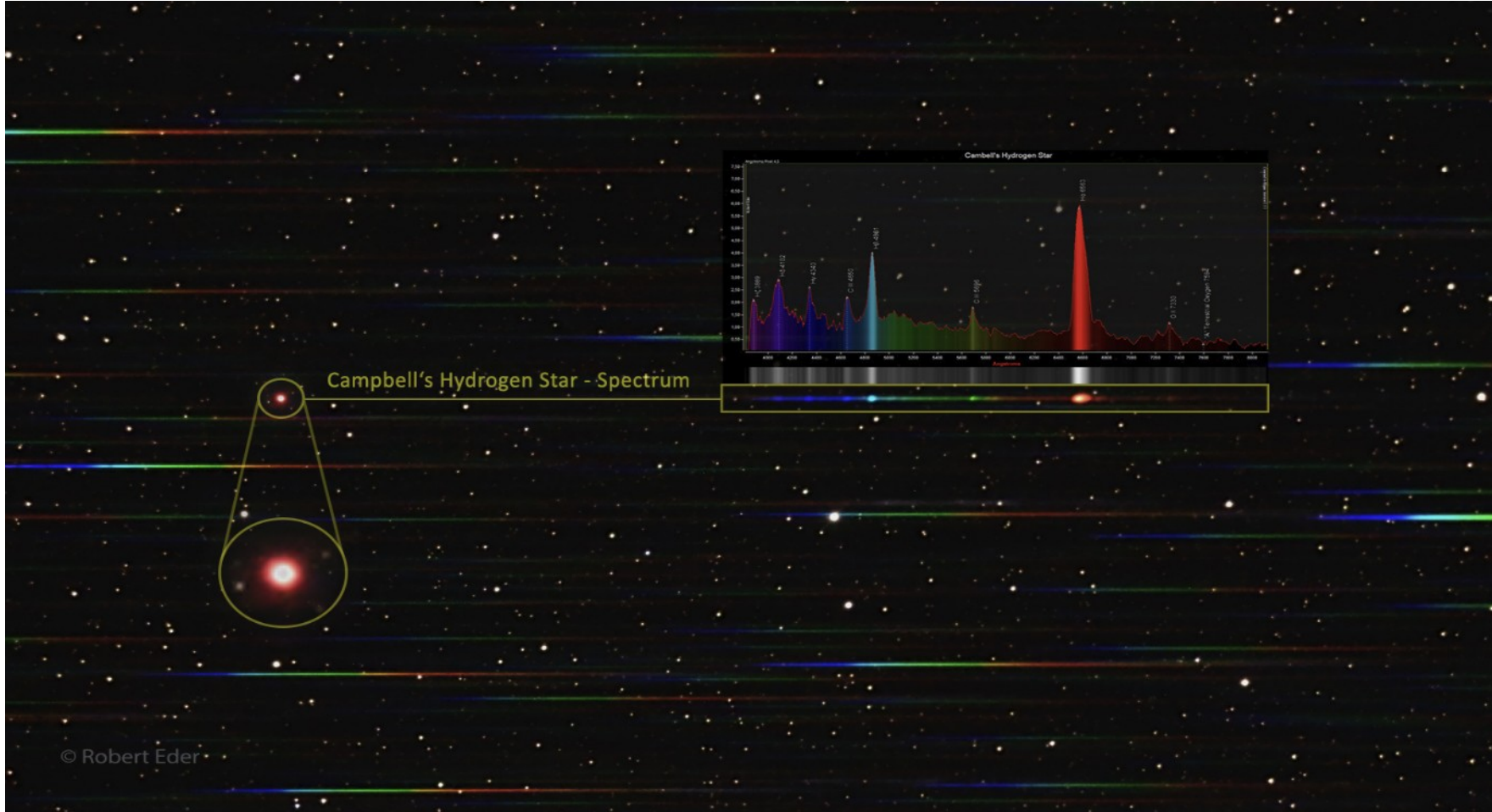
20000

15000

10000

5000

Scientific imaging: Spectroscopy



Astrophotography as an academic hobby

- We **can** get AMAZING results
- BUT there will **also** be better pictures than yours
- The hobby is mostly about learning to improve yourself and plenty of failing along the way
- So temper your expectations and enjoy the progression



Some more pictures





**Thank you for your
attention!**

Questions?